

# Star-formation histories early type dwarfs



Mina Koleva (UGent, Belgium)  
&

Antoine Bouchard, Sven De Rijcke, Philippe Prugniel, Isabelle  
Vauglin, Werner Zeilinger

# Questions

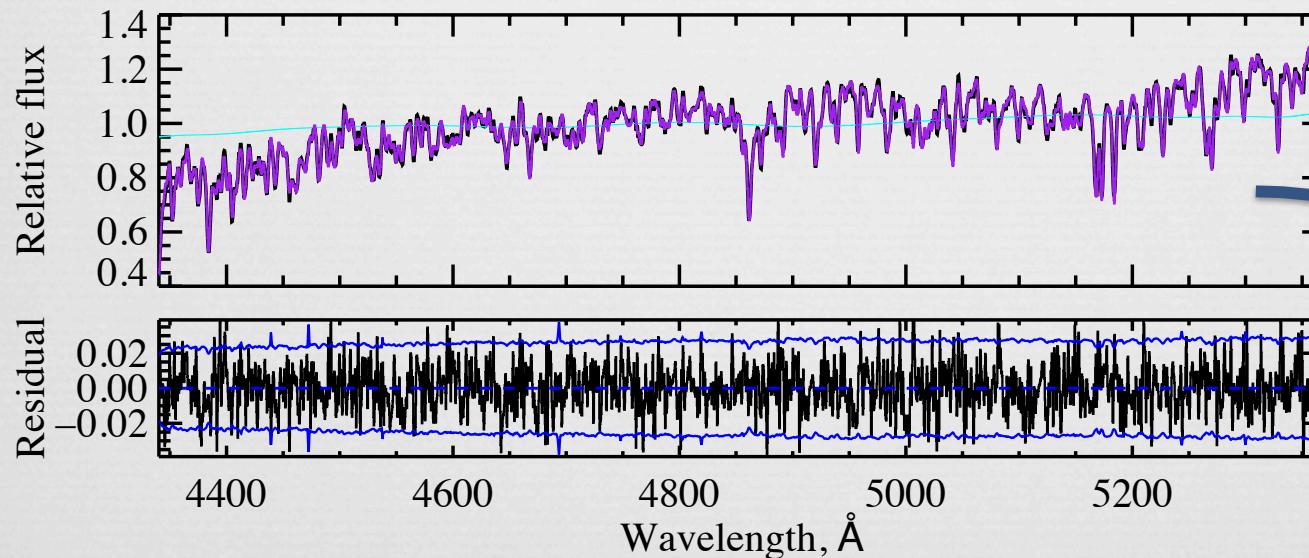
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- ❖ How star formation is regulated in dwarf galaxies?
  - ❖ Cooling, collapse, feedback
  - ❖ DM distribution, mass, angular momentum
- ❖ Why do galaxies stop forming stars?
  - ❖ Internal/environmental processes
- ❖ Which are the progenitors of the quiescent dwarfs?

# Tools



- ❖ High precision, spectroscopic analyses in the optical domain
- ❖ Full spectrum fitting (ULySS, [ulyss.univ-lyon1.fr](http://ulyss.univ-lyon1.fr))



Star-formation  
and metal  
enrichment  
histories

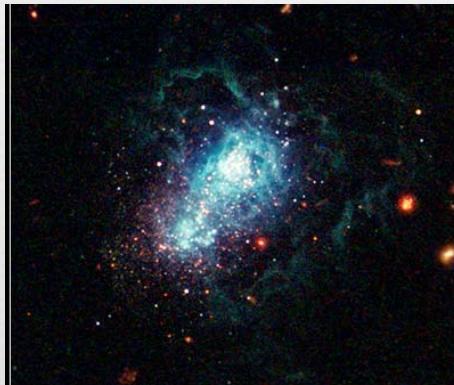
# Tools

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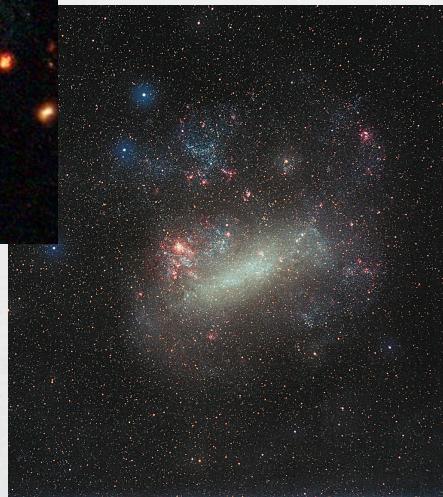


- ❖ High precision, spectroscopic analyses in the optical domain
- ❖ Full spectrum fitting
  - ❖ ULySS ([ulyss.univ-lyon1.fr](http://ulyss.univ-lyon1.fr))
- ❖ Targets
  - ❖ Dwarfs with little star formation ( $M_\star \sim 10^9 M_\odot$  in this talk)

# Metamorphosis



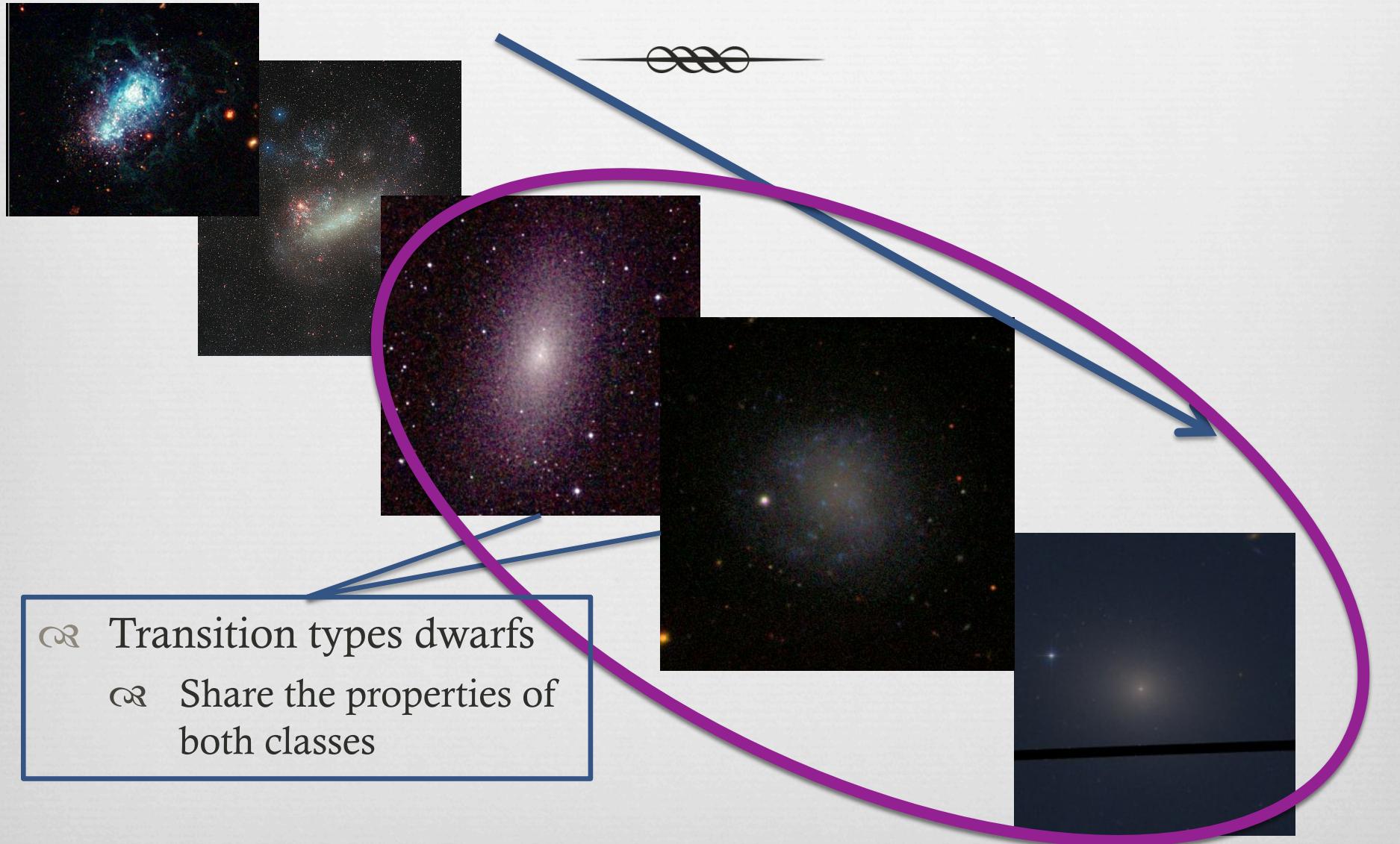
- ❖ Star-forming dwarfs
  - ❖ Star formation
  - ❖ Gas/dust
  - ❖ Irregular shapes



- ❖ Quiescent dwarfs
  - ❖ No star formation
  - ❖ No gas/dust
  - ❖ Regular isophotes



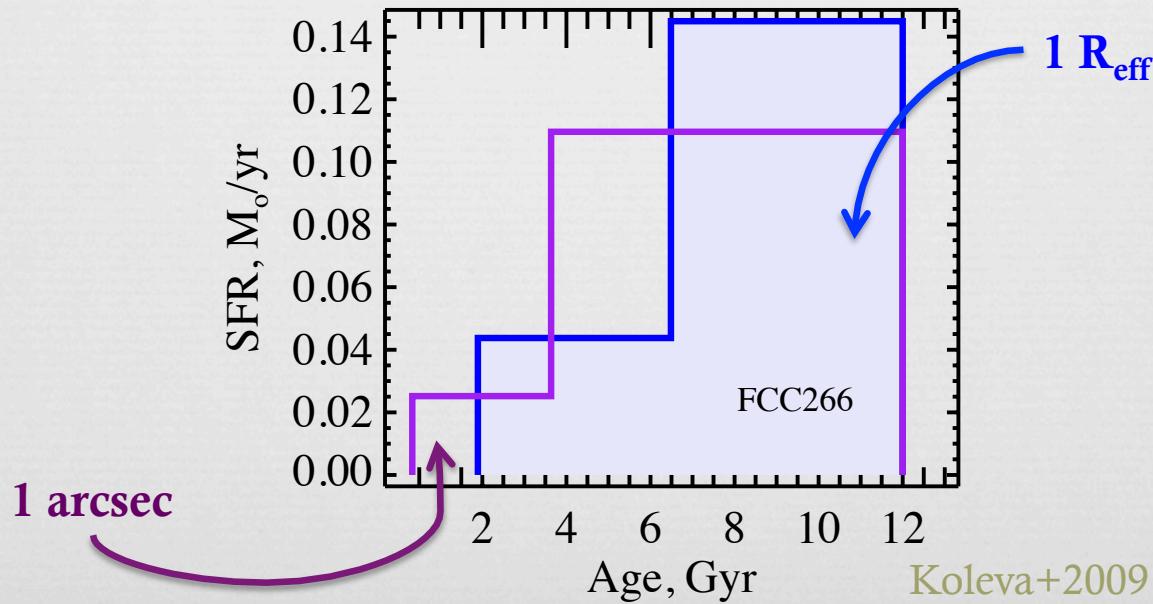
# Metamorphosis



# Dwarf elliptical galaxies



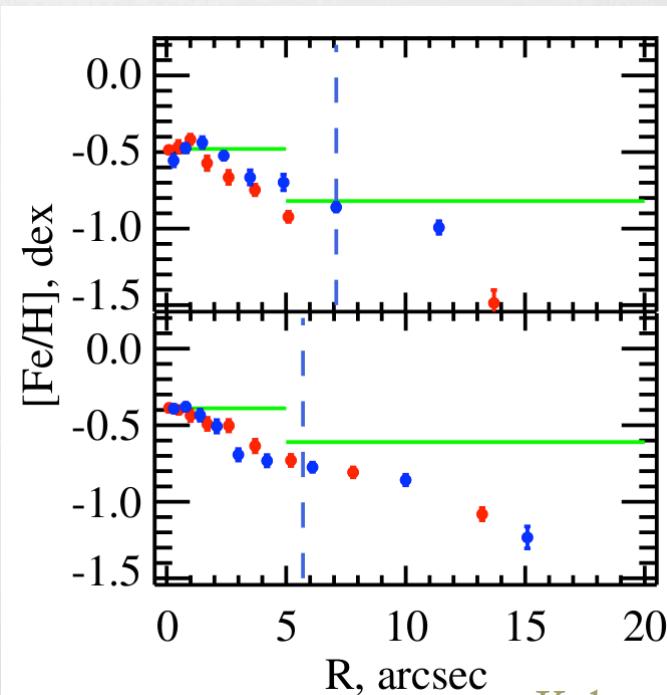
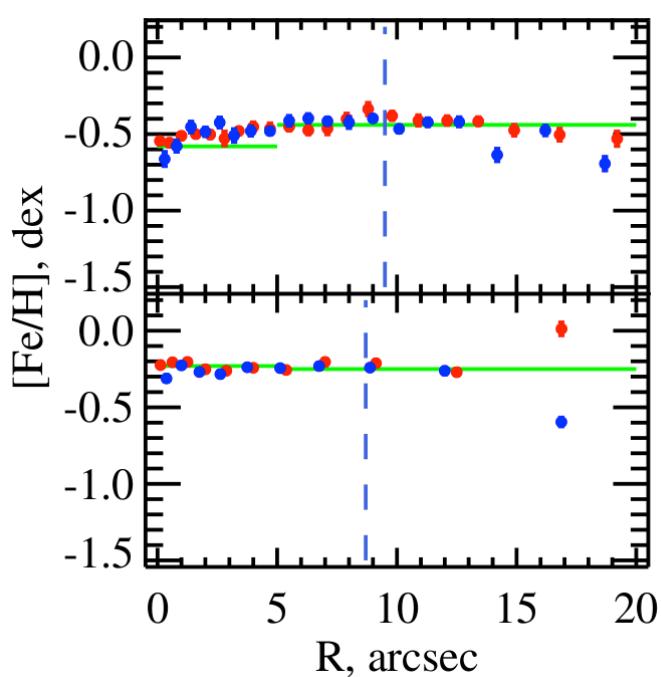
- ❖ Sample (similar to NGC205)
- ❖ Different environment
- ❖  $M_\star \sim 10^9 M_\odot$
- ❖ Long-slit optical spectra
- ❖ Results – SFHs
- ❖ 80% stars  $> 5$  Gyr
- ❖ Zoo of SFHs
- ❖ SF area shrinks with time



# Dwarf elliptical galaxies



- ❖ Results – radial gradients
  - ❖ Flat and positive age gradients
  - ❖ Flat and negative metallicity gradients



Koleva+2009

# Dwarf elliptical galaxies

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## ❖ Discussions

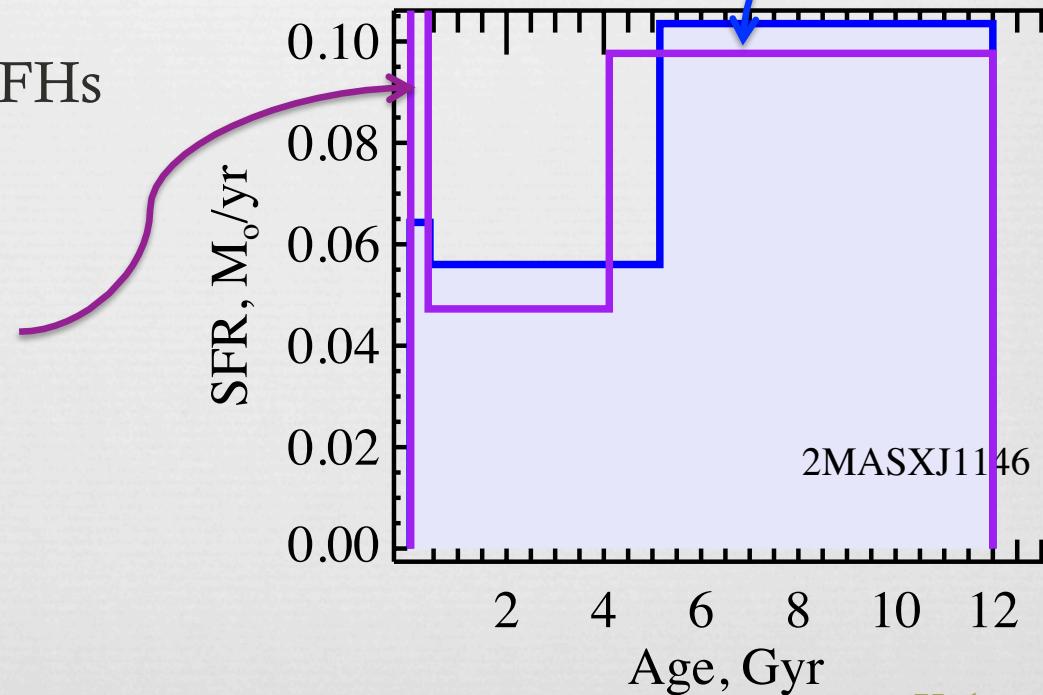
- ❖ Prolonged, heterogeneous star-formation histories
  - ➔ galaxies will keep their gas (and SF) until it is removed by environment (simulations, e.g. Valcke+ 2008)
- ❖ Flat metallicity gradients
  - ➔ **rotating** galaxies (dIrr) with continuous SF (see the talk of J. Schroyen)
- ❖ Negative metallicity gradients
  - ➔ galaxies with episodic SF (cf. BCDs), shrinking with time

# Transition type galaxies



- ❖ Sample – SDSS, long-slit
- ❖ Results
  - ❖ Prolonged SFHs

Vazdekis  
SP models

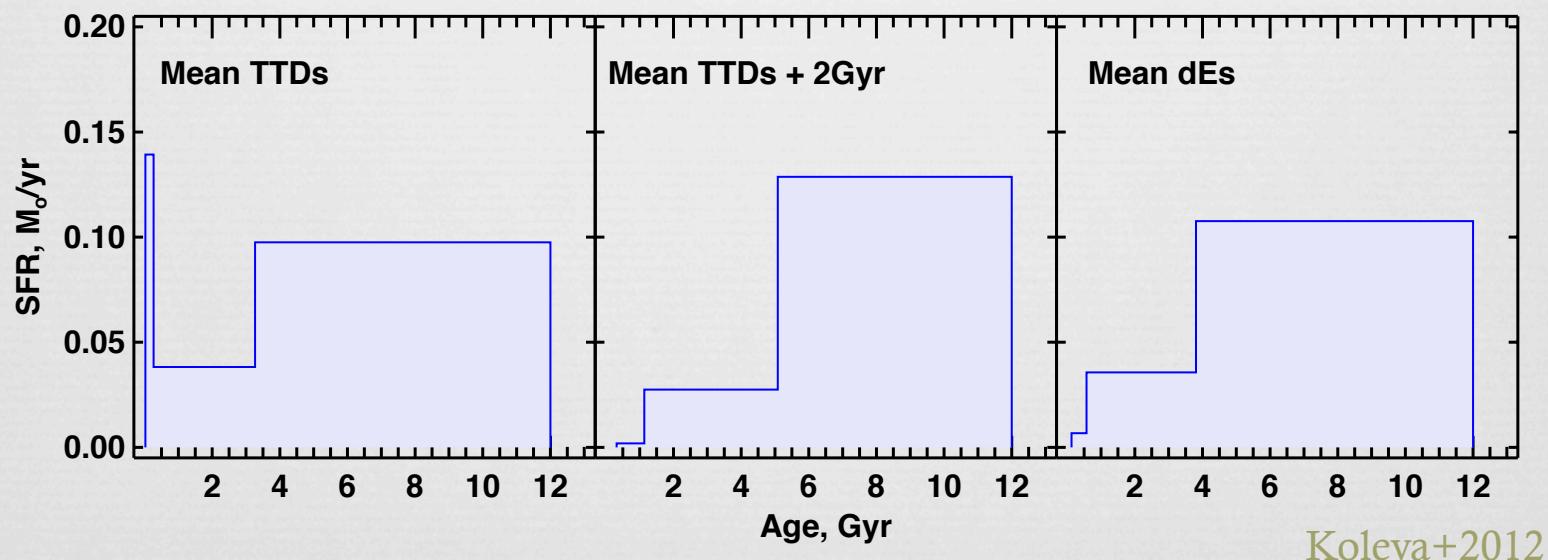


Koleva+2012

# Results: TTD vs. dE



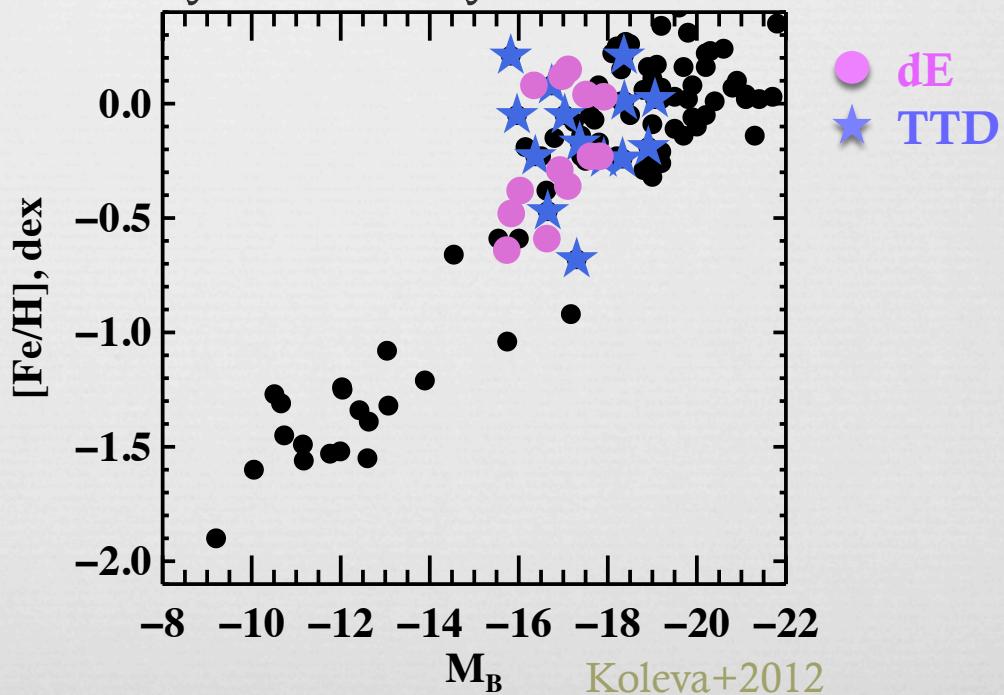
- ❖ Passive evolution of a TTD leads to a galaxy equivalent to a dE
- ❖ Identical star-formation histories



# Results: TTD vs. dE



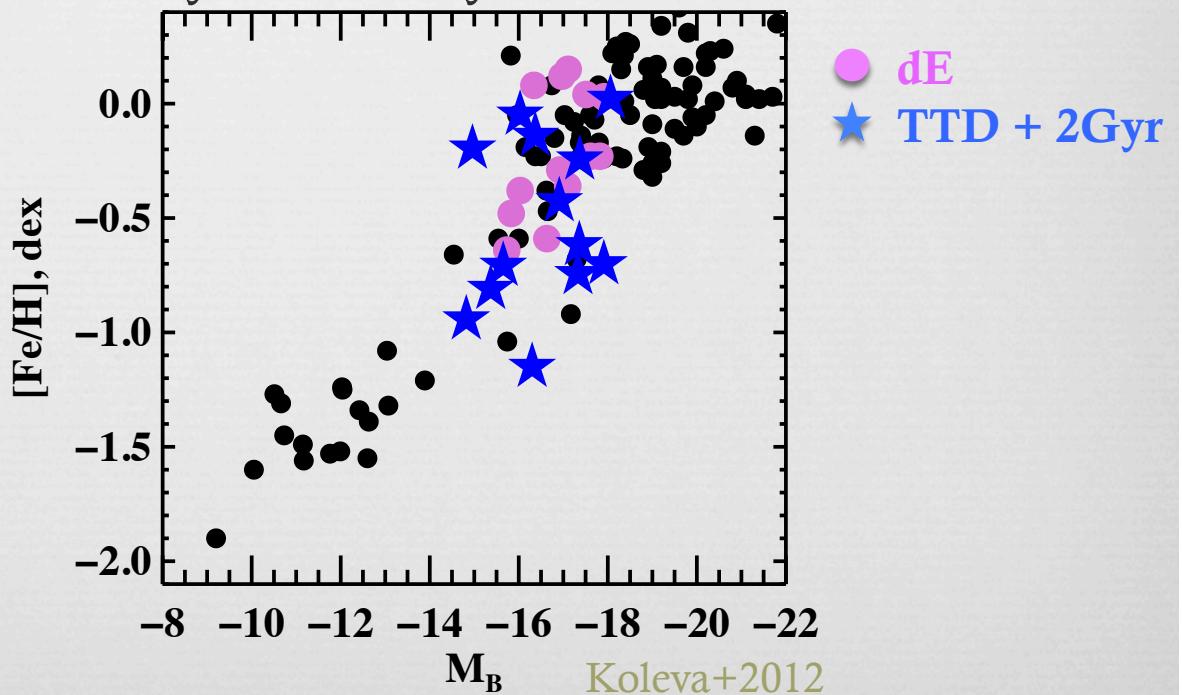
- ❖ Passive evolution of a TTD leads to a galaxy equivalent to a dE
- ❖ On the same metallicity-luminosity relation



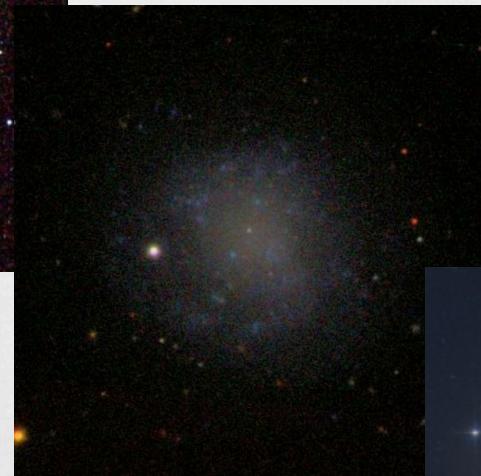
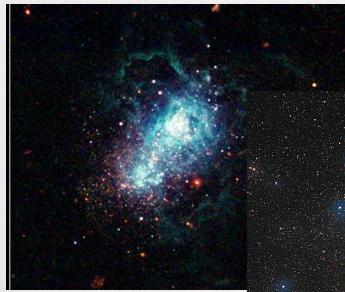
# Results: TTD vs. dE



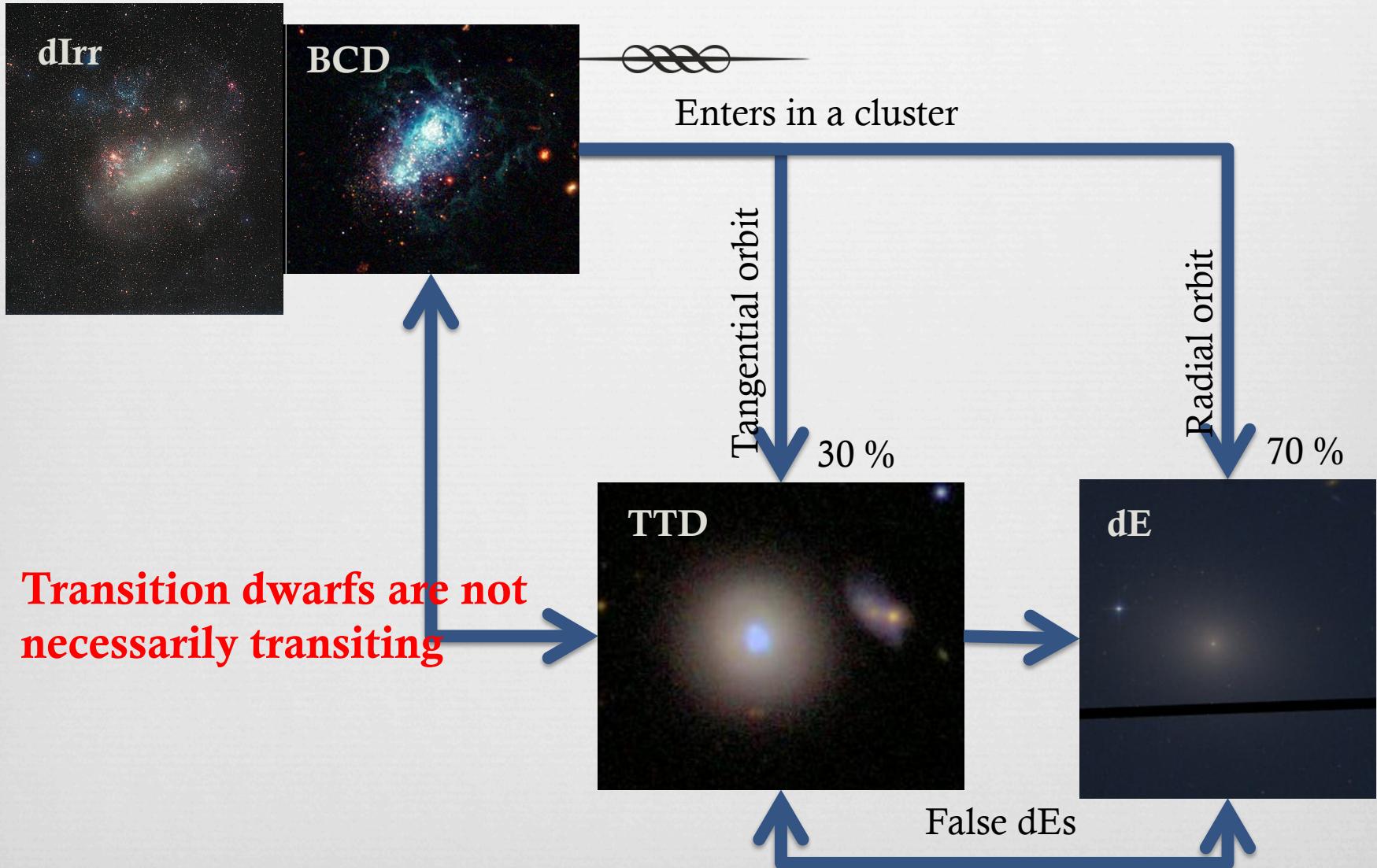
- ❖ Passive evolution of a TTD leads to a galaxy equivalent to a dE
- ❖ On the same metallicity-luminosity relation



# Possible interpretation



# Possible interpretation



# Conclusions

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- ❖ *TTDs and dEs are similar*
  - ❖ *Stellar populations alike between TTDs and dEs (Koleva + 2012)*
  - ❖ Also for less massive TTDs (Grebel + 2003, Weisz + 2011)
- ❖ TTDs should be found only in the outskirts
  - ❖ Yes (Conselice + 2003, Bouchard + 2009, Drinkwater + 2001)
- ❖ ~30% dwarf galaxies have gas. Hence, most of the galaxies should be on radial orbits
  - ❖ Yes (Fornax, De Rijcke et al 2010),
- ❖ TTDs and BCDs are similar in sparsely populated environments
  - ❖ Similar metallicities between BCDs and TTDs (Sanchez-Almeida + 2009)
  - ❖ Episodic SFH, steep metallicity gradients..
  - ❖ **Need more data to be tested!**

# Results: TTD vs. dE



- ❖ Passive evolution of a TTD leads to a galaxy equivalent to a dE
- ❖ Metallicity enrichment histories

