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## THE SUN AS A VARIABLE STAR II

Photometric Observations of Uranus, Neptune and Standard Stars in the Years 1953 - 1961

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### Abstract

The method used in the reduction of two-color photometric observations is described. Sixteen standard stars of nearly solar type were regularly observed for 7 years and no systematic changes in their brightness were found. The r.m.s. deviation of the yearly mean blue magnitude for none of these stars exceeds  $\pm 0^m.009$ . No systematic change in the blue magnitudes of Uranus and Neptune over a period of nine oppositions is obtained if the following conditions are fulfilled: 1) the gradient of the energy distribution in the spectrum of each of these planets within the blue filter spectral region is assumed to be the same as for a star with the same B-V color-index as the planet, and 2) with regard to Uranus only, the photometric effects due to its oblateness are half as great as those for the uniform distribution of brightness over the apparent disc of the planet.

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### Acknowledgements

### I. Introduction

An outline of the purpose of this work, the methods used and the preliminary results were given in the first part of this paper by Johnson and Iriarte (1) and in earlier papers by Giclas (2), by Hardie and Giclas (3) and by Mitchell (4).

The determination of the precise magnitudes of Uranus and Neptune in blue color is accomplished in three steps. First, atmospheric extinction and the coefficients of transformation from the instrumental to the BV photometric system are determined from the observations of primary standard stars in blue and yellow colors. (discussed in Section II). Secondly, the magnitudes and colors of the comparison stars situated along the paths of Uranus and Neptune are determined by comparison with the mean values for 16 bright stars called "Ten Year Standards" (Section III). Finally, the difference of blue magnitudes between each of the planets and two comparison stars situated at small angular distances from the planet is determined (Section IV). The resulting magnitudes of Uranus and Neptune can be used as indicators of the variability of the Sun.

The observations and their discussion were carried on at the Lowell Observatory, with the financial support of the U. S. Air Force.\*

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## II. Atmospheric Extinction and Transformations to The BV System

*A. Transformation Equations.* The following observing schedule is used for all two-color photometric observations in the present program. The star is measured first with a blue filter (5mm Corning 5030 filter + 2mm Schott GG 13) and next with a yellow filter (3.6mm Corning 3384). For each color two deflections of the Brown recorder which are separated by the deflection for sky background are always obtained. All the measurements of the star and sky background, both with blue and yellow filters are always taken with the same amplifier gain. The observation is concluded by measuring the deflection for a radioactive standard light source without a filter. This source is green in color and is measured at a fixed amplifier gain, such as that used for measuring 4th magnitude stars. The amplifier gain is selected so that the deflections with the blue filter and with the standard source are larger than 0.6 of full scale on the recorder. The deflections with a yellow filter for early-type stars were sometimes only 0.15 of full scale; this probably diminished the accuracy of measures with the yellow filter.

We denote by  $B'$  the mean of two deflections for a star observed with the blue filter minus the deflection for sky background with the same filter. By  $V'$  we denote the similar difference with a yellow filter and by  $S'$  the deflection for the standard source minus the deflection for dark current at the same gain. Furthermore, let  $S_b$  denote the magnitude difference between the star under consideration and an imaginary star which would give the same deflections at a certain fixed, arbitrarily chosen amplifier gain, the same for all the stars in the program.

We define now the quantities:

$$C_{y_0} = -2.5 \log B'/Y' , \quad (1)$$

$$m_{y_0} = S_b - 2.5 \log Y'/S' , \quad (2)$$

The color index  $B-V$  and the yellow and blue magnitudes of the BV photometric system are connected with  $C_{y_0}$  and  $m_{y_0}$  by the equations:

$$B-V = \frac{A_1 + A_2 C_{y_0} - K_1 M}{1 + K_2 M} , \quad (3)$$

$$V = m_{y_0} + A_5 + A_6 (B-V) - Q_{y_1} M - Q_{y_2} M (B-V) , \quad (4)$$

$$B = m_{y_0} + C_{y_0} + A_7 + A_8 (B-V) - Q_{b_1} M - Q_{b_2} M (B-V) , \quad (5)$$

where  $A_i$  ( $i=1,2,5,6,7,8$ ) are the coefficients of transformations to the BV system,  $K_1 + (B-V)K_2$  is the extinction coefficient for blue-yellow color-index while  $Q_{y_1} + (B-V)Q_{y_2}$  and  $Q_{b_1} + (B-V)Q_{b_2}$  are the extinction coefficients for the yellow and blue magnitudes.\* Equation (5) is obtained by adding the equations (3) and (4) and by making the following substitutions

$$A_7 = (A_1/A_2) + A_5 , \quad (6)$$

$$A_8 = 1 + A_6 - (1/A_2) , \quad (7)$$

$$Q_{b_1} = Q_{y_1} + (K_1/A_2) , \quad (8)$$

$$Q_{b_2} = Q_{y_2} + (K_2/A_2) . \quad (9)$$

The air mass at the time of observation, denoted by  $M$ , is computed by using the hour angle of a star as read on the telescope's setting circle at the middle of each observation. Tables and graphs facilitating computation of air mass accurate to 0.001 were prepared on the basis of Bemporad's tables of air mass reprinted by Schoenberg (6).

The coefficients of equations (3) and (4) are determined only from the observations of the primary standards of the UBV system. The following 8 stars were considered as such standards:  $\beta$  Cnc,  $\eta$  Hya, HR 4550, 90 Leo (A+B),  $\alpha$  Ser,  $\beta$  Lib,  $\epsilon$  CrB and  $\tau$  Her. The magnitudes and colors which were assumed for these stars are those given by Johnson and Harris (7); the values taken from this paper are henceforth denoted by  $(B-V)_s$  and  $V_s$ .

The coefficients  $K_2$  and  $Q_{y_2}$  describing the color-dependence of extinction are small and their variations are also proportionally smaller than the variations of other coefficients in equations (3) and (4). Therefore, instead of using the nightly values of coefficients  $K_2$  and  $Q_{y_2}$  it is more reasonable to use the mean values obtained by averaging the values from many nights. Since, however, the behaviour of these coefficients is not sufficiently well known and since their dependence on  $K_1$  and  $Q_{y_1}$  must be investigated, the coefficients  $K_2$  and  $Q_{y_2}$  are determined for every night on which a sufficient number of primary standards was observed.

\*The coefficients  $K_1$  and  $K_2$  are connected with coefficients  $k_1$  and  $k_2$  in the instrumental system, defined by Johnson (5), by the relations:

$$K_1 = A_2 k_1 - A_1 k_2 \text{ and } K_2 = k_2 .$$

After multiplying equation (3) by  $1 + K_2 M$  we notice that the coefficients  $(B-V)M$  for the unknowns  $K_2$  and  $Q_{y2}$  are not independent of the coefficients of other unknowns. Therefore a simultaneous determination of  $K_2$  and  $Q_{y2}$  with other unknowns of equations (3) and (4) results in a drastic diminution of the weights of these other unknowns, particularly of  $A_2$  and  $A_6$ . To avoid such loss of accuracy, we may write equations (3) and (4) in the form:

$$B-V = -K_2^{\circ}M(B-V) + A'_1 + A_2C_{y0} - K'_1M \quad (10)$$

$$- (K_2 - K_2^{\circ})(M - \bar{M})[(B-V) - (\bar{B}-\bar{V})] ,$$

$$V = m_{y0} + A'_5 + A'_6(B-V) - Q'_{y1}M$$

$$- Q_{y2}(M - \bar{M})[(B-V) - (\bar{B}-\bar{V})] , \quad (11)$$

where  $\bar{M}$  and  $\bar{B}-\bar{V}$  are the mean air mass and the mean color-index of the primary standards observed that night,  $K_2^{\circ}$  is the assumed approximate value of  $K_2$  and the new unknowns are connected with the old ones by the relations:

$$A'_1 = [1 - (K_2 - K_2^{\circ})\bar{M}]A_1$$

$$+ (K_2 - K_2^{\circ})\bar{M}(\bar{B}-\bar{V}) , \quad (12)$$

$$A'_2 = [1 - (K_2 - K_2^{\circ})\bar{M}]A_2 , \quad (13)$$

$$K'_1 = [1 - (K_2 - K_2^{\circ})\bar{M}]K_1$$

$$+ (K_2 - K_2^{\circ})(\bar{B}-\bar{V}) , \quad (14)$$

$$A'_5 = A_5 + Q_{y2}\bar{M}(\bar{B}-\bar{V}) , \quad (15)$$

$$A'_6 = A_6 - Q_{y2}\bar{M} , \quad (16)$$

$$Q'_{y1} = Q_{y1} + Q_{y2}(\bar{B}-\bar{V}) \quad (17)$$

It can be easily shown that equations (10) and (11) are equivalent to equations (3) and (4) if only the term proportional to  $(K_2 - K_2^{\circ})^2$  is neglected. The accuracy with which the primed coefficients are determined from equations (10) and (11) remains practically unchanged whether we determine  $K_2 - K_2^{\circ}$  and  $Q_{y2}$  from these equations or whether we neglect the terms containing  $K_2 - K_2^{\circ}$  and  $Q_{y2}$ . Since the accuracy of determining  $K_2 - K_2^{\circ}$  and  $Q_{y2}$  from one night's observations is very low and since for the purposes of this work (except for improving  $V$  and  $B-V$  of the primary standards) we can assume that  $Q_{y2} = 0$ , the difference between  $A_1$ ,  $A_2$ ,  $A_5$ ,  $A_6$ ,  $K_1$  and  $Q_{y1}$  and the corresponding primed coefficients defined by equations (12) to (17) can be neglected. Henceforth we shall omit the prime symbols and  $A_1$

will mean  $A'_1$  etc.

For a reliable determination of the extinction coefficients for each night, at least two primary standards should be observed at low altitude. It was found most practical to observe such stars at altitudes between  $22^{\circ}$  and  $28^{\circ}$ . The accuracy of such observations is of course much lower than that of the observations made near the zenith. Following Siedentopf (8), we assume that the mean errors of photometric observations are proportional to the air mass for air masses between 1 and 3. Therefore, before making the least-squares solution of equations (10) and (11) for determining the extinction and transformation coefficients, we divide each of these equations by the air mass  $M$ . Moreover, since  $C_{y0}$  is subject to observational errors and  $(B-V)_s$  is assumed to be exactly known, we must transform these equations so that the free term is  $C_{y0}/M$  and not  $(B-V)_s/M$ . The final form\* of these equations which were solved for every night when at least 6 primary standards were observed is

$$(K_1/A_2) - a(A_1/A_2) + B(1/A_2)$$

$$+ c[(K_2 - K_2^{\circ})/A_2] = d , \quad (18)$$

$$- Q_{y1} + aA_5 + BA_6 - cQ_{y2} = D , \quad (19)$$

where the coefficients are

$$a = 1/M , \quad (20)$$

$$b = (B-V)_s [(1/M) + K_2^{\circ}] , \quad (21)$$

$$c = [1 - (\bar{M}/M)] [(B-V)_s - (\bar{B}-\bar{V})_s] , \quad (22)$$

$$d = C_{y0}/M , \quad (23)$$

$$B = (B-V)_s/M , \quad (24)$$

$$D = (V_s - m_{y0})/M \quad (25)$$

One of the main factors limiting the accuracy of photometric observations is the change of extinction during the night. The first step towards eliminating the influence of this change is to assume that extinction is changing linearly during the night. To eliminate extinction changes of this type we need two groups of observations of standard stars during each night, preferably one near the beginning and the other near the end of the observations of the program stars. Each of these groups should consist

\*Unfortunately only for a few nights (in 1961) could the equations in "final" form be solved. For all the other nights the equations (10) and (11) were solved by the least-squares method. This means that for most of the observations of primary standards discussed in this section the change of accuracy with air mass was not taken into consideration. The values of  $A_2$ ,  $A_6$  and  $A_8$  which are used throughout this paper may be subject to small systematic error, usually not exceeding 0.002, resulting from treating  $(B-V)_s$  as a random variable and  $C_{y0}$  as exactly known.

of at least 4 stars, two of them at low altitude and two at high altitude. The least-squares solution of equations (18) and (19) is then made for both groups together, but the extinction coefficients  $K_1$  and  $Q_{y1}$  are determined in this solution for each group separately. When reducing the observations of program stars, values of extinction coefficients interpolated between the two pairs obtained from the least-squares solution are used.

*B. Solution of Transformation Equations.* Since equations (18) and (19) are likely to be used not only by the future observers in this program but also by other photometric observers, it seems worthwhile to give some details of the least-squares solution of these equations.

We assume that the observations of standard stars made on any one night can be divided into two groups; let the number of stars in these groups be denoted by  $n^I$  and  $n^{II}$ . The coefficients  $A_1, A_2, A_3, A_6, K_2 - K_2^0$  and  $Q_{y2}$  are determined for both groups together while  $K_1$  and  $Q_{y1}$  are computed for each group separately, so that two pairs,  $K_1^I, Q_{y1}^I$ , and  $K_1^{II}, Q_{y1}^{II}$ , of these coefficients are obtained. Following Gauss we shall denote (in this section only) by square brackets the sums of the terms in these brackets. If the symbol in brackets has a roman numeral I (or II), this means that it should be summed only over the first (or second) group of standard stars. If there are no roman numerals, the summation is over both groups together.

We introduce the following auxiliary quantities\*

$$\begin{aligned}s_{11} &= (n^I)^{\frac{1}{2}}, \quad s_{22} = (n^{II})^{\frac{1}{2}}, \\s_{31} &= [a^I]/s_{11}, \quad s_{32} = [a^{II}]/s_{22}, \\s_{33} &= [a^2] - s_{31}^2 - s_{32}^2, \\s_{41} &= [b^I]/s_{11}, \quad s_{42} = [b^{II}]/s_{22}, \\s_{43} &= [ab] - s_{31}s_{41} - s_{32}s_{42}, \\s_{51} &= [c^I]/s_{11}, \quad s_{52} = [c^{II}]/s_{22},\end{aligned}$$

$$\begin{aligned}s_{53} &= [ac] - s_{31}s_{51} - s_{32}s_{52}, \\s_{61} &= [d^I]/s_{11}, \quad s_{62} = [d^{II}]/s_{22}, \\s_{63} &= [ad] - s_{31}s_{61} - s_{32}s_{62}, \\s_{44} &= [b^2] - s_{41}^2 - s_{42}^2 - (s_{43}^2/s_{33}), \\s_{54} &= [bc] - s_{41}s_{51} - s_{42}s_{52} - (s_{43}s_{53}/s_{33}), \\s_{64} &= [bd] - s_{41}s_{61} - s_{42}s_{62} - (s_{43}s_{63}/s_{33}), \\s_{55} &= [c^2] - s_{51}^2 - s_{52}^2 - (s_{53}^2/s_{33}) - (s_{54}^2/s_{44}), \\s_{65} &= [cd] - s_{51}s_{61} - s_{52}s_{62} \\&\quad - (s_{53}s_{63}/s_{33}) - (s_{54}s_{64}/s_{44}), \\t_{41} &= [B^I]/s_{11}, \quad t_{42} = [B^{II}]/s_{22}, \\t_{43} &= [aB] - s_{31}t_{41} - s_{32}t_{42}, \\t_{61} &= [D^I]/s_{11}, \quad t_{62} = [D^{II}]/s_{22}, \\t_{63} &= [aD] - s_{31}t_{61} - s_{32}t_{62}, \\t_{44} &= [B^2] - t_{41}^2 - t_{42}^2 - (t_{43}^2/s_{33}), \\t_{54} &= [Bc] - t_{41}s_{51} - t_{42}s_{52} - (t_{43}s_{53}/s_{33}), \\t_{64} &= [BD] - t_{41}t_{61} - t_{42}t_{62} - (t_{43}t_{63}/s_{33}), \\t_{55} &= [c^2] - s_{51}^2 - s_{52}^2 - (s_{53}^2/s_{33}) - (t_{54}^2/t_{44}), \\t_{65} &= [cD] - s_{51}t_{61} - s_{52}t_{62} \\&\quad - (s_{53}t_{63}/s_{33}) - (t_{54}t_{64}/t_{44}).\end{aligned}$$

\*The least-squares solution presented here is based on the cracovian algorithm introduced by Banachiewicz (9) and thoroughly described by Kopal (10). Our  $s_{ij}$  are connected with Kopal's  $r_{ij}$  by the relations:  $s_{ij} = r_{ij}$  for  $j = 1, 2$  and  $s_{ij} = r_{ij}r_{jj}$  for  $j \geq 3$ .

The unknowns are computed from the formulas:

$$\begin{aligned} A_2 &= s_{44}/\left\{s_{64} - (s_{54}s_{65}/s_{55})\right\}, \\ K_2 - K_2^o &= A_2 s_{65}/s_{55}, \\ A_1 &= \left\{s_{43} - A_2 s_{63} + (K_2 - K_2^o)s_{53}\right\}/s_{33}, \\ K_1^I &= \left\{A_1 s_{31} + A_2 s_{61} - (K_2 - K_2^o)s_{51} - s_{41}\right\}/s_{11}, \\ K_1^{II} &= \left\{A_1 s_{32} + A_2 s_{62} - (K_2 - K_2^o)s_{52} - s_{42}\right\}/s_{22}, \\ Q_{y2} &= -t_{65}/t_{55}, \quad A_6 = (t_{64} + Q_{y2} t_{54})/t_{44}, \\ A_5 &= (t_{63} - A_6 t_{43} + Q_{y2} s_{53})/s_{33}, \\ Q_{y1}^I &= (A_5 s_{31} + A_6 t_{41} - Q_{y2} s_{51} - t_{61})/s_{11}, \\ Q_{y1}^{II} &= (A_5 s_{32} + A_6 t_{42} - Q_{y2} s_{52} - t_{62})/s_{22}. \end{aligned}$$

The mean errors of a single observation of B-V and V at zenith, are

$$\begin{aligned} \epsilon_{B-V} &= \left\{ \sum_i^{n^I+n^{II}} (a \delta_{B-V})^2 / (n^I + n^{II} - m) \right\}^{\frac{1}{2}}, \\ \epsilon_V &= \left\{ \sum_i^{n^I+n^{II}} (a \delta_V)^2 / (n^I + n^{II} - m) \right\}^{\frac{1}{2}}, \end{aligned}$$

respectively, where  $(a \delta_{B-V})$  and  $(a \delta_V)$  are deviations, reduced to zenith, which should be computed for every star from the formulas

$$\begin{aligned} a \delta_{B-V} &= a A_1 + d A_2 - c \cdot (K_2 - K_2^o) - b - K_1, \\ a \delta_V &= a A_5 + b A_6 - c Q_{y2} - d - Q_{y1}; \end{aligned}$$

where m denotes the number of unknowns determined from each of the systems of unknowns described in equations (18) and (19). If the stars are divided into two groups, we have  $m = 5$ . For only one group  $n^{II} = 0$  and  $m = 4$ ; if, moreover,  $K_2 - K_2^o$  and  $Q_{y2}$  are not determined we should replace c by zero, omit all the terms  $s_{i2}$ ,  $t_{i2}$  and  $s_{5i}$  ( $i=1, \dots, 6$ ) and assume  $m=3$ .

The mean errors of  $Q_{y2}$  and  $A_6$  are given by

$$\begin{aligned} \epsilon(Q_{y2}) &= \epsilon_V/t_{55}^{\frac{1}{2}}, \\ \epsilon(A_6) &= \epsilon_V \left\{ (1/t_{44}) + (t_{54}^2/t_{44}^2 t_{55}) \right\}^{\frac{1}{2}}; \\ \text{for the mean errors of } K_2 - K_2^o, A_2 \text{ and of the extinction coefficients the approximate expressions can be used:} \\ \epsilon(K_2 - K_2^o) &\approx A_2 \epsilon_{B-V}/s_{55}^{\frac{1}{2}}, \quad \epsilon(A_2) \approx A_2^2 \cdot \\ \epsilon_{B-V} &\left\{ (1/s_{44}) + (s_{54}^2/s_{44}^2 s_{55}) \right\}^{\frac{1}{2}}, \\ \epsilon(K_1^I) &\approx \epsilon_{B-V} / \left\{ \sum_i^n (M - \bar{M})^2 \right\}^{\frac{1}{2}}, \\ \epsilon(Q_{y1}^I) &\approx \epsilon_V / \left\{ \sum_i^n (M - \bar{M})^2 \right\}^{\frac{1}{2}}. \end{aligned}$$

If the unknowns are calculated correctly, the following two control equations are fulfilled:

$$\begin{aligned} [ab] A_1 + [bd] A_2 - [bc] (K_2 - K_2^o) - [b^I] K_1^I \\ - [b^{II}] K_1^{II} = [b^2] - (n^I + n^{II} - m) \epsilon_{B-V}^2, \\ [ad] A_5 + [bd] A_6 - [cd] Q_{y2} - [d^I] Q_{y1}^I \\ - [d^{II}] Q_{y1}^{II} = [d^2] - (n^I + n^{II} - m) \epsilon_V^2; \end{aligned}$$

the terms containing  $\epsilon_{B-V}$  and  $\epsilon_V$  are usually negligible if, as seems to be most convenient, all the calculations are made with four digits after the decimal point.

*C. Observers and Instruments.* The observers working on the photometry of Uranus and Neptune and the instruments used are listed in Table I.

The tube 1P21 No. 12 which has been used since December 1957 is characterized by an exceptionally large dark current when unrefrigerated. When the tube is refrigerated with dry ice the dark current from the cathode amounts to only about 20 electrons per minute. The sensitivity of this tube increases by several per cent, or more, when the tube is illuminated by a bright star, especially if the tube was previously illuminated without applying voltage. When the tube is illuminated by a second magnitude star in the focus of the 21-inch telescope, an appreciable in-

crease of sensitivity can be noticed during the first 10 minutes of illumination. After this time the sensitivity becomes stable and drops down very slowly when the tube is not illuminated. When the tube is initially illuminated by a fainter star, the time necessary for reaching the stable sensitivity is proportionally longer. The sensitivity of multiplier tubes used until 1957 decreased, rather than increased, after illumination by a bright star.

To secure stable sensitivity of the multiplier tube the voltage is applied 1½ hours before observations are started and the tube is illuminated throughout this interval by the standard radioactive light source. Dry ice is put in immediately after applying the voltage. During all the observations, the voltage applied to the multiplier tube is taken from a 900-volt battery constructed from 30-volt batteries connected in series; this is believed to affect the linearity of the multiplier tube less than other types of power supplies.

The calibration of the amplifier and the testing of its linearity were performed at intervals of several months or years. A precision decade resistance box was used for this purpose. The presently used General Radio type 1230-A d.c. amplifier (with gain resistors taken from the previously used amplifier No. 4) is calibrated by a procedure similar to that described by Borgman (11). For determining the exact value of any 2.5-magnitude step, the switch is set in position B (Figure 1) and the fine gain (½ magnitude steps) of the amplifier is adjusted so as to give a full scale deflection of the recorder. The switch is

now changed to A and the resistance box is adjusted so as to give the same deflection of the recorder.

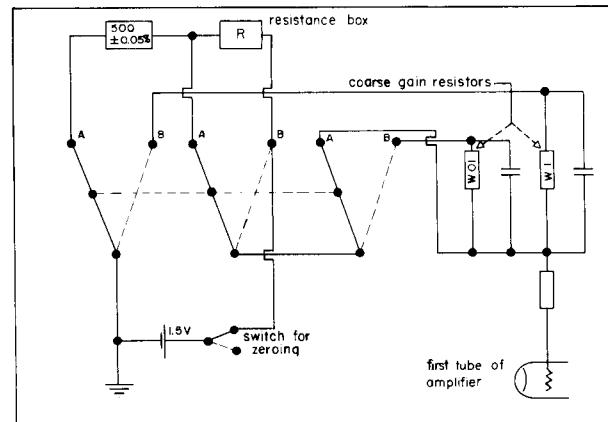


Figure 1. Wiring diagram of simple arrangement for calibrating the General Radio d.c. amplifier and testing its linearity.

The fine gain steps are calibrated and linearity is tested with the switch at position A. For every fine-gain step the resistance box is adjusted so as to give the same deflection of the recorder. Linearity was checked at every fine-gain step separately by changing the setting of the resistance box and comparing the deflections of the recorder with the expected values. No corrections for nonlinearity were applied.

In all the amplifiers used in this program, both coarse-gain and fine-gain resistors were wire-wound.

TABLE I

Observers and Instruments

Period		Observer	Telescope	Tube Type	Tube Desig.
From	To				
1953, Jan. 21	1953, Feb. 11	H. L. Johnson	McDonald 13-inch	1P21	$A_8 = +0.030$
		H. L. Johnson	McDonald 82-inch	1P21	$A_8 = +0.018$
1953, Feb. 23	1953, Apr. 25	H. L. Johnson	Lowell 42-inch	1P21	$A_8 = +0.046$
1953, Jun. 10	1953, Jul. 3	R. H. Hardie	Lowell 21-inch	1P21	$A_8 = +0.024$
1953, Oct. 30	1954, Mar. 6	R. H. Hardie	Lowell 21-inch	1P21 Y	$A_8 = +0.017$
1954, Mar. 15	1954, Jun. 16	R. H. Hardie	Lowell 21-inch	Lall.	$A_8 = +0.063$
1954, Jun. 17	1954, Jul. 1	R. H. Hardie	Lowell 21-inch	1P21 Y	$A_8 = +0.017$
1954, Oct. 11	1955, Apr. 26	R. H. Hardie	Lowell 21-inch	1P21 2—Ref*	$A_8 = +0.073$
1955, Apr. 28	1955, Jul. 8	R. H. Hardie	Lowell 21-inch	1P21 Y—Unref**	
1955, Sep. 23	1956, Jun. 27	C. F. Knuckles	Lowell 21-inch	1P21 10—Ref.	$A_8 = +0.005$
1956, Sep. 22	1957, Jun. 24	R. I. Mitchell	Lowell 21-inch	1P21 10—Ref.	amplifier No. 4
1957, Sep. 21	1957, Dec. 12	W. M. Sinton	Lowell 21-inch	1P21 10—Ref.	amplifier No. 4
1957, Dec. 26	1958, Jul. 5	W. M. Sinton	Lowell 21-inch	1P21 12—Ref.	amplifier No. 4
1958, Oct. 3	1959, Jun. 10	B. Iriarte	Lowell 21-inch	1P21 12—Ref.	amplifier No. 4
1960, Jan. 28	1960, Oct. 20	K. Serkowski	Lowell 21-inch	1P21 12—Ref.	amplifier No. 4
1960, Nov. 17	1961, Jun. 28	J. Priser & K. Serkowski	Lowell 21-inch	1P21 12—Ref.	General Radio Amplifier No. 838

\*Refrigerated

\*\*Unrefrigerated

Lall. = Lallemand

The corrections for the gain steps, based on the calibration of the amplifier, however small, were applied to all the observations.

The coefficients  $A_2$ ,  $A_6$  and  $A_8$ , determined on particular nights, are given in Table II. The dates

in this table, as in all other tables in this paper, refer to U.T. The third column gives the number of observations of primary standard stars used for determining the transformation coefficients. The extinction coefficients,  $K_1$  and  $Q_{y1}$  were determined together with

TABLE II. Transformation Coefficients

TUBE	DATE	n	$A_2$	$A_6$	$A_8$	REMARKS
1955						
1P21 No. 2-U**	Feb. 2	2	1.076	-0.040	+0.031:	
	24	11	1.072	-0.001	+0.066	
1P21 No. 2-R*	Apr. 4	18	1.090	-0.019	+0.064:	sky dusty
	11	7	1.087	-0.008	+0.072	wind
	20	13	1.091	-0.005	+0.079	
	21	17	1.077	-0.006	+0.066	
	25	7	1.090	-0.007	+0.076	
				mean	+0.073	
1P21 Y-U	May 4	11	1.014	+0.010	+0.024	clouds at end
2nd set of filters	6	18	1.012	+0.021	+0.033	
	28	11	1.018	+0.012	+0.030	clouds at end
				mean	+0.029	
1P21 Y-U	Jun. 20	8	1.055	+0.034	+0.086	
1st set of filters	21	8	1.055	+0.035	+0.087	
	22	8	1.051	+0.029	+0.078	
	27	8	1.045	+0.028	+0.071	
				mean	+0.080	
1P21 No. 10-R	Oct. 27	4	1.012	-0.011	+0.001	
	28	4	1.022	-0.017	+0.005	
	30	4	1.015	-0.014	+0.001	
	31	4	1.023	-0.005	+0.017	
	Nov. 4	2	1.024	-0.015	+0.009	
	7	4	1.014	-0.003	+0.011	hazy
	8	4	1.023	-0.010	+0.012	
	16	2	1.020	-0.018	+0.002	
	19	2	1.018	-0.015	+0.003	
	20	2	1.027	-0.021	+0.005	
	27	2	1.028	-0.022	+0.005	hazy
1956						
	Jan. 17	25	1.029	-0.018	+0.008	
	19	4	1.021	0.000	+0.021:	
	20	4	1.027	+0.006	+0.032:	
	Feb. 3	19	1.006	-0.004	+0.002	
	10	22	1.026	-0.019	+0.006	
	14	23	1.054	-0.031	+0.020	
	Mar. 5	5	1.021	-0.024	-0.003	
	14	8	1.011	-0.006	+0.005	clouds
	19	36	1.006	+0.003	+0.009	
	Apr. 5	24	1.012	-0.005	+0.006	
	18	12	1.005	-0.011	-0.006	clouds
				mean	+0.005	

\*Refrigerated

\*\*Unrefrigerated

TABLE II. Transformation Coefficients (Cont'd)

TUBE	DATE 1957	n	A <sub>2</sub>	A <sub>6</sub>	A <sub>8</sub>	REMARKS
1P21 No. 10-R	Mar. 23	13	1.024	-0.014	+0.010	
	24	15	1.019	-0.004	+0.015	wind
	26	3	1.012	-0.008	+0.004	clouds at end
	28	4	1.011	+0.018	+0.029	
	31	3	1.022	-0.013	+0.009	clouds
	Apr. 5	13	1.015	-0.001	+0.014	poor seeing
	6	8	1.015	+0.013	+0.028	
	7	4	1.018	-0.006	+0.012	
	9	3	1.018	-0.015	+0.003	clouds at end
	15	10	1.015	0.000	+0.015	
	16	3	1.019	-0.014	+0.005	clouds at end
	18	7	1.013	+0.011	+0.024	clouds at end
	19	4	1.011	-0.003	+0.008	clouds at end
	20	5	1.020	+0.003	+0.017	clouds at end
	24	4	1.018	+0.030	+0.048	
	26	5	1.008	-0.002	+0.006	clouds at end
	May 1	10	1.015	-0.007	+0.008	extinction changing
	3	8	1.017	-0.002	+0.015	
	5	6	1.015	-0.006	+0.009	clouds
mean					+0.015	
	May 30	5	1.009	+0.019	+0.028	
	Jun. 1	4	1.016	+0.016	+0.032	
	13	2	1.009	+0.018	+0.027	
	14	2	1.022	+0.007	+0.029	
mean					+0.029	
	Jun. 22	4	1.014	+0.026	+0.040	
	24	4	1.016	+0.025	+0.037	
	25	6	1.012	+0.021	+0.037	
mean					+0.038	
1P21 No. 12-R	Dec. 27	5	1.043	-0.070	-0.029:	clouds ?
	1958					
	Jan. 8	6	1.046	-0.046	-0.002	clouds ?
	9	8	1.036	-0.027	+0.008	clouds ?
	16	9	1.039	-0.035	+0.003	
	Feb. 11	8	1.039	-0.036	+0.002	
	27	9	1.043	-0.016	+0.025	clouds ?
	Mar. 20	9	1.046	-0.023	+0.021	clouds
	Apr. 10	10	1.038	-0.027	+0.010	
	11	10	1.041	-0.026	+0.013	
	15	10	1.043	-0.025	+0.016	
	May 4	10	1.042	-0.029	+0.011	
	9	10	1.044	-0.023	+0.019	
	10	10	1.047	-0.027	+0.018	
mean					+0.013	
1959						
1P12 No. 12-R	Jan. 5	4	1.025	-0.034	-0.010	
	11	4	1.043	-0.034	+0.007	
	24	4	1.047	-0.048	-0.003	
	29	2	1.031	-0.044	-0.014	
	31	4	1.047	-0.043	+0.002	clouds

TABLE II. Transformation Coefficients (Cont'd)

TUBE	DATE	n	A <sub>2</sub>	A <sub>6</sub>	A <sub>8</sub>	REMARKS
1P21 No. 12-R	1959					
Feb.	2	2	1.061	-0.051	+0.007	clouds
	4	4	1.051	-0.046	+0.003	
	5	4	1.049	-0.049	--0.002	
	26	4	1.040	-0.038	0.000	
	28	4	1.043	-0.032	+0.003	
	Mar. 1	4	1.039	-0.041	-0.003	
	3	4	1.041	-0.041	-0.002	
				mean	0.000	
	May 9	2	1.034	-0.009	+0.024	
	15	2	1.044	--0.015	+0.027	
1960	16	2	1.041	-0.028	+0.011	clouds
	27	2	1.068	-0.019	+0.045:	
	28	2	1.061	-0.015	+0.043:	
	Jan. 20	2	1.026	-0.067	-0.042	clouds at end
	28	4	1.026	-0.045	-0.020	
	29	2	1.022	-0.052	--0.020	
	Feb. 17	6	1.023	-0.040	-0.018	
	18	2	1.026	-0.051	-0.026	
	21	4	1.021	-0.040	-0.019	
	23	6	1.025	-0.041	-0.017	
Mar.	25	2	1.030	-0.052	-0.023	clouds
	15	6	1.027	-0.042	-0.016	
	17	6	1.023	-0.040	-0.018	
	18	8	1.018	-0.035	-0.017	
	19	6	1.021	-0.032	-0.011	
	21	6	1.024	-0.034	-0.010	
				mean	-0.020	
	Mar. 26	4	1.028	-0.024	+0.003	
	30	6	1.026	-0.030	-0.005	
Apr.	3	6	1.033	-0.038	-0.006	clouds at end
	4	2	1.017	-0.046	-0.029	
	6	4	1.044	-0.038	+0.004	
				mean	-0.007	
	1961					
Feb.	6	6	1.024	-0.052	-0.029	clouds at end
	8	9	1.028	-0.036	-0.009	
	10	8	1.044	-0.027	-0.009	
	15	4	1.022	-0.046	-0.024	
	22	4	1.005	-0.041	-0.036	
	24	3	1.038	-0.057	-0.020	
	Mar. 1	8	1.032	-0.050	-0.019	
	8	4	1.034	-0.050	-0.017	
	23	8	1.027	-0.032	-0.006	
	31	4	1.029	-0.057	-0.029	
Apr.	4	6	1.023	-0.040	-0.018	clouds at end
	9	7	1.024	-0.040	-0.017	
	12	13	1.024	-0.036	-0.013	
				mean	-0.019	

TABLE III. Extinction Coefficients

DATE	$\Sigma(M - \bar{M})$	Err of single observation		Extinction for B-V colors		Extinction for V magnitudes		Extinction for B mags.		Coefficients of color-dependence of extinction		
		$\epsilon_{B-V}$ m.e.	$\epsilon_V$ m.e.	$K_1$ m.e.	$Q_{y1}$ m.e.	$Q_{bl}$	$K_2$ m.e.	$Q_{y2}$ m.e.				
1955												
Feb.	24	0.4	+0.006	+0.018	0.099	+0.009	0.145	+0.029	0.237			
Apr.	4	0.7	+0.014	+0.025	0.161	+0.017	0.138	+0.031	0.286			
	21	1.0	+0.015	+0.017	0.117	+0.014	0.132	+0.017	0.239			
May	6	5.9	+0.010	+0.009	0.113	+0.004	0.176	+0.004	0.287	-0.033	+0.005	+0.002
Oct.	27	0.8	+0.006	+0.015	0.126	+0.007	0.133	+0.018	0.257			
	28	0.9	+0.007	+0.009	0.129	+0.007	0.120	+0.011	0.246			
	30	3.1	+0.008	+0.015	0.128	+0.005	0.114	+0.009	0.240			
	31	1.4	+0.006	+0.007	0.126	+0.005	0.184	+0.006	0.207			
Nov.	7	0.9	+0.008	+0.001	0.117	+0.008	0.089	+0.001	0.204			
	8	0.7	+0.012	+0.017	0.088	+0.014	0.124	+0.021	0.210			
1956												
Jan.	17	3.5	+0.009	+0.010	0.108	+0.004	0.138	+0.005	0.243	-0.023	+0.006	-0.007
Feb.	3	1.2	+0.011	+0.024	0.118	+0.007	0.151	+0.018	0.268			
	10	3.5	+0.009	+0.015	0.117	+0.004	0.158	+0.008	0.272	-0.017	+0.006	0.000
	14	2.2	+0.011	+0.016	0.102	+0.007	0.164	+0.011	0.261	-0.034	+0.010	+0.008
Mar.	19	9.8	+0.008	+0.012	0.122	+0.002	0.144	+0.004	0.265	-0.028	+0.004	-0.010
	19	0.5	+0.011	+0.009	0.105	+0.015	0.158	+0.012	0.262			
Apr.	5	1.0	+0.011	+0.017	0.128	+0.010	0.198	+0.017	0.325			
	5	1.0	+0.006	+0.016	0.128	+0.005	0.172	+0.016	0.298			
1957												
Mar.	23	2.6	+0.005	+0.010	0.109	+0.003	0.185	+0.006	0.291	-0.028	+0.005	+0.003
	24	1.0	+0.008	+0.014	0.138	+0.008	0.142	+0.014	0.277			
Apr.	5	0.8	+0.010	+0.014	0.110	+0.010	0.136	+0.016	0.244			
	6	0.4	+0.009	+0.016	0.124	+0.015	0.160	+0.028	0.282			
	15	1.3	+0.007	+0.011	0.132	+0.006	0.212	+0.010	0.342			
	26	1.5	+0.005	+0.008	0.132	+0.004	0.264	+0.007	0.395			
May	1	1.3	+0.009	+0.027	0.145	+0.008	0.296	+0.024	0.439			
	3	0.4	+0.007	+0.017	0.113	+0.011	0.178	+0.027	0.289			
1958												
Jan.	8	1.2	+0.021	+0.013	0.116	+0.019	0.121	+0.012	0.232			
	9	0.8	+0.010	+0.021	0.132	+0.010	0.156	+0.027	0.283			
	16	1.0	+0.006	+0.013	0.116	+0.006	0.134	+0.013	0.246			
Apr.	10	1.3	+0.007	+0.011	0.110	+0.006	0.258	+0.010	0.364	-0.026	+0.008	-0.002
	11	2.1	+0.004	+0.014	0.105	+0.003	0.257	+0.010	0.358	-0.031	+0.003	+0.015
	15	3.5	+0.008	+0.007	0.121	+0.004	0.174	+0.004	0.290	-0.034	+0.005	+0.007
May	4	3.6	+0.009	+0.016	0.130	+0.004	0.187	+0.009	0.312	-0.030	+0.006	+0.007
	9	5.0	+0.008	+0.016	0.127	+0.003	0.219	+0.007	0.341	-0.032	+0.004	+0.008
	10	7.2	+0.008	+0.025	0.116	+0.003	0.228	+0.009	0.339	-0.030	+0.004	+0.002
1959												
Feb.	28	0.6	+0.009	+0.011	0.118	+0.011	0.127	+0.014	0.240			
1960												
Mar.	15	0.4	+0.005	+0.010	0.088	+0.007	0.167	+0.015	0.253			
	17	0.5	+0.008	+0.015	0.114	+0.011	0.203	+0.020	0.314			
	18	0.5	+0.004	+0.022	0.125	+0.006	0.172	+0.032	0.295			
	19	0.4	+0.008	+0.016	0.090	+0.012	0.140	+0.026	0.228			
	21	0.6	+0.005	+0.012	0.105	+0.005	0.142	+0.015	0.244			
1961												
Jan.	10	1.4		0.117		0.105		0.219				
	11	0.4		0.105		0.151		0.253				
Feb.	6	1.8	+0.005	+0.007	0.122	+0.004	0.147	+0.005	0.266	-0.032	+0.012	+0.008
	8	2.8	+0.007	+0.012	0.097	+0.004	0.107	+0.007	0.201	-0.057	+0.013	-0.055
	10	1.3	+0.005	+0.009	0.130	+0.004	0.131	+0.008	0.258	-0.029	+0.012	-0.001
	15	1.1	+0.005	+0.007	0.115	+0.005	0.146	+0.007	0.258			
	24	0.5		0.115	+0.012	0.171	+0.014	0.283				
Mar.	1	2.5	+0.009	+0.009	0.094	+0.006	0.097	+0.006	0.189	-0.057	+0.009	-0.092
	22	1.8		0.126		0.160		0.283				
	23	4.7	+0.009	+0.018	0.108	+0.004	0.133	+0.008	0.238			
	31	2.9	+0.011	+0.005	0.102	+0.007	0.134	+0.003	0.233			
Apr.	4	1.9	+0.004	+0.008	0.129	+0.003	0.135	+0.006	0.261	-0.044	+0.010	-0.013
	9	3.4	+0.004	+0.012	0.127	+0.002	0.147	+0.006	0.271			
	12	8.4	+0.005	+0.014	0.097	+0.003	0.147	+0.007	0.241	-0.017	+0.005	+0.003
	29	0.5		0.078		0.105		0.241				
May	9	1.0		0.110		0.151		0.258				
	16	0.5		0.131		0.202		0.330				
	17	2.5		0.119		0.214		0.330				
	19	0.9		0.111		0.179		0.287				
	21	4.5		0.122		0.177		0.296				
Jun.	2	1.1		0.120		0.157		0.274				
	8	1.1		0.113		0.175		0.285				
	10	0.8		0.115		0.247		0.359				
	28	0.8		0.146		0.226		0.368				

the transformation coefficients on only those nights for which  $\Sigma(M - \bar{M})^2$  for the primary standards was larger than 0.3. These nights are listed in Table III.

Examining Table II we notice that the values of  $A_8$  are usually higher in the summer months when the temperature of the air is high and refrigeration with dry ice is not so effective as in the winter. Numerous experiments made by H. L. Johnson proved that for most multiplier tubes the transformation coefficients depend strongly on temperature when the tubes are not refrigerated.

*D. Extinction Coefficients.* The values of the extinction coefficients determined from the observations of primary standard stars are given with their mean errors in Table III. The second column gives  $\Sigma(M - \bar{M})^2$  which can be considered as the weight of the extinction determination. The third and fourth columns give the mean errors of a single observation of B-V and V as obtained from the least-squares solution. The number of observations used in the solution can be found in the third column of Table II. The mean errors are not given in Table III for those nights when only 2 or 3 primary standards were observed. The values of  $K_1$ ,  $Q_{y1}$  and  $Q_{b1}$  listed in the 5th, 6th and 7th columns of Table III are plotted as a function of season in Figures 2, 3 and 4. The coefficient  $K_1$  does not seem to vary with the season. Its mean value can be assumed as 0<sup>m</sup>115. The mean seasonal values of the extinction coefficients with yellow and blue filters (coefficients  $Q_{y2}$  and  $Q_{b2}$  used for reducing the observations are:

	Yellow $Q_{y1}$	Blue $Q_{b1}$
October, November	0 <sup>m</sup> 120	0 <sup>m</sup> 232
December through March 10	.138	.250
March 11 through April 20	.150	.262
April 21 through July	.190	.300

The dependence of extinction on humidity was investigated but no clearly expressed correlation was found.

The last two columns of Table III give the coefficients  $K_2$  and  $Q_{y2}$  which describe the color-dependence of extinction;  $K_2 = -0.030$  and  $Q_{y2} = -0.002$  are their mean values. In reducing the observations it was, however, assumed that  $Q_{y2} = 0$  and that  $Q_{b2}$  had the value  $-0.033$ .

### III. MAGNITUDES AND COLORS OF STANDARD AND COMPARISON STARS

*A. Accuracy of Two-Color Observations.* The magnitudes and colors of the Ten-Year Standards and comparison stars were computed from equations (3) and (4), where the values  $K_2 = -0.03$  and  $Q_{y2} = 0$  were always assumed. The values of  $A_1$ ,  $A_2$ ,  $A_5$  and  $A_6$  were determined from the observations of primary standards for every night when two color observations were made. They were never averaged over several nights.

The nightly values of the extinction coefficients  $K_1$  and  $Q_{y1}$  were used only for those nights which are listed in Table III. For other nights the seasonal

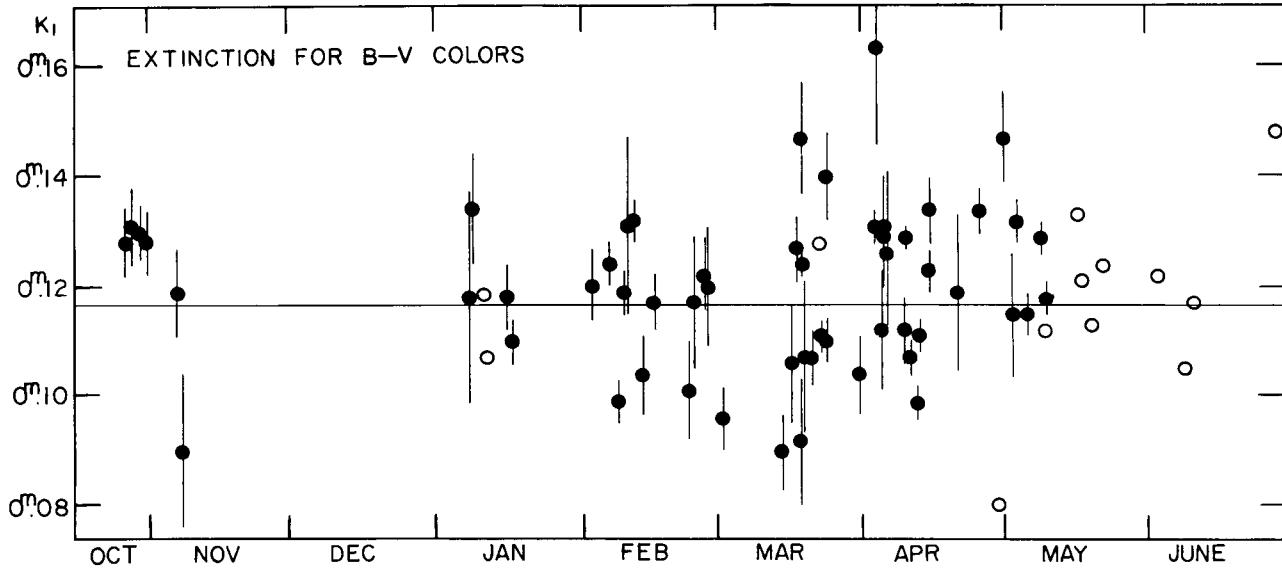


Figure 2. The atmospheric extinction coefficients  $K_1$  for the blue-yellow color index as determined at Lowell Observatory. Vertical lines show the mean errors of nightly values. Open circles represent values based on 2 or 3 standard stars, filled circles those based on 4 or more stars. Horizontal line represents the assumed mean value.

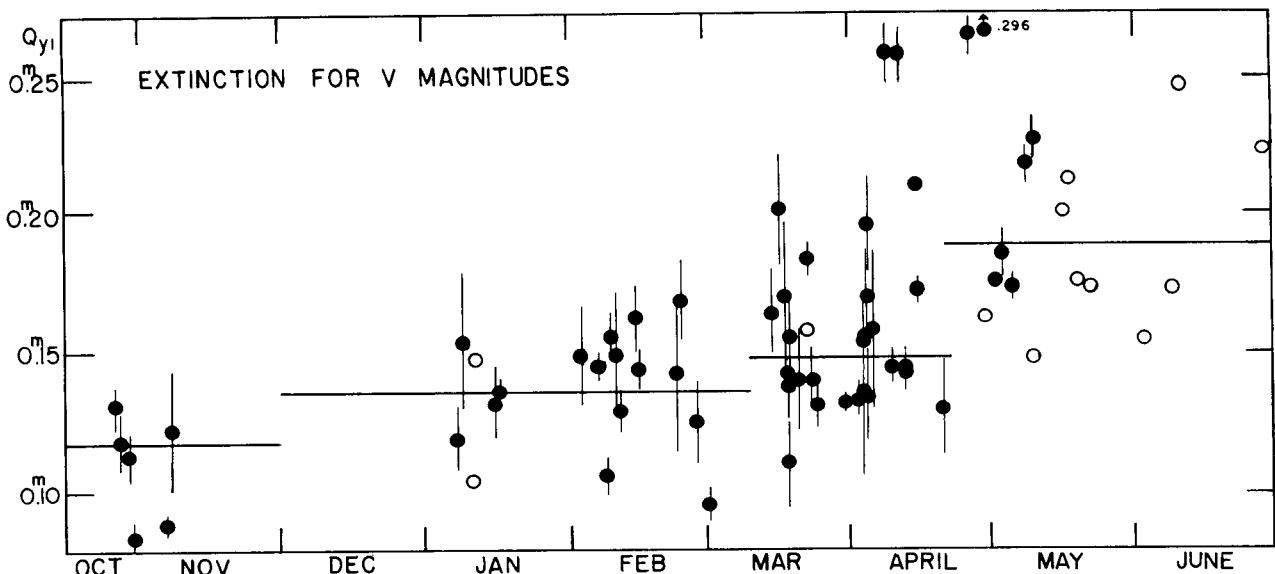


Figure 3. The atmospheric extinction coefficients  $Q_{y1}$  for the yellow magnitudes. Notations are the same as in Figure 2. Horizontal lines represent the assumed mean seasonal values.

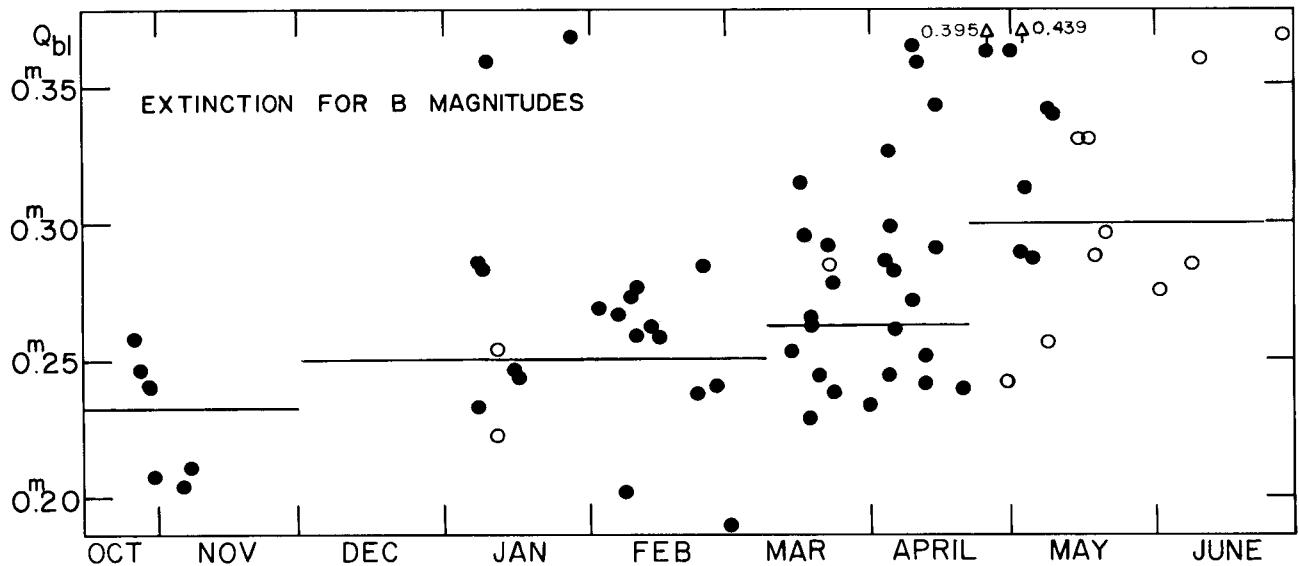


Figure 4. The atmospheric extinction coefficients  $Q_{b1}$  for the blue magnitudes. Notations are the same as in Figures 2 and 3.

mean values were used. From the scatter in Figures 2 and 3, it can be roughly estimated that the r.m.s. deviations of nightly values of  $K_1$  and  $Q_{y1}$  from seasonal averages are  $\sigma(K_1) = \pm 0^m 025$  and  $\sigma(Q_{y1}) = \pm 0^m 04$ , respectively. The mean errors for a typical observation of the color and magnitude of a bright star at the zenith can be assumed to be  $\epsilon_{B-V} (M=1) = \pm 0^m 0075$  and  $\epsilon_V (M=1) = \pm 0^m 012$ . These values are used for computing the weights of the observations of the Ten-Year Standards and comparison stars, as described below.

We denote by  $\bar{M}$  the mean air mass for the primary standard stars used for determining the transformation coefficients  $A_1$ ,  $A_2$ ,  $A_5$  and  $A_6$ . If these coefficients were the only unknowns determined from the equations (3) and (4),  $A_1$  and  $A_5$  would be distorted by any deviation of extinction coefficients  $K_1$  and  $Q_{y1}$  from the assumed seasonal mean values  $K_1^\circ$  and  $Q_{y1}^\circ$ . Instead of the true values for  $A_1$  and  $A_5$ , we obtain from the least-squares solution of equations (3) and (4) the coefficients  $A''_1$  and  $A''_5$  connected with the true coefficients by the relations:

$$A''_1 = A_1 - (K_1 - K_1^{\circ}) \bar{M} , \quad (21)$$

$$A''_5 = A_5 - (Q_{y1} - Q_{y1}^{\circ}) \bar{M} , \quad (22)$$

which are valid when the deviations of  $K_2$  and  $Q_{y2}$  from their assumed values can be neglected. The equations (3) and (4) now take the form

$$(B-V)(1+K_2^{\circ}M) = A''_1 + A_2 C_{y_0} - K_1 - K_1^{\circ} (\bar{M} - \bar{M}) , \quad (23)$$

$$V - m_{y_0} = A''_5 + A_6 (B-V) - (Q_{y1} - Q_{y1}^{\circ}) (\bar{M} - M) . \quad (24)$$

Assuming the seasonal mean values of extinction coefficients we neglect the last terms of equations (23) and (24). Therefore, if we assume that the errors of photometric measurements are proportional to the air mass, the mean errors of the color and magnitude computed by taking into account the uncertainty of extinction are

Table IV. Observations of Ten Year Standards

DATE	$\rho$ Gem				
	V System of primary stds.	B-V System of primary stds.	Wt.	V System of 10-year stds.	B-V System of 10-year stds.
1955					
Feb. 24	4.169	+0.322	3	4.159	+0.319
Apr. 4	4.210	.328	4	4.214	.318
11	4.170	.308	4		
20	4.162	.311	5	4.171	.310
21	4.191	.326	4	4.192	.322
May 4	4.169	.327	3	4.172	.330
6	4.168	.321	3	4.163	+0.325
			mean 4.180		+0.320
1956					
Jan. 17	4.173	.321	4	4.181	+0.318
Feb. 3	4.190	.298	4	4.188	.300
10	4.193	.314	4	4.184	.313
14	4.193	.308	4	4.183	.304
Mar. 5	4.140	.316	2	4.178	.321
19	4.179	.323	3	4.182	.315
Apr. 5	4.194	.300	4	4.187	.302
18	4.174	.333	2	4.170	+0.342
			mean 4.183		+0.312
1957					
Mar. 23	4.182	.302	5	4.181	+0.304
24	4.194	.311	4	4.199	.305
Apr. 5	4.167	.312	4		
9	4.187	.316	2		
15	4.166	.304	3	4.163	.308
20	4.157	.308	2		

$$\varepsilon_{B-V}(M) = [M^2 \varepsilon_{B-V}^2(M=1) + (M - \bar{M})^2 \sigma^2(K_1)]^{1/2} , \quad (25)$$

$$\varepsilon_V(M) = [M^2 \varepsilon_V^2(M=1) + (M - \bar{M})^2 \sigma^2(Q_{y1})]^{1/2} . \quad (26)$$

If the weight of the observation at the zenith is taken as 5, the weight of an observation at air mass  $M$  is

$$W = \frac{5}{M^2 + (M - \bar{M})^2 [\sigma(Q_{y1})/\varepsilon_V(M=1)]^2} \quad (27)$$

This equation was used for computing the weights of all observations given in Tables IV and V. For

DATE	$\rho$ Gem				
	V System of primary stds.	B-V System of primary stds.	Wt.	V System of 10-year stds.	B-V System of 10-year stds.
1957					
May 1	4.188	+0.331	2	4.177	0.327
	3	4.191	.312	4.194	+0.311
			mean 4.184		+0.309
Dec. 27	4.192	.320	4		
1958					
Jan. 9	4.176	.331	3		
	16	4.177	.315	5	4.180 +0.314
Feb. 11	4.179	.313	2	4.188	.307
	27	4.149	.311	4	
Mar. 20	4.181	.309	2	4.175	.307
Apr. 10	4.178	.314	4	4.178	.309
	11	4.168	.321	4	4.168 .315
	15	4.175	.308	4	4.171 .304
May 4	4.194	.318	3	4.196	.317
	9	4.190	.302	3	4.187 .310
	10	4.226	.311	3	4.220 +0.306
1959					
		mean 4.183			+0.310
Jan. 5	4.141	.322	4	4.173	+0.326
	12	4.185	.321	2	4.190 .323
	24	4.187	.301	3	4.191 .317
	29	4.177	.309	5	
	31	4.183	.312	4	4.192 .317
Feb. 2	4.187	.308	4		
	5	4.168	.331	4	4.173 .332
	26	4.167	.322	4	4.183 .314

Table IV. Observations of Ten Year Standards (Cont'd)

		$\rho$ Gem				10 U Ma			
DATE		V System of primary stds.	B-V	V System of 10-year stds.	B-V	V System of primary stds.	B-V	V System of 10-year stds.	B-V
		Wt.		Wt.		Wt.		Wt.	
1959									
Feb.	28	4.215	+ 0.330	3	4.209	+ 0.331			
Mar.	1	4.185	.313	3	4.187	.323			
	3	4.195	.316	3	4.218	+ 0.330			
				mean	4.189	+ 0.324			
1960									
Jan.	20	4.184	.315	4					
	28	4.181	.314	3	4.184	+ 0.314			
Feb.	21	4.190	.321	2	4.181	.321			
	23	4.192	.315	1	4.193	.313			
Mar.	15	4.170	.317	2	4.171	.324			
	17	4.179	.315	2	4.173	.321			
	18	4.201	.314	3	4.186	.313			
	19	4.173	.320	3	4.181	.321			
	21	4.170	.340	1	4.176	.315			
	26	4.169	.327	0	4.162	.329			
	30	4.178	.313	2	4.183	.315			
Apr.	3	4.169	.309	3	4.175	.313			
	6	4.200	.319	2	4.192	+ 0.325			
				mean	4.181	+ 0.318			
1961									
Feb.	6	4.183	.310	5	4.183	+ 0.308			
	8	4.211	.311	4	4.206	.302			
	10	4.181	.315	4	4.174	.309			
	15	4.186	.314	3	4.173	.314			
	21	4.181	.315	4					
	22	4.167	.313	4	4.177	.309			
	24	4.165	.318	3	4.173	.304			
Mar.	1	4.183	.309	4	4.176	.306			
	8	4.181	.313	4	4.169	.311			
	23	4.158	.312	4	4.154	.311			
	31	4.189	.310	3	4.183	.325			
Apr.	4	4.192	.323	3	4.185	.318			
	9	4.171	.306	2	4.177	.312			
	12	4.183	.316	4	4.180	+ 0.315			
				mean	4.177	+ 0.311			
		10 U Ma							
DATE		V System of primary stds.	B-V	V System of 10-year stds.	B-V	1959		mean	
		Wt.		Wt.		Jan.		3.980	+ 0.427
1955									
Feb.	24	3.966	+ 0.432	4	3.956	+ 0.429			
Apr.	4	3.967	.453	3	3.971	.443			
	11	3.974	.433	3					
	20	3.977	.440	4	3.986	.439			
May	4	3.966	.433	3	3.969	.436			
	6	3.964	.436	4	3.959	.440			

**Table IV. Observations of Ten Year Standards (Cont'd)**

10 U Ma								11 L Mi							
DATE		V	B-V	V		B-V		V	B-V	V		B-V			
		System of primary stds.	Wt.	System of 10-year stds.				System of primary stds.	Wt.	System of 10-year stds.					
1959															
Feb.	26	3 <sup>m</sup> .947	+ 0 <sup>m</sup> .445	5	3 <sup>m</sup> .963	+ 0 <sup>m</sup> .437		May	6	5 <sup>m</sup> .402	+ 0 <sup>m</sup> .771	4	5 <sup>m</sup> .397	+ 0 <sup>m</sup> .775	
	28	4 .001	.433	3	3 .995	.434		28	5 .414	.782	3	5 .413	+ 0 .780		
Mar.	1	3 .977	.432	2	3 .979	.442						mean	5 .412	+ 0 .772	
	3	3 .974	.435	4	3 .997	+ 0 .449									
				mean	3 .976	+ 0 .439									
1960								1956							
Jan.	20	3 .975	.430	4				Jan.	17	5 .406	.772	4	5 .414	+ 0 .769	
	28	3 .981	.427	3	3 .984	+ 0 .427		Feb.	3	5 .394	.780	4	5 .392	.782	
Feb.	21	3 .984	.438	4	3 .975	.438		10	5 .412	.775	4	5 .403	.774		
	23	3 .974	.442	4	3 .975	.440		14	5 .422	.764	5	5 .412	.760		
Mar.	15	3 .968	.429	3	3 .969	.436		Mar.	5	5 .380	.755	2	5 .418	.760	
	17	3 .985	.429	3	3 .979	.435		19	5 .414	.770	4	5 .417	.762		
	18	4 .007	.436	2	3 .992	.435		Apr.	5	5 .407	.766	4	5 .400	.768	
	19	3 .958	.441	2	3 .966	.442		18	5 .440	.756	2	5 .436	+ 0 .765		
	21	3 .966	.464	2	3 .972	.439						mean	5 .409	+ 0 .768	
	26	3 .977	.437	1	3 .970	.439		1957							
	30	3 .962	.440	3	3 .967	.442		Mar.	23	5 .402	.763	5	5 .401	+ 0 .765	
Apr.	3	3 .959	.427	4	3 .965	.431		24	5 .420	.770	4	5 .425	.764		
	6	3 .989	.420	3	3 .981	+ 0 .426		Apr.	5	5 .416	.773	4			
				mean	3 .974	+ 0 .435		9	5 .415	.776	2				
1961								15	5 .414	.758	3	5 .411	.762		
Feb.	6	3 .974	.434	4	3 .974	+ 0 .432		20	5 .426	.778	2				
	8	3 .965	.436	3	3 .960	.427		26	5 .424	.760	1				
	10	3 .983	.435	3	3 .976	.429		May	1	5 .459	.777	3	5 .448	.773	
	15	3 .987	.425	4	3 .974	.425		3	5 .425	.766	4	5 .428	+ 0 .765		
	21	3 .973	.435	4								mean	5 .421	+ 0 .766	
	22	3 .957	.442	4	3 .967	.438		Dec.	27	5 .397	.769	4			
	24	3 .968	.444	4	3 .976	.430		1958							
Mar.	1	3 .973	.428	5	3 .966	.425		Jan.	8	5 .414	.766	1			
	8	3 .981	.431	4	3 .969	.429		9	5 .395	.783	3				
	23	3 .947	.434	3	3 .943	.433		16	5 .426	.756	4	5 .429	+ 0 .755		
	31	3 .980	.418	3	3 .974	.433		Feb.	11	5 .396	.778	2	5 .405	.772	
Apr.	4	3 .980	.436	4	3 .973	.431		27	5 .393	.770	4				
	9	3 .965	.434	3	3 .971	.440		Mar.	20	5 .420	.778	2	5 .414	.776	
	12	3 .981	.429	5	3 .978	+ 0 .428		Apr.	10	5 .424	.762	4	5 .424	.757	
				mean	3 .970	+ 0 .431		11	5 .411	.777	4	5 .411	.771		
								15	5 .428	.782	4	5 .424	.778		
								May	4	5 .407	.773	3	5 .409	.772	
								9	5 .418	.757	3	5 .415	.765		
								10	5 .433	.770	3	5 .427	+ 0 .765		
												mean	5 .419	+ 0 .767	
11 L Mi															
DATE		V	B-V	V		B-V									
		System of primary stds.	Wt.	System of 10-year stds.											
1955								1959							
Feb.	24	5 <sup>m</sup> .412	+ 0 <sup>m</sup> .775	2	5 <sup>m</sup> .402	+ 0 <sup>m</sup> .772		Jan.	5	5 .367	.763	3	5 .399	+ 0 .767	
Apr.	4	5 .413	.778	3	5 .417	.768		12	5 .381	.788	2	5 .386	.790		
	11	5 .407	.766	4				24	5 .394	.758	2	5 .398	.774		
	20	5 .412	.771	5	5 .421	.770		29	5 .382	.768	5				
	21	5 .416	.775	4	5 .417	.771		31	5 .384	.768	4	5 .393	.773		
May	4	5 .406	.764	4	5 .409	.767		Feb.	2	5 .401	.775	4			
								4	5 .398	.765	5	5 .413	.770		
								5	5 .409	.775	5	5 .414	.776		

**Table IV. Observations of Ten Year Standards (Cont'd)**

11 L Mi								40 Leo							
DATE		V	B-V	V	B-V	V	B-V	V	B-V	Wt.	Wt.	V	B-V		
		System of primary stds.	Wt.	System of 10-year stds.	Wt.	System of primary stds.	Wt.								
1959															
Feb.	26	5 <sup>m</sup> 402	+ 0 <sup>m</sup> .763	5	5 <sup>m</sup> .418	+ 0 <sup>m</sup> .755	Jan.	17	4 <sup>m</sup> .783	+ 0 <sup>m</sup> .456	4	4 <sup>m</sup> .791	+ 0 <sup>m</sup> .453		
	28	5 .441	.764	3	5 .435	.765	Feb.	3	4 .788	.459	4	4 .786	.461		
Mar.	1	5 .405	.757	2	5 .407	.767		10	4 .793	.454	4	4 .784	.453		
	3	5 .404	.752	4	5 .427	+ 0 .766		14	4 .801	.441	4	4 .791	.437		
1960				mean	5 .411	+ 0 .769	Mar.	5	4 .759	.441	2	4 .797	.446		
Jan.	20	5 .412	.772	4				14	4 .771	.453	4				
	28	5 .402	.776	3	5 .405	+ 0 .776		19	4 .783	.466	3	4 .786	.458		
Feb.	17	5 .406	.765	2	5 .404	.770	Apr.	5	4 .804	.445	3	4 .797	.447		
	21	5 .419	.776	4	5 .410	.776		18	4 .801	.436	2	4 .797	+ 0 .445		
Mar.	15	5 .412	.769	3	5 .413	.776						mean	4 .790	+ 0 .451	
	17	5 .418	.756	3	5 .412	.762	1957								
	18	5 .428	.771	2	5 .413	.770	Mar.	23	4 .794	.444	5	4 .793	+ 0 .446		
	19	5 .398	.774	2	5 .406	.775		24	4 .786	.464	4	4 .791	.458		
	21	5 .404	.795	3	5 .410	.770	Apr.	5	4 .782	.446	4				
	26	5 .422	.766	2	5 .415	.768		9	4 .802	.444	2				
	30	5 .408	.763	3	5 .413	.765		15	4 .790	.441	3	4 .787	.445		
Apr.	3	5 .406	.758	4	5 .412	.762		20	4 .786	.454	2				
	6	5 .419	.766	3	5 .411	+ 0 .772		26	4 .797	.453	1				
1961				mean	5 .410	+ 0 .770	May	1	4 .822	.470	2	4 .811	.466		
Feb.	6	5 .400	.774	4	5 .400	+ 0 .772		3	4 .788	.442	4	4 .791	.441		
	8	5 .406	.785	3	5 .401	.776		5	4 .799	.445	1	4 .788	.448		
	10	5 .413	.784	3	5 .406	.778		30	4 .807	.448	2	4 .807	+ 0 .448		
	15	5 .425	.762	4	5 .412	.762						mean	4 .794	+ 0 .449	
	21	5 .400	.767	4			Dec.	27	4 .792	.447	4				
	22	5 .389	.766	4	5 .399	.762	1958								
	24	5 .395	.773	4	5 .403	.759	Jan.	8	4 .793	.451	1				
Mar.	1	5 .402	.768	5	5 .395	.765		9	4 .784	.458	3				
	8	5 .402	.768	4	5 .390	.766		16	4 .784	.453	4	4 .787	+ 0 .452		
	23	5 .390	.771	3	5 .386	.770	Feb.	11	4 .782	.461	2	4 .791	.455		
	31	5 .402	.759	4	5 .396	.774		27	4 .768	.457	4				
Apr.	4	5 .402	.776	5	5 .395	.771	Mar.	20	4 .792	.461	2	4 .786	.459		
	9	5 .399	.762	4	5 .405	.768	Apr.	10	4 .794	.451	4	4 .794	.446		
	12	5 .422	.768	5	5 .419	+ 0 .767		11	4 .794	.453	4	4 .794	.447		
				mean	5 .401	+ 0 .768		15	4 .807	.448	4	4 .803	.444		
				40 Leo			May	4	4 .793	.450	3	4 .795	.449		
				V	B-V			9	4 .809	.438	3	4 .806	.446		
				System of primary stds.	Wt.	System of 10-year stds.		10	4 .814	.452	3	4 .808	+ 0 .447		
1955												mean	4 .796	+ 0 .449	
Feb.	24	4 <sup>m</sup> .801	+ 0 <sup>m</sup> .450	2	4 <sup>m</sup> .791	+ 0 <sup>m</sup> .447	1959								
Apr.	4	4 .793	.458	3	4 .797	.448	Jan.	5	4 .754	.446	2	4 .786	+ 0 .450		
	11	4 .794	.455	4				12	4 .782	.453	3	4 .787	.455		
	20	4 .796	.446	5	4 .805	.445		24	4 .784	.441	2	4 .788	.457		
	21	4 .800	.455	4	4 .801	.451		29	4 .776	.439	5				
May	4	4 .797	.441	4	4 .800	.444		31	4 .781	.441	4	4 .790	.446		
	6	4 .791	.454	4	4 .786	.458	Feb.	2	4 .785	.444	4				
	28	4 .796	.457	3	4 .795	+ 0 .455		4	4 .779	.445	5	4 .794	.450		
				mean	4 .797	+ 0 .450		5	4 .805	.430	5	4 .810	.431		
								26	4 .782	.458	5	4 .798	.450		

**Table IV. Observations of Ten Year Standards (Cont'd)**

40 Leo							36 U Ma (A)						
DATE	V	B-V	Wt.	V	B-V	Wt.	V	B-V	Wt.	V	B-V	Wt.	mean
	System of primary stds.	System of 10-year stds.		System of primary stds.	System of 10-year stds.		mean	System of primary stds.		mean	System of primary stds.		
1959							1955						
Feb. 28	4 <sup>m</sup> .797	+ 0 <sup>m</sup> .452	3	4 <sup>m</sup> .791	+ 0 <sup>m</sup> .453		May 28	4 <sup>m</sup> .841	+ 0 <sup>m</sup> .518	4	4 <sup>m</sup> .840	+ 0 <sup>m</sup> .517	
Mar. 1	4 .792	.445	2	4 .794	.455					mean	4 .840	+ 0 .515	
3	4 .775	.432	4	4 .798	+ 0 .446		Jan. 17	4 .823	.519	4	4 .831	+ 0 .516	
			mean	4 .795	+ 0 .448		Feb. 3	4 .833	.518	4	4 .831	.520	
1960							10	4 .847	.507	4	4 .838	.506	
Jan. 20	4 .812	.456	4				14	4 .833	.538	4	4 .823	.534	
28	4 .789	.447	3	4 .792	+ 0 .447		Mar. 5	4 .802	.509	2	4 .840	.514	
Feb. 17	4 .797	.443	2	4 .795	.448		19	4 .832	.527	3	4 .835	.519	
21	4 .807	.446	4	4 .798	.446		Apr. 5	4 .830	.521	3	4 .823	.523	
23	4 .792	.452	3	4 .793	.450		18	4 .843	.500	2	4 .839	+ 0 .509	
Mar. 15	4 .793	.450	3	4 .794	.457				mean	4 .832	+ 0 .518		
17	4 .804	.440	3	4 .798	.446		1957						
18	4 .812	.454	2	4 .797	.453		Mar. 23	4 .830	.518	4	4 .829	+ 0 .520	
19	4 .791	.440	2	4 .799	.441		24	4 .831	.526	4	4 .836	.520	
21	4 .793	.473	3	4 .799	.448		Apr. 5	4 .830	.522	3			
26	4 .798	.449	2	4 .791	.451		9	4 .836	.521	2			
30	4 .791	.449	3	4 .796	.451		15	4 .835	.513	3	4 .832	.517	
Apr. 3	4 .789	.448	4	4 .795	.452		20	4 .832	.512	2			
6	4 .806	.444	3	4 .792	+ 0 .450		26	4 .848	.522	1			
			mean	4 .795	+ 0 .449		May 1	4 .874	.529	2	4 .863	.525	
1961							3	4 .830	.530	4	4 .833	.529	
Feb. 6	4 .786	.456	2	4 .786	+ 0 .454		5	4 .839	.512	1	4 .828	.515	
8	4 .795	.457	3	4 .790	.448		30	4 .839	.513	3	4 .839	+ 0 .513	
10	4 .800	.464	2	4 .793	.458		Jun. 1	4 .837	.528	3			
15	4 .808	.440	3	4 .795	.440				mean	4 .836	+ 0 .521		
21	4 .794	.450	4				Dec. 27	4 .821	.519	4			
22	4 .793	.454	4	4 .803	.450		1958						
24	4 .782	.459	3	4 .790	.445		Jan. 8	4 .855	.516	1			
Mar. 8	4 .805	.450	4	4 .793	.448		9	4 .830	.528	3			
23	4 .800	.446	3	4 .796	.445		16	4 .827	.517	4	4 .830	+ 0 .516	
31	4 .802	.438	4	4 .796	.453		Feb. 11	4 .838	.535	2	4 .847	.529	
Apr. 4	4 .792	.458	4	4 .785	.453		27	4 .800	.525	4			
9	4 .783	.448	4	4 .789	.454		Mar. 20	4 .839	.514	2	4 .833	.512	
12	4 .807	.454	4	4 .804	+ 0 .453		Apr. 10	4 .823	.526	4	4 .823	.521	
			mean	4 .793	+ 0 .450		11	4 .838	.516	4	4 .838	.510	
							15	4 .837	.523	4	4 .833	.519	
							May 4	4 .826	.516	3	4 .828	.515	
							9	4 .845	.507	3	4 .842	.515	
							10	4 .858	.519	3	4 .852	+ 0 .514	
									mean	4 .835	+ 0 .516		
1955							1959						
Feb. 24	4 <sup>m</sup> .836	+ 0 <sup>m</sup> .522	2	4 <sup>m</sup> .826	+ 0 <sup>m</sup> .519		Jan. 5	4 .824	.502	4	4 .856	+ 0 .506	
Apr. 4	4 .831	.529	3	4 .835	.519		12	4 .833	.502	2	4 .838	.504	
11	4 .830	.528	4				24	4 .835	.503	2	4 .839	.519	
20	4 .852	.496	3	4 .861	.495		31	4 .830	.513	4	4 .839	.518	
21	4 .847	.521	4	4 .848	.517		Feb. 4	4 .837	.502	5	4 .852	.507	
May 4	4 .831	.516	4	4 .834	.519		5	4 .827	.528	5	4 .832	.529	
6	4 .838	.515	4	4 .833	.519								

**Table IV. Observations of Ten Year Standards (Cont'd)**

36 U Ma (A)								$\beta$ Vir								
DATE	V System of primary stds.	B-V		V System of 10-year stds.	B-V		DATE	V System of primary stds.	B-V		V System of 10-year stds.	B-V		V System of 10-year stds.	B-V	
		Wt.			Wt.				Wt.			Wt.			Wt.	
1959																
Feb.	26	4.823	+ 0.525	5	4.823	+ 0.517	1955	May	4	3.597	+ 0.547	3	3.600	+ 0.550		
	28	4.844	.521	3	4.838	.522			6	3.602	.550	3	3.597	.554		
Mar.	1	4.843	.508	2	4.845	.518			28	3.599	.563	3	3.598	+ 0.561		
	3	4.805	.494	3	4.828	+ 0.508	1956					mean	3.612	+ 0.551		
1960			mean	4.839	+ 0.515		Jan.	17	3.589	.554	3	3.597	+ 0.551			
Jan.	28	4.829	.520	3	4.832	+ 0.520	Feb.	3	3.593	.568	3	3.591	.570			
Feb.	17	4.836	.513	3	4.834	.518		10	3.612	.540	3	3.603	.539			
	21	4.837	.516	4	4.828	.516		14	3.599	.549	3	3.589	.545			
	23	4.835	.514	3	4.836	.512	Mar.	5	3.566	.545	2	3.604	.550			
Mar.	15	4.831	.518	3	4.832	.525		14	3.600	.558	3					
	17	4.843	.508	3	4.837	.514		19	3.615	.547	2	3.618	.539			
	18	4.859	.523	1	4.844	.522	Apr.	5	3.611	.542	3	3.604	.544			
	19	4.822	.519	1	4.830	.520		18	3.620	.534	1	3.616	+ 0.543			
	21	4.825	.545	3	4.831	.520					mean	3.601	+ 0.548			
	26	4.838	.508	3	4.831	.510	1957									
	30	4.828	.513	2	4.833	.515	Mar.	23	3.613	.542	3	3.612	+ 0.544			
Apr.	3	4.832	.516	4	4.838	.520		24	3.600	.553	4	3.605	.547			
	6	4.839	.512	4	4.831	+ 0.518	Apr.	5	3.608	.558	2					
			mean	4.833	+ 0.517			15	3.608	.545	3	3.605	.549			
1961								20	3.591	.554	0					
Feb.	6	4.837	.517	3	4.837	+ 0.515		26	3.612	.550	1					
	8	4.849	.538	3	4.844	.529	May	1	3.621	.558	3	3.610	.554			
	10	4.848	.533	3	4.841	.527		3	3.603	.551	3	3.606	.550			
	15	4.850	.523	3	4.837	.523		5	3.611	.543	1	3.600	.546			
	21	4.841	.517	4				30	3.613	.533	2	3.613	+ 0.533			
	22	4.832	.526	4	4.842	.522	Jun.	1	3.617	.545	2					
	24	4.829	.528	3	4.837	.514				mean	3.608	+ 0.547				
Mar.	1	4.846	.523	4	4.839	.520	1958									
	8	4.859	.522	4	4.847	.520	Dec.	27	3.602	.548	1					
	23	4.836	.512	3	4.832	.511	Jan.	8	3.600	.546	2					
	31	4.840	.509	3	4.834	.524		9	3.598	.550	3					
Apr.	4	4.892	.510	4	4.885	.505		16	3.601	.557	3	3.604	+ 0.556			
	9	4.824	.509	4	4.830	.515	Feb.	11	3.587	.563	4	3.596	.557			
	12	4.844	.515	4	4.841	+ 0.514	Apr.	10	3.599	.561	3	3.599	.556			
			mean	4.842	+ 0.518			11	3.654	.569	5	3.654	.563			
								15	3.615	.552	3	3.611	.548			
							May	4	3.600	.545	3	3.602	.544			
								9	3.618	.547	3	3.615	.555			
DATE		V System of primary stds.	B-V		V System of 10-year stds.	B-V		10	3.615	.558	3	3.609	+ 0.553			
1955										mean	3.614	+ 0.555				
Feb.	10	3.602	+ 0.548	4			1959									
	24	3.645	.555	4	3.635	+ 0.552	Jan.	5	3.571	.552	1	3.603	+ 0.556			
Apr.	4	3.602	.554	4	3.606	.544		12	3.605	.537	3	3.610	.539			
	11	3.612	.564	3				24	3.596	.532	3	3.600	.548			
	20	3.611	.554	4	3.620	.553		31	3.595	.545	4	3.604	.550			
	21	3.622	.549	3	3.623	.545	Feb.	4	3.576	.552	3	3.591	.557			
								5	3.597	.538	2	3.602	.539			

**Table IV. Observations of Ten Year Standards (Cont'd)**

$\beta$ Vir								$\beta$ CVn								
DATE	V	B-V	System of primary stds.	Wt.	V	B-V	System of 10-year stds.	DATE	V	B-V	System of primary stds.	Wt.	V	B-V	System of 10-year stds.	
1959								1955								
Feb.	26	3 <sup>m</sup> .577	+0 <sup>m</sup> .563	2	3 <sup>m</sup> .593	+0 <sup>m</sup> .555		May	28	4 <sup>m</sup> .261	+0 <sup>m</sup> .585	4	4 <sup>m</sup> .260	+0 <sup>m</sup> .583		
	28	3.617	.547	3	3.611	.548						mean	4.255	+0.585		
Mar.	1	3.606	.533	3	3.608	.543	1956	Jan.	17	4.241	.589	4	4.249	+0.586		
	3	3.582	.536	3	3.605	+0.550		Feb.	3	4.242	.595	4	4.240	.597		
				mean	3.603	+0.548			10	4.267	.583	4	4.258	.582		
1960									14	4.279	.596	4	4.269	.592		
Jan.	28	3.608	.536	3	3.611	+0.536		Mar.	5	4.212	.570	2	4.250	.575		
Feb.	17	3.621	.558	3	3.619	.563			19	4.253	.590	5	4.256	.582		
	21	3.622	.543	3	3.613	.543		Apr.	5	4.253	.600	2	4.246	.602		
Mar.	15	3.609	.556	3	3.610	.554			18	4.253	.571	2	4.249	+0.580		
	17	3.608	.538	3	3.609	.545						mean	4.253	+0.587		
	18	3.612	.544	3	3.606	.550		1957	Mar.	23	4.262	.590	5	4.261	+0.592	
	19	3.622	.537	2	3.607	.536			24	4.245	.598	4	4.250	.592		
	21	3.603	.548	2	3.611	.549		Apr.	5	4.252	.581	4				
	21	3.607	.568	3	3.613	.543			6	4.264	.582	4				
	26	3.612	.549	2	3.605	.551			15	4.264	.589	2	4.261	.593		
	30	3.603	.547	3	3.608	.549		May	1	4.279	.596	2	4.268	.592		
Apr.	3	3.620	.547	3	3.626	.551			3	4.254	.587	2	4.257	.586		
	6	3.620	.548	3	3.612	+0.554			5	4.272	.580	2	4.261	.583		
				mean	3.612	+0.548			30	4.264	.581	5	4.264	+0.581		
1961								1958	mean	4.260	+0.588					
Feb.	6	3.604	.550	4	3.604	+0.548		Jan.	8	4.246	.585	1				
	10	3.623	.556	0	3.616	.550			9	4.253	.591	3				
	15	3.618	.554	2	3.605	.554			16	4.256	.588	2	4.259	+0.587		
	21	3.604	.543	2				Feb.	27	4.245	.589	3				
	22	3.600	.558	1	3.610	.554		Mar.	20	4.256	.592	2	4.250	.590		
	24	3.592	.558	2	3.600	.544		Apr.	10	4.256	.594	4	4.256	.589		
Mar.	1	3.626	.555	3	3.619	.552			11	4.255	.585	4	4.255	.579		
	8	3.627	.551	3	3.615	.549			15	4.260	.597	5	4.256	.593		
	23	3.597	.552	3	3.593	.551		May	4	4.254	.586	4	4.256	.585		
Apr.	4	3.606	.548	4	3.599	.543			9	4.262	.578	3	4.259	.586		
	9	3.585	.543	3	3.591	.549			10	4.270	.590	4	4.264	+0.585		
	12	3.611	.561	3	3.608	+0.560						mean	4.257	+0.587		
				mean	3.609	+0.550										
$\beta$ CVn								1959								
DATE	V	B-V	System of primary stds.	Wt.	V	B-V	System of 10-year stds.	1959	Jan.	5	4.208	.562	2	4.240	.566	
1955									12	4.259	.573	3	4.264	.575		
Feb.	10	4 <sup>m</sup> .259	+0 <sup>m</sup> .587	4					24	4.252	.572	3	4.256	.588		
	24	4.269	.591	4	4 <sup>m</sup> .259	+0 <sup>m</sup> .588			31	4.248	.579	4	4.257	.584		
Apr.	4	4.238	.607	2	4.242	.597		Feb.	4	4.241	.581	5	4.256	.586		
	20	4.243	.584	5	4.252	.583			5	4.251	.574	4	4.256	.575		
	21	4.254	.593	4	4.255	.589			26	4.244	.599	4	4.260	.591		
May	4	4.245	.583	4	4.248	.586			28	4.261	.581	3	4.255	.582		
	6	4.265	.576	5	4.260	.580		Mar.	1	4.259	.575	2	4.261	.585		

**Table IV. Observations of Ten Year Standards (Cont'd)**

β CVn						78 UMa					
DATE	V System of primary stds.	B-V		V System of 10-year stds.	Wt.	DATE	V System of primary stds.	B-V		Wt.	V System of 10-year stds.
		Wt.	B-V					Wt.	B-V		
1959						1955					
Mar. 3	4.228	+ 0.578	4	4.251	+ 0.592	May 28	4.931	+ 0.360	4	4.930	+ 0.358
			mean	4.256	+ 0.583				mean	4.927	+ 0.364
1960						1956					
Jan. 28	4.255	.585	2	4.258	+ 0.585	Jan. 17	4.915	.370	4	4.923	+ 0.367
Feb. 17	4.262	.585	5	4.260	.590	Feb. 3	4.927	.361	4	4.925	.363
21	4.275	.580	4	4.266	.580	10	4.925	.368	4	4.916	.367
23	4.261	.584	5	4.262	.582	14	4.927	.368	4	4.917	.364
Mar. 15	4.253	.581	3	4.254	.588	Mar. 5	4.877	.363	2	4.915	.368
17	4.266	.580	3	4.260	.586	19	4.937	.359	4	4.940	.351
18	4.282	.586	1	4.267	.585	Apr. 5	4.933	.362	3	4.926	.364
19	4.252	.579	1	4.260	.580	18	4.929	.356	2	4.925	+ 0.365
21	4.245	.611	3	4.251	.586				mean	4.924	+ 0.363
26	4.263	.586	5	4.256	.588	1957					
30	4.249	.584	2	4.254	.586	Mar. 23	4.939	.356	4	4.938	+ 0.358
Apr. 3	4.253	.581	4	4.259	.585	24	4.910	.378	4	4.915	.372
6	4.262	.577	4	4.254	+ 0.583	Apr. 6	4.928	.358	3		
			mean	4.258	+ 0.585	15	4.934	.359	2	4.931	.363
1961						May 1	4.946	.371	2	4.935	.367
Feb. 6	4.261	.583	5	4.261	+ 0.581	3	4.926	.365	4	4.929	.364
8	4.275	.598	2	4.270	.589	5	4.948	.356	2	4.937	.359
10	4.267	.607	2	4.260	.601	30	4.920	.371	5	4.920	+ 0.371
15	4.266	.590	3	4.253	.590	Jun. 1	4.915	.377	4		
21	4.268	.589	3						mean	4.928	+ 0.366
22	4.261	.594	3	4.271	.590	1958					
24	4.257	.601	3	4.265	.587	Jan. 8	4.911	.357	2		
Mar. 1	4.269	.593	3	4.262	.590	9	4.915	.369	3		
8	4.279	.591	4	4.267	.589	16	4.916	.363	2	4.919	+ 0.362
23	4.278	.589	4	4.274	.588	Feb. 11	4.920	.377	3	4.929	.371
31	4.266	.576	3	4.260	.591	27	4.934	.369	2		
Apr. 4	4.252	.594	5	4.245	.589	Mar. 20	4.929	.374	2	4.923	.372
9	4.240	.582	5	4.246	.588	Apr. 10	4.924	.373	4	4.924	.368
12	4.255	.589	5	4.252	+ 0.588	11	4.920	.370	4	4.920	.364
			mean	4.259	+ 0.587	15	4.938	.371	4	4.934	.367
						May 4	4.928	.368	4	4.930	.367
						9	4.936	.358	3	4.933	.366
						10	4.927	.376	3	4.921	+ 0.371
									mean	4.926	+ 0.367
						1959					
						Jan. 5	4.902	.363	2	4.934	+ 0.367
1955						12	4.921	.363	3	4.926	.365
Feb. 10	4.942	+ 0.348	4			24	4.922	.347	3	4.926	.363
24	4.939	.371	4	4.929	+ 0.368	31	4.924	.351	4	4.933	.356
Apr. 4	4.925	.377	3	4.929	.367	Feb. 4	4.910	.372	4	4.925	.377
20	4.911	.366	5	4.920	.365	5	4.924	.355	3	4.929	.356
21	4.928	.361	4	4.929	.357	26	4.914	.373	3	4.930	.365
May 4	4.925	.363	4	4.928	.366	28	4.933	.362	3	4.927	.363
6	4.934	.360	4	4.929	.364	Mar. 1	4.924	.356	2	4.924	.366

Table IV. Observations of Ten Year Standards (Cont'd)

78 UMa							HD 115043						
DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V	Wt.	DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.		
1959							1956						
Mar. 3	4 <sup>m</sup> .899	+ 0 <sup>m</sup> .356	4	4 <sup>m</sup> .922	+ 0 <sup>m</sup> .370		Jan. 17	6 <sup>m</sup> .822	+ 0 <sup>m</sup> .604	4	6 <sup>m</sup> .830	+ 0 <sup>m</sup> .601	
			mean	4.927	+ 0.365		Feb. 3	6.860	.598	4	6.858	.600	
							10	6.849	.592	4	6.840	.591	
1960							14	6.856	.620	4	6.846	.616	
Jan. 28	4.919	.369	2	4.922	+ 0.369		Mar. 5	6.814	.575	2	6.852	.580	
Feb. 17	4.925	.365	5	4.923	.370		19	6.837	.612	4	6.840	.604	
21	4.931	.364	4	4.922	.364		Apr. 5	6.860	.598	3	6.853	.600	
23	4.925	.365	5	4.927	.363		18	6.822	.604	2	6.818	+ 0.613	
Mar. 15	4.942	.330	3	4.943	.337					mean	6.843	+ 0.601	
17	4.932	.367	3	4.926	.373								
18	4.945	.372	1	4.930	.371								
19	4.913	.361	1	4.921	.362		Mar. 23	6.816	.602	4	6.815	+ 0.604	
21	4.924	.386	3	4.930	.361		24	6.812	.609	4	6.817	.603	
26	4.940	.360	4	4.933	.362		Apr. 6	6.833	.610	3			
30	4.914	.367	2	4.919	.369		15	6.821	.603	2	6.818	.607	
Apr. 3	4.935	.366	4	4.941	.370		May 1	6.841	.610	2	6.830	.606	
6	4.932	.361	4	4.924	+ 0.367		3	6.815	.605	4	6.818	.604	
			mean	4.928	+ 0.365		5	6.830	.601	2	6.819	.604	
1961							30	6.824	.605	5	6.824	+ 0.605	
Feb. 6	4.935	.368	4	4.935	+ 0.366		Jun. 1	6.814	.605	4			
8	4.931	.387	2	4.926	.378				mean	6.820	+ 0.604		
10	4.941	.365	2	4.934	.359								
15	4.949	.377	2	4.936	.377								
21	4.964	.354	3				1958						
22	4.933	.378	3	4.943	.374		Jan. 8	6.819	.602	1			
24	4.929	.380	3	4.937	.366		9	6.812	.607	3			
Mar. 1	4.963	.380	3	4.956	.377		16	6.822	.611	2	6.825	+ 0.610	
8	4.957	.375	4	4.945	.373		Feb. 11	6.809	.615	3	6.818	.609	
23	4.945	.371	3	4.941	.370		27	6.820	.605	2			
31	4.942	.352	3	4.936	.367		Mar. 20	6.814	.609	1	6.808	.607	
Apr. 4	4.936	.370	4	4.929	.365		Apr. 10	6.830	.615	4	6.830	.610	
9	4.908	.353	4	4.914	.359		11	6.835	.603	4	6.835	.597	
12	4.931	.383	4	4.928	+ 0.382		15	6.836	.610	4	6.832	.606	
			mean	4.935	+ 0.370		May 4	6.832	.595	4	6.834	.594	
							9	6.833	.603	3	6.830	.611	
							10	6.831	.607	3	6.825	+ 0.602	
									mean	6.829	+ 0.604		
HD 115043							1959						
DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V	Wt.	Jan.	12	6.836	.593	3	6.841	+ 0.595
1955							24	6.820	.581	3	6.824	.597	
Feb. 10	6 <sup>m</sup> .845	+ 0 <sup>m</sup> .610	4				31	6.809	.591	4	6.818	.596	
24	6.836	.603	3	6 <sup>m</sup> .826	+ 0 <sup>m</sup> .600		Feb. 4	6.799	.602	3	6.814	.607	
Apr. 4	6.821	.624	3	6.825	.614		5	6.814	.608	3	6.819	.609	
20	6.819	.606	5	6.828	.605		26	6.804	.617	3	6.820	.609	
21	6.837	.594	4	6.838	.590		28	6.818	.607	3	6.812	.608	
May 4	6.845	.603	4	6.848	.606		Mar. 1	6.837	.592	2	6.839	.602	
6	6.842	.596	4	6.837	.600		3	6.800	.580	4	6.823	+ 0.594	
28	6.850	.606	4	6.849	+ 0.604				mean	6.823	+ 0.601		
			mean	6.836	+ 0.602								
							1960						
							Jan. 28	6.811	.611	2	6.814	+ 0.611	

**Table IV. Observations of Ten Year Standards (Cont'd)**

HD 115043							$\beta$ Com							
DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V	DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V			
<b>1960</b>														
Feb.	17	6 <sup>m</sup> .815	+ 0 <sup>m</sup> .599	5	6 <sup>m</sup> .813	+ 0 <sup>m</sup> .604	1956	Mar.	5	4 <sup>m</sup> .206	+ 0 <sup>m</sup> .572	2	4 <sup>m</sup> .244	+ 0 <sup>m</sup> .577
	21	6 .826	.614	4	6 .817	.614			14	4 .235	.590	4		
	23	6 .816	.603	5	6 .817	.601			19	4 .239	.586	5	4 .242	.578
Mar.	15	6 .811	.572	3	6 .812	.579		Apr.	5	4 .273	.564	3	4 .266	.566
	17	6 .815	.592	3	6 .809	.598			18	4 .254	.560	2	4 .250	+ 0 .569
	18	6 .830	.605	1	6 .815	.604						mean	4 .249	+ 0 .574
	19	6 .810	.602	1	6 .818	.603								
	21	6 .810	.627	3	6 .816	.602	1957	Mar.	23	4 .264	.561	4	4 .263	+ 0 .563
	26	6 .825	.605	4	6 .818	.607			24	4 .233	.574	3	4 .238	.568
	30	6 .813	.594	2	6 .818	.596		Apr.	6	4 .256	.572	3		
Apr.	3	6 .820	.608	4	6 .826	.612			15	4 .263	.576	2	4 .260	.580
	6	6 .828	.594	4	6 .820	+ 0 .600		May	1	4 .252	.566	2	4 .241	.562
					mean 6 .817	+ 0 .603			3	4 .246	.569	4	4 .249	.568
									5	4 .247	.564	2	4 .236	.567
<b>1961</b>														
Feb.	6	6 .822	.611	4	6 .822	+ 0 .609		30	4 .254	.573	5	4 .254	+ 0 .573	
	8	6 .827	.608	3	6 .822	.599		Jun.	1	4 .254	.575	4		
	10	6 .832	.601	2	6 .825	.595	1958				mean	4 .250	+ 0 .569	
	15	6 .843	.610	2	6 .830	.610		Jan.	8	4 .232	.566	2		
	21	6 .824	.609	2					9	4 .243	.582	3		
	22	6 .833	.606	2	6 .843	.602			16	4 .250	.574	4	4 .253	+ 0 .573
	24	6 .814	.631	3	6 .822	.617		Feb.	11	4 .244	.575	3	4 .253	.569
Mar.	1	6 .844	.616	3	6 .837	.613		Mar.	20	4 .276	.581	2	4 .270	.579
	8	6 .851	.614	4	6 .839	.612		Apr.	10	4 .249	.581	4	4 .249	.576
	23	6 .849	.610	3	6 .845	.609			11	4 .243	.582	4	4 .243	.576
	31	6 .826	.591	3	6 .820	.606			15	4 .255	.577	4	4 .251	.573
Apr.	4	6 .842	.612	4	6 .835	.607		May	4	4 .248	.572	4	4 .250	.571
	9	6 .821	.596	4	6 .827	.602			9	4 .253	.566	4	4 .250	.574
	12	6 .820	.610	4	6 .817	+ 0 .609			10	4 .250	.576	4	4 .244	+ 0 .571
					mean 6 .829	+ 0 .607					mean	4 .250	+ 0 .573	
$\beta$ Com														
DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V	DATE	Jan.							
<b>1955</b>														
Feb.	24	4 <sup>m</sup> .266	+ 0 <sup>m</sup> .568	3	4 <sup>m</sup> .256	+ 0 <sup>m</sup> .565		1959	5	4 .221	.570	2		
Apr.	4	4 .245	.588	2	4 .249	.578			12	4 .258	.567	3	4 .263	+ 0 .569
	20	4 .235	.566	5	4 .244	.565			24	4 .250	.555	3	4 .254	.571
	21	4 .243	.581	4	4 .244	.577			31	4 .250	.571	4	4 .259	.576
May	4	4 .247	.564	4	4 .250	.567		Feb.	4	4 .244	.560	3	4 .259	.565
	6	4 .261	.561	5	4 .256	.565			5	4 .243	.572	3	4 .248	.573
	28	4 .251	.567	4	4 .250	+ 0 .565			26	4 .240	.582	3	4 .256	.574
					mean 4 .250	+ 0 .568			28	4 .246	.572	3	4 .240	.573
							Mar.	1	4 .258	.556	2	4 .260	.566	
								3	4 .239	.554	4	4 .262	+ 0 .568	
											mean	4 .256	+ 0 .571	
<b>1956</b>														
Jan.	17	4 .241	.580	4	4 .249	+ 0 .577	1960	Jan.	28	4 .250	.570	2	4 .253	+ 0 .570
Feb.	3	4 .240	.591	4	4 .238	.593		Feb.	17	4 .261	.569	5	4 .259	.574
	10	4 .264	.567	4	4 .255	.567			21	4 .256	.579	4	4 .247	.579
	14	4 .262	.568	5	4 .252	.564			23	4 .249	.579	5	4 .250	.577
							Mar.	15	4 .259	.571	3	4 .260	.578	

**Table IV. Observations of Ten Year Standards (Cont'd)**

		β Com				61 Vir								
DATE		V System of primary stds.	B-V	V System of 10-year stds.	B-V	DATE	V System of primary stds.	B-V	V System of 10-year stds.					
1960						1956								
Mar.	17	4.263	+0.566	4	4.257	+0.572	Apr.	18	4.775	+0.691				
	18	4.273	.574	1	4.258	.573			0	4.771	+0.700			
	19	4.254	.557	1	4.262	.558			mean	4.734	+0.707			
	21	4.241	.601	3	4.247	.576	1957	Mar.	23	4.726	.715	1	4.725	+0.717
	26	4.257	.564	5	4.250	.566		24	4.724	.701	0	4.729	.695	
	30	4.251	.567	2	4.256	.569	Apr.	6	4.761	.706	4			
Apr.	3	4.244	.572	4	4.250	.576		15	4.730	.712	1	4.727	.716	
	6	4.250	.569	4	4.242	+0.575	May	1	4.714	.680	1	4.703	.676	
				mean	4.252	+0.574		3	4.748	.710	1	4.751	.709	
1961								5	4.774	.749	1	4.763	.752	
Feb.	6	4.256	.573	5	4.256	+0.571		30	4.741	.697	1	4.741	+0.697	
	8	4.249	.569	3	4.244	.560	Jun.	1	4.750	.706	1			
	10	4.250	.576	1	4.243	.570			mean	4.735	+0.711			
	15	4.278	.567	3	4.265	.567	1958	Jan.	8	4.761	.711	2		
	22	4.241	.578	4	4.251	.574		16	4.734	.707	2	4.737	+0.706	
	24	4.246	.588	4	4.254	.574	Mar.	20	4.739	.715	1	4.733	.713	
Mar.	1	4.251	.566	4	4.244	.563	Apr.	10	4.723	.725	1	4.723	.720	
	8	4.251	.569	4	4.239	.567		11	4.726	.699	2	4.726	.693	
	23	4.259	.578	3	4.255	.577		15	4.730	.706	2	4.726	.702	
	31	4.249	.559	4	4.243	.574	May	4	4.738	.701	2	4.740	.700	
Apr.	4	4.246	.579	5	4.239	.574		9	4.735	.699	2	4.732	.707	
	9	4.248	.568	3	4.254	.574		10	4.735	.720	2	4.729	+0.715	
	12	4.245	.577	5	4.242	+0.576								
				mean	4.248	+0.571	1959			mean	4.731	+0.706		
							Jan.	5	4.680	.718	0			
							12	4.711	.730	0	4.716	+0.732		
							24	4.707	.688	0	4.711	.704		
							31	4.726	.705	0	4.735	.710		
							Feb.	4	4.704	.675	0	4.719	.680	
								5	4.713	.693	0	4.718	.694	
								26	4.714	.718	0	4.730	.710	
								28	4.760	.699	1	4.754	.700	
							Mar.	1	4.724	.696	0	4.726	.706	
								3	4.715	.686	1	4.738	+0.700	
										mean	(4.746) (+0.700)			
1956						1960								
						Jan.	28	4.726	.705	1	4.729	+0.705		
Jan.	17	4.730	.703	2	4.738	+0.700	Feb.	17	4.722	.706	0	4.720	.711	
Feb.	3	4.742	.695	1	4.740	.697		21	4.739	.706	1	4.730	.706	
	10	4.742	.702	2	4.733	.701		23	4.744	.702	1	4.745	.700	
	14	4.734	.729	2	4.724	.725	Mar.	15	4.723	.716	2	4.724	.723	
Mar.	5	4.692	.708	1	4.730	.713		17	4.682	.687	0	4.676	.693	
	19	4.727	.712	2	4.730	.704		18	4.729	.706	2	4.714	.705	
Apr.	5	4.749	.702	2	4.742	.704		19	4.728	.722	2	4.736	.723	
								21	4.733	.740	2	4.739	.715	
								26	4.756	.707	0	4.749	.709	
								30	4.741	.704	1	4.746	.706	

**Table IV. Observations of Ten Year Standards (Cont'd)**

61 Vir							70 Vir								
DATE	V	B-V	System of primary stds.	Wt.	V	B-V	System of 10-year stds.	DATE	V	B-V	System of primary stds.	Wt.	V	B-V	System of 10-year stds.
1960								1957							
Apr.	3	4 <sup>m</sup> .757	+ 0 <sup>m</sup> .728	0	4 <sup>m</sup> .763	+ 0 <sup>m</sup> .732		May	3	4 <sup>m</sup> .983	+ 0 <sup>m</sup> .710	4	4 <sup>m</sup> .986	+ 0 <sup>m</sup> .709	
	6	4.773	.707	1	4.765	+ 0.713		5	4.991	.714	2	4.980	.717		
			mean 4.734			+ 0.712		30	4.979	.710	5	4.979	+ 0.710		
1961								Jun.	1	4.968	.727	4			
Feb.	6	4.732	.713	2	4.732	+ 0.711					mean 4.980		+ 0.708		
	8	4.772	.722	1	4.767	.713		1958							
	15	4.731	.714	1	4.718	.714		Jan.	16	4.968	.717	3	4.971	+ 0.716	
	22	4.715	.702	1	4.725	.698		Feb.	11	4.961	.723	3	4.970	.717	
Mar.	1	4.755	.725	2	4.748	.722		Mar.	20	5.013	.715	1	5.007	.713	
	8	4.747	.705	1	4.735	.703		Apr.	10	4.972	.725	3	4.972	.720	
	23	4.752	.715	2	4.748	.714		11	4.972	.724	4	4.972	.718		
	31	4.751	.694	2	4.745	.709		15	4.988	.725	4	4.984	.721		
Apr.	4	4.719	.707	2	4.712	.702		May	4	4.968	.724	4	4.970	.723	
	12	4.746	.723	2	4.743	+ 0.722		9	4.976	.696	3	4.973	.704		
			mean 4.738			+ 0.712		10	4.973	.717	3	4.967	+ 0.712		
	70 Vir							1959							
DATE	V	B-V	System of primary stds.	Wt.	V	B-V	System of 10-year stds.	Jan.							
1955								5	4.959	.722	1				
Feb.	8	4 <sup>m</sup> .984	+ 0 <sup>m</sup> .707	4				12	4.964	.715	2	4.969	+ 0.717		
	10	4.990	.707	4				24	4.969	.699	3	4.973	.715		
	24	4.998	.717	3	4 <sup>m</sup> .988	+ 0 <sup>m</sup> .714		31	4.967	.706	4	4.976	.711		
Apr.	4	4.953	.738	3	4.957	.728		Feb.	4	4.954	.702	2	4.969	.707	
	20	4.963	.716	5	4.972	.715		5	4.956	.716	2	4.961	.717		
	21	4.959	.729	4	4.960	.725		26	4.958	.726	2	4.974	.718		
May	4	4.977	.706	4	4.980	.709		28	4.980	.704	3	4.974	.705		
	6	4.995	.710	4	4.990	.714		Mar.	1	4.968	.707	3	4.970	.717	
	28	4.975	.711	4	4.974	+ 0.709		3	4.944	.696	4	4.967	+ 0.710		
			mean 4.974			+ 0.716					mean 4.971		+ 0.713		
1956								1960							
Jan.	17	4.959	.722	3	4.967	+ 0.719		Jan.	28	4.971	.713	2	4.974	+ 0.713	
Feb.	3	4.966	.718	4	4.964	.720		Feb.	17	4.975	.707	5	4.973	.712	
	10	4.986	.714	4	4.977	.713		21	4.990	.714	4	4.981	.714		
	14	4.990	.703	4	4.980	.699		23	4.981	.713	5	4.982	.711		
Mar.	5	4.928	.725	2	4.966	.730		Mar.	15	4.976	.700	3	4.977	.707	
	14	4.983	.707	4				17	4.984	.708	3	4.978	.714		
	19	4.975	.708	4	4.978	.700		18	4.991	.718	1	4.976	.717		
Apr.	5	4.994	.719	3	4.987	.721		19	4.980	.705	1	4.988	.706		
	18	4.982	.691	2	4.978	+ 0.700		21	4.983	.734	3	4.989	.709		
			mean 4.975			+ 0.712		26	4.995	.714	4	4.988	.716		
1957								30	4.979	.714	2	4.984	.716		
Mar.	23	4.991	.709	4	4.990	+ 0.711		Apr.	3	4.977	.704	4	4.983	.708	
	24	4.965	.710	3	4.970	.704		6	4.986	.709	4	4.978	+ 0.715		
Apr.	6	4.983	.712	3											
	15	4.987	.706	3	4.984	.710		1961							
May	1	4.978	.702	3	4.967	.698					mean 4.981		+ 0.712		

**Table IV. Observations of Ten Year Standards (Cont'd)**

70 Vir								70 Vir							
DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V	Wt.	DATE 1961	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V	Wt.		
1961															
Feb. 24	4 <sup>m</sup> .964	+ 0 <sup>m</sup> .731	3	4 <sup>m</sup> .972	+ 0 <sup>m</sup> .717		Apr. 9	4 <sup>m</sup> .975	+ 0 <sup>m</sup> .723	4	4 <sup>m</sup> .968	+ 0 <sup>m</sup> .718			
Mar. 1	4 .977	.714	4	4 .970	.711		9	4 .973	.705	3	4 .979	.711			
8	4 .981	.717	4	4 .969	.715		12	4 .972	.717	4	4 .969	+ 0 .716			
23	4 .978	.712	3	4 .974	.711								mean	4 .971 + 0 .711	
31	4 .977	.699	3	4 .971	.704										
$\tau$ Boo															
DATE								Wt.							
1955								V System of 10-year stds.							
Feb								B-V							
8	4 <sup>m</sup> .498	+ 0 <sup>m</sup> .487						4							
24	4 .503	.488						3	4 <sup>m</sup> .493	+ 0 <sup>m</sup> .485					
Apr.								3	4 .492	.470					
4	4 .488	.480						5	4 .491	.478					
20	4 .482	.479						3	4 .476	.490					
21	4 .475	.494						4	4 .494	.479					
May								4	4 .507	.472					
4	4 .491	.476						4	4 .482	+ 0 .477					
6	4 .512	.468													
28	4 .483	.479											mean	4 .491 + 0 .478	
1956															
Jan.								V System of primary stds.							
17	4 .493	.470						3	4 .501	+ 0 .467					
Feb.								4	4 .523	.453					
3	4 .525	.451						4	4 .511	.507					
10	4 .520	.508						4	4 .495	.490					
14	4 .505	.494													
Mar.								2	4 .488	.470					
5	4 .450	.465						4	4 .482	+ 0 .477					
14	4 .482	.499											mean	4 .491 + 0 .478	
19	4 .486	.502													
Apr.								3	4 .491	.493					
5	4 .498	.491						2	4 .491	+ 0 .484					
18	4 .495	.475											mean	4 .500 + 0 .483	
1957															
Mar.								V System of primary stds.							
23	4 .500	.479						4	4 .499	+ 0 .481					
24	4 .482	.485						4	4 .487	.479					
Apr.								3	4 .484	.485					
6	4 .494	.484						3	4 .498	.485					
15	4 .501	.481													
24	4 .516	.470						1							
May								2	4 .468	.470					
1	4 .479	.474						4	4 .496	.484					
3	4 .493	.485						1	4 .513	.466					
5	4 .524	.463						4	4 .489	+ 0 .483					
30	4 .489	.483													
Jun.								1	4 .487	.486					
								4	4 .492	+ 0 .480					
1958															
Jan.								V System of primary stds.							
8	4 .484	.483						2							
16	4 .488	.485						3	4 .491	+ 0 .484					
Feb.								4	4 .488	.475					
11	4 .479	.481													
Mar.								1	4 .515	.495					
20	4 .521	.497						3	4 .495	.484					
Apr.								11	4 .499	.479					
								4	4 .499	.473					

Table IV. Observations of Ten Year Standards (Cont'd)

 $\tau$  Boo

DATE		V	B-V	Wt.	V	B-V
		System of primary stds.			System of 10-year stds.	
<b>1958</b>						
Apr.	15	4 <sup>m</sup> .498	+ 0 <sup>m</sup> .486	4	4 <sup>m</sup> .494	+ 0 <sup>m</sup> .482
May	4	4 .489	.486	4	4 .491	.485
	9	4 .489	.470	3	4 .486	.478
	10	4 .484	.486	3	4 .478	+ 0 .481
				mean 4 .491		+ 0 .481
<b>1959</b>						
Jan.	5	4 .479	.481	1	4 .511	+ 0 .485
	12	4 .489	.476	2	4 .494	.478
	24	4 .495	.461	3	4 .499	.477
	31	4 .484	.475	4	4 .493	.480
Feb.	4	4 .480	.472	2	4 .495	.477
	5	4 .491	.473	2	4 .496	.474
	26	4 .480	.476	2	4 .496	.468
	28	4 .486	.476	3	4 .480	.477
Mar.	1	4 .484	.470	3	4 .486	.480
	3	4 .471	.452	4	4 .494	+ 0 .466
				mean 4 .493		+ 0 .476
<b>1960</b>						
Jan.	28	4 .497	.477	2	4 .500	+ 0 .477
Feb.	17	4 .528	.439	5	4 .526	.444
	21	4 .510	.474	4	4 .501	.474
	23	4 .486	.490	5	4 .487	.488
Mar.	15	4 .493	.477	3	4 .494	.484
	17	4 .492	.481	3	4 .486	.487
	18	4 .507	.485	1	4 .492	.484
	19	4 .492	.471	1	4 .500	.472
	21	4 .492	.504	3	4 .498	.479
	26	4 .502	.480	5	4 .495	.482
	30	4 .493	.478	2	4 .498	.480
Apr.	3	4 .490	.467	4	4 .496	.471
	6	4 .503	.465	4	4 .495	+ 0 .471
				mean 4 .498		+ 0 .475
<b>1961</b>						
Feb.	6	4 .493	.485	4	4 .493	+ 0 .483
	8	4 .494	.511	3	4 .489	.502
	10	4 .494	.475	2	4 .487	.469
	15	4 .517	.480	3	4 .504	.480
	22	4 .471	.484	4	4 .481	.480
	24	4 .487	.502	2	4 .495	.488
Mar.	1	4 .498	.484	4	4 .491	.481
	8	4 .507	.472	4	4 .495	.470
	23	4 .533	.479	3	4 .529	.478
	31	4 .508	.460	3	4 .502	.475
Apr.	4	4 .500	.483	4	4 .493	.478
	9	4 .494	.469	2	4 .500	.475
	12	4 .499	.474	3	4 .496	+ 0 .473
				mean 4 .496		+ 0 .479

**Table IV. Observations of Ten Year Standards (Cont'd)**

DATE		$\eta$ Boo			Wt.	V	B-V
		V	B-V	System of primary stds.			
<b>1955</b>							
Feb.	8	2.681	+ 0.588		4		
	24	2.693	.577		3	2.683	+ 0.574
Apr.	4	2.663	.601		3	2.667	.591
	20	2.666	.587		5	2.675	.586
	21	2.677	.582		4	2.678	.578
May	4	2.679	.577		4	2.682	.580
	6	2.697	.571		4	2.692	.575
	28	2.667	.584		4	2.666	+ 0.582
					mean	2.678	+ 0.581
<b>1956</b>							
Jan.	17	2.686	.578		3	2.694	+ 0.575
Feb.	3	2.695	.573		4	2.693	.575
	10	2.689	.585		4	2.680	.584
	14	2.684	.592		4	2.674	.588
Mar.	5	2.643	.584		2	2.681	.589
	19	2.654	.605		4	2.657	.597
Apr.	5	2.698	.581		3	2.691	.583
	18	2.685	.569		2	2.681	+ 0.578
					mean	2.681	+ 0.584
<b>1957</b>							
Mar.	23	2.697	.583		4	2.696	+ 0.585
	24	2.684	.594		4	2.689	.588
Apr.	6	2.691	.583		3		
	15	2.702	.576		3	2.699	.580
	24	2.768	.588		1		
May	1	2.662	.581		1	2.651	.577
	3	2.684	.584		4	2.687	.583
	30	2.686	.581		5	2.686	+ 0.581
Jun.	1	2.677	.594		4		
					mean	2.689	+ 0.583
<b>1958</b>							
Jan.	8	2.682	.585		2		
	16	2.677	.590		3	2.680	.589
Feb.	11	2.666	.590		4	2.675	.584
Mar.	20	2.705	.581		1	2.699	.579
Apr.	10	2.679	.596		4	2.679	.591
	11	2.691	.575		3	2.691	.569
	15	2.693	.583		4	2.689	.579
May	4	2.677	.591		4	2.679	.590
	9	2.678	.582		3	2.675	.590
	10	2.678	.595		3	2.672	+ 0.590
					mean	2.681	+ 0.585
<b>1959</b>							
Jan.	5	2.675	.580		1	2.707	+ 0.584
	12	2.671	.586		2	2.676	.588
	24	2.676	.570		3	2.680	.586
	31	2.670	.578		4	2.679	.583

**Table IV. Observations of Ten Year Standards (Cont'd)**

DATE		$\eta$ Boo			Wt.	V	B-V
		V	B-V	System of primary stds.			
<b>1959</b>							
Feb.	4	2 <sup>m</sup> .660	+ 0 <sup>m</sup> .578		2	2 <sup>m</sup> .675	+ 0 <sup>m</sup> .583
	5	2 .684	.574		2	2 .689	.575
	26	2 .644	.574		1	2 .660	.566
	28	2 .670	.591		3	2 .664	.592
Mar.	1	2 .687	.568		3	2 .689	.578
	3	2 .641	.576		4	2 .664	+ 0 .590
					mean	2 .677	+ 0 .584
<b>1960</b>							
Jan.	28	2 .676	.586		2	2 .679	+ 0 .586
Feb.	17	2 .677	.579		5	2 .675	.584
	21	2 .695	.582		4	2 .686	.582
	23	2 .683	.584		5	2 .684	.582
Mar.	15	2 .683	.586		3	2 .684	.593
	17	2 .691	.574		3	2 .685	.580
	18	2 .702	.587		1	2 .687	.586
	19	2 .681	.571		1	2 .689	.572
	21	2 .678	.607		3	2 .684	.582
	26	2 .689	.582		5	2 .682	.584
	30	2 .681	.581		2	2 .686	.583
Apr.	3	2 .666	.584		4	2 .672	+ 0 .588
					mean	2 .682	+ 0 .584
<b>1961</b>							
Feb.	6	2 .688	.584		4	2 .688	+ 0 .582
	8	2 .683	.588		3	2 .678	.579
	10	2 .703	.579		2	2 .696	.573
	15	2 .689	.584		3	2 .676	.584
	22	2 .660	.584		4	2 .670	.580
	24	2 .683	.608		3	2 .691	.594
Mar.	1	2 .683	.584		4	2 .676	.581
	8	2 .684	.586		4	2 .672	.584
	23	2 .719	.588		3	2 .715	.587
	31	2 .693	.556		3	2 .687	.571
Apr.	4	2 .687	.582		4	2 .680	.577
	9	2 .674	.582		3	2 .680	.588
	12	2 .686	.585		3	2 .683	+ 0 .584
					mean	2 .683	+ 0 .582
<b><math>\sigma</math> Boo</b>							
DATE		System of primary stds.			Wt.	V	B-V
		V	B-V	System of 10-year stds.			
<b>1955</b>							
Feb.	8	4 <sup>m</sup> .474	+ 0 <sup>m</sup> .373		4		
	24	4 .475	.360		2	4 <sup>m</sup> .465	+ 0 <sup>m</sup> .357
Apr.	4	4 .445	.384		2	4 .449	.374
	20	4 .442	.368		5	4 .451	.367
	21	4 .449	.371		4	4 .450	.367
May	28	4 .470	.357		4	4 .469	+ 0 .355
					mean	4 .456	+ 0 .364
<b>1956</b>							
Jan.	17	4 .466	.368		4	4 .474	+ 0 .365

**Table IV. Observations of Ten Year Standards (Cont'd)**

DATE		$\sigma$ Boo			Wt.	V	B-V
		V	B-V	System of primary stds.			System of 10-year stds.
<b>1956</b>							
Feb.	3	4.482	+ 0.353		4	4.480	+ 0.355
	10	4.484	.364		5	4.475	.363
	14	4.487	.371		5	4.477	.367
Mar.	5	4.424	.357		2	4.462	.362
	19	4.470	.369		3	4.473	.361
Apr.	5	4.463	.366		3	4.456	+ 0.368
						mean 4.473	+ 0.363
<b>1957</b>							
Mar.	24	4.450	.370		4	4.455	+ 0.364
Apr.	5	4.460	.356		4		
	15	4.476	.357		2	4.473	.361
	24	4.436	.365		2		
May	1	4.447	.357		2	4.436	.353
	3	4.447	.366		4	4.450	.365
	30	4.458	.364		5	4.458	+ 0.364
Jun.	1	4.456	.368		4		
						mean 4.455	+ 0.363
<b>1958</b>							
Jan.	9	4.456	.363		3		
	16	4.458	.369		3	4.461	+ 0.368
Feb.	11	4.462	.360		4	4.471	.354
Mar.	20	4.495	.371		1	4.489	.369
Apr.	10	4.461	.372		3	4.461	.367
	11	4.451	.372		4	4.451	.366
	15	4.454	.369		4	4.450	.365
May	4	4.459	.366		5	4.461	.365
	9	4.452	.357		3	4.449	.365
	10	4.447	.370		4	4.441	+ 0.365
						mean 4.457	+ 0.364
<b>1959</b>							
Jan.	5	4.441	.357		1	4.473	+ 0.361
	12	4.450	.363		2	4.455	.365
	24	4.468	.342		3	4.472	.358
	31	4.465	.347		4	4.474	.352
Feb.	4	4.449	.356		2	4.464	.361
	5	4.453	.370		2	4.458	.371
	26	4.439	.379		2	4.455	.371
	28	4.459	.352		3	4.453	.353
Mar.	1	4.457	.352		3	4.459	.362
	3	4.423	.346		4	4.446	+ 0.360
						mean 4.460	+ 0.360
<b>1960</b>							
Jan.	28	4.467	.365		2	4.470	+ 0.365
	17	4.458	.365		5	4.456	.370
	21	4.476	.353		4	4.467	.353
	23	4.460	.363		5	4.461	.361
Mar.	15	4.467	.352		4	4.468	.359
	17	4.472	.356		4	4.466	.362

**Table IV. Observations of Ten Year Standards (Cont'd)**

DATE		$\sigma$ Boo			Wt.	V	B-V	
		V	B-V	System of primary stds.				
		System of 10-year stds.						
<b>1960</b>								
Mar.	18	4.474	+ 0.376		1	4.459	+ 0.375	
	19	4.471	.353		1	4.479	.354	
	21	4.460	.388		4	4.466	.363	
	26	4.475	.362		5	4.468	.364	
	30	4.465	.360		2	4.470	.362	
Apr.	3	4.451	.358		4	4.457	.362	
	6	4.470	.351		4	4.462	+ 0.357	
					mean	4.464	+ 0.362	
<b>1961</b>								
Feb.	8	4.466	.366		3	4.461	+ 0.357	
	10	4.462	.369		2	4.455	.363	
	15	4.472	.369		3	4.459	.369	
	22	4.455	.375		4	4.465	.371	
	24	4.465	.376		2	4.473	.362	
Mar.	1	4.479	.362		4	4.472	.359	
	8	4.476	.365		4	4.464	.363	
	23	4.466	.348		4	4.462	.347	
	31	4.483	.336		3	4.477	.351	
Apr.	4	4.468	.365		4	4.461	.360	
	9	4.464	.356		3	4.470	.362	
	12	4.466	.356		5	4.463	+ 0.355	
					mean	4.465	+ 0.360	
<b>1955</b>								
Apr.	4	4.531	+ 0.781		3	4.535	+ 0.771	
	20	4.544	.780		5	4.553	.779	
	21	4.530	.766		4	4.531	.762	
May	28	4.552	.764		4	4.551	+ 0.762	
					mean	4.544	+ 0.769	
<b>1956</b>								
Jan.	17	4.557	.765		4	4.565	+ 0.762	
Feb.	3	4.548	.752		4	4.546	.754	
	10	4.546	.771		4	4.537	.770	
	14	4.551	.778		4	4.541	.774	
Mar.	5	4.501	.773		2	4.539	.778	
	19	4.553	.777		3	4.556	.769	
Apr.	5	4.571	.758		3	4.564	+ 0.760	
					mean	4.550	+ 0.766	
<b>1957</b>								
Mar.	24	4.542	.768		4	4.547	+ 0.762	
Apr.	5	4.548	.763		4			
	6	4.563	.782		3			
	15	4.558	.772		2	4.555	.776	
	24	4.538	.773		2			
May	1	4.528	.753		2	4.517	.749	
	3	4.535	.765		4	4.538	.764	
	30	4.548	.771		5	4.548	+ 0.771	
					mean			
<b><math>\zeta</math> Boo</b>								
DATE		V	B-V	Wt.	$\zeta$ Boo		Corrections to system of 10-year stds.	
					V	B-V		
<b>1955</b>								
Apr.	4	4.531	+ 0.781		3	4.535	+ 0.771	
	20	4.544	.780		5	4.553	.779	
	21	4.530	.766		4	4.531	.762	
May	28	4.552	.764		4	4.551	+ 0.762	
					mean	4.544	+ 0.769	
<b>1956</b>								
Jan.	17	4.557	.765		4	4.565	+ 0.762	
Feb.	3	4.548	.752		4	4.546	.754	
	10	4.546	.771		4	4.537	.770	
	14	4.551	.778		4	4.541	.774	
Mar.	5	4.501	.773		2	4.539	.778	
	19	4.553	.777		3	4.556	.769	
Apr.	5	4.571	.758		3	4.564	+ 0.760	
					mean	4.550	+ 0.766	
<b>1957</b>								
Mar.	24	4.542	.768		4	4.547	+ 0.762	
Apr.	5	4.548	.763		4			
	6	4.563	.782		3			
	15	4.558	.772		2	4.555	.776	
	24	4.538	.773		2			
May	1	4.528	.753		2	4.517	.749	
	3	4.535	.765		4	4.538	.764	
	30	4.548	.771		5	4.548	+ 0.771	
					mean			

Table IV. Observations of Ten Year Standards (Cont'd)

DATE		V System of primary stds.	B-V	Wt.	$\xi$ Boo		Corrections to system of 10-year stds.
					V System of 10-year stds.	B-V	
1957							
Jun.	1	4 <sup>m</sup> .548	+0 <sup>m</sup> .766	4			
					mean 4 .543	+0 .765	
1958							
Jan.	16	4 .552	.763	3	4 .555	+0 .762	+0 <sup>m</sup> .003 -0 <sup>m</sup> .001
Mar.	20	4 .542	.756	5	4 .536	.754	-0 .006 -0 .002
Apr.	10	4 .548	.778	3	4 .548	.773	0 .000 -0 .005
	11	4 .547	.769	3	4 .547	.763	0 .000 -0 .006
	15	4 .541	.772	4	4 .537	.768	-0 .004 -0 .004
May	4	4 .546	.766	4	4 .548	.765	+0 .002 -0 .001
	9	4 .550	.755	3	4 .547	.763	-0 .003 +0 .008
	10	4 .539	.760	3	4 .533	+0 .755	-0 .006 -0 .005
					mean 4 .543	+0 .763	
1959							
Jan.	5	4 .541	.753	0	4 .573	+0 .757	+0 .032 +0 .004
	12	4 .524	.774	1	4 .529	.776	+0 .005 +0 .002
	24	4 .554	.746	2	4 .558	.762	+0 .004 +0 .016
	31	4 .541	.766	3	4 .550	.771	+0 .009 +0 .005
Feb.	4	4 .533	.752	1	4 .548	.757	+0 .015 +0 .005
	5	4 .551	.766	1	4 .556	.767	+0 .005 +0 .001
	26	4 .535	.768	1	4 .551	.760	+0 .016 -0 .008
	28	4 .546	.758	3	4 .540	.759	-0 .006 +0 .001
Mar.	1	4 .531	.755	2	4 .533	.765	+0 .002 +0 .010
	3	4 .507	.749	4	4 .530	+0 .763	+0 .023 +0 .014
					mean 4 .542	+0 .764	
1960							
Jan.	28	4 .551	.767	2	4 .554	+0 .767	+0 .003 0 .000
Feb.	17	4 .541	.772	5	4 .539	.777	-0 .002 +0 .005
	21	4 .556	.765	4	4 .547	.765	-0 .009 0 .000
	23	4 .550	.764	5	4 .551	.762	+0 .001 -0 .002
Mar.	15	4 .558	.757	3	4 .559	.764	+0 .001 +0 .007
	17	4 .556	.754	3	4 .550	.760	-0 .006 +0 .006
	18	4 .561	.764	1	4 .546	.763	-0 .015 -0 .001
	19	4 .532	.758	1	4 .540	.759	+0 .008 +0 .001
	21	4 .536	.796	3	4 .542	.771	+0 .006 -0 .025
	26	4 .550	.761	5	4 .543	.763	-0 .007 +0 .002
	30	4 .543	.765	2	4 .548	.767	+0 .005 +0 .002
Apr.	3	4 .536	.761	3	4 .542	+0 .765	+0 .006 +0 .004
					mean 4 .547	+0 .766	
1961							
Feb.	8	4 .565	.773	2	4 .560	+0 .764	-0 .005 -0 .009
	10	4 .576	.779	1	4 .569	.773	-0 .007 -0 .006
	15	4 .563	.768	3	4 .550	.768	-0 .013 0 .000
	22	4 .541	.774	4	4 .551	.770	+0 .010 -0 .004
	24	4 .525	.785	2	4 .533	.771	+0 .008 -0 .014
Mar.	1	4 .561	.775	4	4 .554	.772	-0 .007 -0 .003
	8	4 .570	.773	4	4 .558	.771	-0 .012 -0 .002
	23	4 .538	.778	3	4 .534	.777	-0 .004 -0 .001
	31	4 .555	.742	1	4 .549	.757	-0 .006 +0 .015

**Table IV. Observations of Ten Year Standards (Cont'd)**

DATE	V System of primary stds.	B-V	Wt.	$\xi$ Boo		Corrections to system of 10-year stds.	
				V System of 10-year stds.	B-V		
1961	Apr. 4	4.584	+0.778	4	4.577	+0.773	-0.007
	9	4.557	.754	3	4.563	.760	+0.006
	12	4.550	+0.748	4	4.547	+0.747	-0.003
				mean 4.554		+0.767	

**Table V. Observations of Comparison Stars**

UA 53 = HD 55052 = 48 Gem						NB 54 = HD 118705					
DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V	DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V
1954						1954					
Feb. 28	5.882	+0.364	3			May 31	9.171	+0.476	2		
Mar. 1	5.853	.372	3			June 2	9.181	.476	2		
1961						12	9.170	.492	1		
Feb. 6	5.864	.370	4	5.864	+0.368	14	9.166	.499	2		
8	5.856	.373	3	5.851	.364	16	9.134	.494	2		
10	5.867	.380	3	5.860	.374	17	9.180	.469	1		
21	5.885	.375	3			1961					
22	5.852	.369	4	5.862	.365	Feb. 8	9.172	.475	2	9.167	+0.466
24	5.856	.375	3	5.864	.361	10	9.171	.468	1	9.164	.462
Mar. 1	5.858	.374	4	5.851	.371	15	9.176	.475	2	9.163	.475
8	5.861	.377	4	5.849	.375	22	9.147	.462	2	9.157	.458
23	5.846	.365	3	5.842	.364	Mar. 1	9.190	.458	3	9.183	.455
31	5.880	.366	3	5.874	.381	8	9.178	.470	2	9.166	.468
Apr. 4	5.861	.379	2	5.854	.374	23	9.196	.445	2	9.192	.444
12	5.873	+0.374	4	5.870	+0.373	Apr. 4	9.183	.462	3	9.176	.457
				mean 5.859	+0.370	12	9.169	+0.470	3	9.166	+0.469
										mean 9.171	+0.461
NA 54 = HD 118704						UA 54 = HD 61997					
DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V	DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V
1954						1953					
May 31	8.501	+0.543	2			Oct. 30	7.105	+0.422	3		
Jun. 2	8.507	.564	2			Nov. 6	7.100	.416	3		
12	8.504	.552	1			8	7.102	.431	2		
14	8.491	.568	2			23	7.113	.415	2		
16	8.454	.562	2			1954					
17	8.500	.539	1			Feb. 17	7.120	.421	3		
1961						23	7.136	.422	3		
Feb. 8	8.515	.546	2	8.510	+0.537	24	7.144	.434	3		
10	8.495	.545	2	8.488	.539	28	7.146	.413	4		
15	8.520	.548	2	8.507	.548	Mar. 29	7.119	.421	3		
22	8.475	.531	2	8.485	.527	1961					
Mar. 1	8.511	.547	3	8.504	.544	Feb. 6	7.133	.421	4	7.133	+0.419
8	8.501	.540	2	8.489	.538	10	7.143	.415	3	7.136	.409
23	8.506	.553	2	8.502	.552	15	7.143	.412	3	7.130	.412
Apr. 4	8.489	.542	3	8.482	.537	21	7.132	.411	3		
12	8.489	+0.544	3	8.486	+0.543	22	7.129	.409	4	7.139	.405
				mean 8.494	+0.541	24	7.108	.423	3	7.116	.409

**Table V. Observations of Comparison Stars (Cont'd)**

UA 54 = HD 61997							UA' 54 = HD 58551						
DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V	DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V		
1961													
Mar. 1	7.129	+0.406	4	7.122	+0.403	1961	24	6.526	+0.472	3	6.534	+0.458	
8	7.139	.412	4	7.127	.410	Mar. 1	6.554	.458	4	6.547	.455		
23	7.122	.410	3	7.118	.409	8	6.551	.458	4	6.539	.456		
31	7.139	.414	3	7.133	.429	23	6.530	.459	3	6.526	.458		
Apr. 4	7.137	.418	3	7.130	.413	31	6.562	.458	2	6.556	.473		
12	7.136	+0.410	3	7.133	+0.409	Apr. 4	6.562	.455	2	6.555	.450		
				mean 7.129	+0.411	12	6.562	+0.464	3	6.559	+0.463		
										mean 6.545	+0.460		
UB 54 = HD 60914													
DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V	DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V		
1953													
Oct. 30	6.940	+1.047	2			1954	Mar. 30	7.143	+0.931	2			
Nov. 6	6.938	1.047	2			Apr. 1	7.147	.927	3				
1954							6	7.143	.934	2			
Feb. 17	6.963	1.049	3			22	7.153	.931	2				
1961							26	7.141	.927	2			
Feb. 8	6.969	1.052	3	6.964	+1.043	1961	Feb. 8	7.168	.920	3	7.163	+0.911	
10	6.986	1.067	3	6.979	1.061	10	7.188	.946	3	7.181	.940		
15	6.980	1.046	3	6.967	1.046	15	7.168	.919	3	7.155	.919		
21	6.977	1.044	3			21	7.178	.921	1				
22	6.966	1.034	4	6.976	1.030	22	7.156	.915	4	7.166	.911		
24	6.942	1.064	3	6.950	1.050	24	7.135	.947	3	7.143	.933		
Mar. 1	6.961	1.052	4	6.954	1.049	Mar. 1	7.157	.937	4	7.150	.934		
8	6.963	1.056	4	6.951	1.054	8	7.165	.930	4	7.153	.928		
23	6.961	1.046	3	6.957	1.045	23	7.155	.922	3	7.151	.921		
31	6.972	1.046	3	6.966	1.061	31	7.170	.918	2	7.164	.933		
Apr. 4	6.974	1.048	3	6.967	1.043	Apr. 12	7.180	+0.921	3	7.177	+0.920		
12	6.979	+1.051	4	6.976	+1.050					mean 7.160	+0.925		
				mean 6.964	+1.048								
UA' 54 = HD 58551													
DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V	DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V		
1954													
Mar. 30	6.544	+0.468	2			1955	Apr. 20	8.939	+0.354	4	8.948	+0.353	
Apr. 1	6.544	.474	3			25	8.943	.340	4				
6	6.540	.466	2			May 4	8.961	.361	2	8.964	.364		
22	6.545	.476	2			1956	Jan. 17	8.943	.353	4	8.951	.350	
26	6.540	.474	2			Feb. 3	8.944	.358	4	8.942	.360		
1961							10	8.968	.352	4	8.959	.351	
Feb. 6	6.547	.454	4	6.547	+0.452	14	8.965	.325	5	8.955	.321		
10	6.553	.465	3	6.546	.459	Mar. 5	8.932	.348	2	8.970	.353		
15	6.560	.455	3	6.547	.455	14	8.943	.349	3				
21	6.549	.472	1			19	8.952	.357	3	8.955	.349		
22	6.538	.484	4	6.548	.480								
UA 55 = HD 63772													
DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V	DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V		
1955													
Apr. 20	8.939	+0.354	4			1956	Jan. 17	8.943	.353	4	8.951	.350	
25	8.943	.340	4			Feb. 3	8.944	.358	4	8.942	.360		
May 4	8.961	.361	2			10	8.968	.352	4	8.959	.351		
1956	Jan. 17	8.943	.353			14	8.965	.325	5	8.955	.321		
Feb. 3	8.944	.358	4			Mar. 5	8.932	.348	2	8.970	.353		
10	8.968	.352	4			14	8.943	.349	3				
14	8.965	.325	5			19	8.952	.357	3	8.955	.349		

**Table V. Observations of Comparison Stars (Cont'd)**

UA 55 = HD 63772							NB 55 = HD 119869														
DATE	V			B-V		Wt.	V			B-V		Wt.	V			B-V					
	System of primary stds.		B-V	System of 10-year stds.			System of primary stds.		B-V	System of 10-year stds.			System of primary stds.		B-V	System of 10-year stds.					
1956													1955								
Apr.	5	8.958	+ 0.345	4	8.951	+ 0.347							Apr.	21	8.922	+ 0.442	3	8.923	+ 0.438		
	18	8.960	.370	2	8.956	.379							25	8.902	.467	2					
1961													May	4	8.916	.470	2	8.919	.473		
Apr.	12	8.952	+ 0.359	3	8.949	+ 0.358							Jun.	20	8.910	.454	2				
					mean 8.953	+ 0.350							21	8.905	.473	2					
	UB 55 = HD 62720												1956								
	V			B-V			V			B-V			Jan.	17	8.893	.452	3	8.901	.449		
	System of primary stds.		B-V	System of 10-year stds.			System of primary stds.		B-V	System of 10-year stds.			Feb.	10	8.901	.460	3	8.892	.459		
	DATE												Mar.	14	8.910	.452	2				
1955													19	8.873	.480	4	8.876	.472			
Apr.	20	7.400	+ 0.386	4	7.409	+ 0.385							Apr.	5	8.892	.459	3	8.885	.461		
	25	7.388	.391	4									18	8.913	.429	1	8.909	.438			
May	4	7.416	.393	2	7.419	.396							Apr.	12	8.895	+ 0.472	3	8.892	+ 0.471		
1956													mean	8.897	+ 0.459						
Jan.	17	7.393	.394	4	7.401	.391							UA 56 = HD 67228 = $\mu$ Cnc.								
Feb.	3	7.410	.378	4	7.408	.380							V			B-V		V		B-V	
													DATE				System of primary stds.	Wt.	System of 10-year stds.		
1955													1955								
Apr.	10	7.425	.378	4	7.416	.377							Apr.	11	5.256	+ 0.619	4				
	14	7.423	.357	5	7.413	.353							12	5.301	.622	5					
Mar.	5	7.370	.390	2	7.408	.395							21	5.321	.642	4	5.322	+ 0.638			
	14	7.400	.378	3									25	5.287	.640	4					
	19	7.401	.391	3	7.404	.383							May	4	5.308	.637	2	5.311	.640		
Apr.	5	7.404	.388	4	7.397	.390							Oct.	27	5.297	.629	3				
	18	7.415	.378	2	7.411	.387							28	5.293	.632	3					
1961													30	5.280	.633	2					
Apr.	12	7.424	+ 0.384	3	7.421	+ 0.383							31	5.283	.625	3					
					mean 7.409	+ 0.381							Nov.	4	5.308	.639	3				
	NA 55 = HD 120186												19	5.304	.633	4					
	V			B-V			V			B-V			20	5.304	.641	4					
	System of primary stds.		B-V	System of 10-year stds.			System of primary stds.		B-V	System of 10-year stds.			27	5.304	.637	5					
1955													1956								
Apr.	21	7.730	+ 0.544	3	7.731	+ 0.540							Jan.	17	5.287	.646	4	5.295	.643		
	25	7.763	.496	2									Feb.	3	5.292	.644	4	5.290	.646		
May	4	7.729	.541	2	7.732	.544							10	5.306	.636	4	5.297	.635			
Jun.	20	7.723	.557	2									14	5.326	.620	5	5.316	.616			
	21	7.731	.553	2									Mar.	5	5.271	.628	2	5.309	.633		
1956													14	5.307	.625	3					
Jan.	17	7.704	.546	3	7.712	.543							19	5.302	.633	3	5.305	.625			
	10	7.704	.544	3	7.695	.543							Apr.	5	5.287	.634	4	5.280	.636		
Mar.	14	7.713	.549	2									18	5.308	.630	2	5.304	.639			
	19	7.711	.561	4	7.714	.553							1957								
Apr.	5	7.717	.546	3	7.710	.548							Mar.	23	5.296	.633	4	5.295	.635		
	18	7.718	.530	1	7.714	.539							24	5.309	.644	4	5.314	.638			
1961													26	5.307	.643	1					
Apr.	12	7.714	+ 0.550	3	7.711	+ 0.549							mean 7.714	+ 0.546							

**Table V. Observations of Comparison Stars (Cont'd)**

UA 56 = HD 67228 = $\mu$ Cnc.							UB 56 = HD 68256 = $\delta$ Cnc (A + B + C)						
DATE	V System of primary stds.	B-V	V System of 10-year stds.	B-V	DATE	V System of primary stds.	B-V	V System of primary stds.	B-V	V System of 10-year stds.	B-V		
1957					1961								
Apr.	5	5.293	+ 0.633	4	Apr.	12	4.681	+ 0.540	3	4.678	+ 0.539		
	9	5.307	.632	2					mean	4.674	+ 0.530		
	15	5.305	.626	2									
	16	5.310	.633	1									
	19	5.305	.632	1									
	20	5.299	.638	2									
May	1	5.325	.650	1									
	3	5.308	.639	3									
	5	5.310	.628	1									
1961					1956								
Apr.	12	5.313	+ 0.634	3	Jan.	17	6.907	+ 0.542	2	6.915	+ 0.539		
						19	6.900	.541	1				
						20	6.899	.535	1				
						Feb.	6.941	.527	2	6.939	.529		
							10	6.905	.546	2	6.896	.545	
						Mar.	6.907	.566	4	6.910	.558		
						Apr.	6.918	.534	2	6.911	.536		
							18	6.933	.529	0	6.929	.538	
1955					1957								
Oct.	27	4.663	+ 0.531	3	Mar.	23	6.900	.546	2	6.899	.548		
	28	4.671	.527	3		24	6.906	.542	2	6.911	.536		
	30	4.657	.531	2	Apr.	5	6.942	.533	1				
	31	4.648	.528	3		6	6.905	.531	4				
Nov.	4	4.678	.539	3		18	6.905	.539	2				
	19	4.669	.535	4		24	7.086	.531	0				
	20	4.670	.538	4	May	1	6.861	.529	0	6.850	.525		
	27	4.675	.536	5		3	6.906	.552	1	6.909	.551		
1956						30	6.910	.531	1	6.910	.531		
Jan.	17	4.661	.543	4	Jun.	1	6.915	.544	1				
Feb.	3	4.652	.543	4	1961	12	6.925	+ 0.540	2	6.922	+ 0.539		
	10	4.675	.532	4					mean	6.912	+ 0.543		
	14	4.686	.520	4									
Mar.	5	4.637	.533	2									
	14	4.658	.533	4									
	19	4.677	.525	3									
Apr.	5	4.683	.510	4									
	18	4.675	.533	2									
1957													
Mar.	23	4.674	.526	4									
	24	4.691	.532	4									
	26	4.681	.540	1									
Apr.	5	4.665	.526	3									
	9	4.679	.528	1									
	15	4.671	.531	2									
	16	4.681	.531	1									
	19	4.677	.518	1									
	20	4.665	.536	2									
May	1	4.689	.539	1									
	3	4.683	.529	3									
	5	4.670	.531	1									

**Table V. Observations of Comparison Stars (Cont'd)**

NB 56 = HD 121111						UB 57 = HD 73665 = 39 Cnc						
DATE	V System of primary stds.	B-V	V System of 10-year stds.	B-V	DATE	V System of primary stds.	B-V	V System of 10-year stds.	B-V			
1957					1955							
Apr.	5	7 <sup>m</sup> .735	+ 0 <sup>m</sup> .538	2	Oct.	27	6 <sup>m</sup> .380	+ 0 <sup>m</sup> .966	3			
	6	7.690	.540	4		28	6.381	.981	3			
	18	7.736	.509	2		30	6.370	.980	2			
	24	7.799	.561	1		31	6.378	.976	3			
May	1	7.656	.546	1	7 <sup>m</sup> .645	+ 0 <sup>m</sup> .542						
	3	7.689	.550	2	7.692	.549	Nov.	4	6.397	.983	3	
	30	7.690	.540	2	7.690	.540		7	6.382	.979	3	
Jun.	1	7.695	.554	2		8	6.387	.973	3			
1961						16	6.383	.982	5			
Apr.	12	7.702	+ 0.552	3	7.699	+ 0.551		19	6.386	.982	4	
					mean	7.691	+ 0.549	20	6.387	.980	4	
								27	6.388	.986	5	
UA 57 = HD 72779 = 35 Cnc												
DATE	V System of primary stds.	B-V	V System of 10-year stds.	B-V	1956							
1955					Jan.	17	6.384	.975	3	6 <sup>m</sup> .392	+ 0 <sup>m</sup> .972	
Nov.	16	6 <sup>m</sup> .578	+ 0 <sup>m</sup> .681	5	Feb.	3	6.389	.978	4	6.387	.980	
1956						10	6.398	.977	4	6.389	.976	
Jan.	17	6.573	.681	4	14	6.381	.970	3	6.371	.966		
Feb.	3	6.585	.679	4	Mar.	5	6.355	.988	3	6.393	.993	
	10	6.592	.683	4		14	6.394	.978	4			
	14	6.589	.664	4		19	6.391	.990	4	6.394	.982	
Mar.	5	6.539	.688	2	1957	23	6.388	.975	5	6.387	.977	
	14	6.584	.666	4		24	6.410	.978	4	6.415	.972	
	19	6.578	.684	3		26	6.380	.982	1			
Apr.	5	6.582	.679	4		31	6.397	.990	1			
	18	6.594	.668	2		Apr.	5	6.389	.965	4		
1957							9	6.407	.973	2		
Mar.	23	6.573	.685	5			15	6.394	.970	3	6.391	.974
	24	6.596	.683	4		16	6.387	.982	2			
	26	6.567	.691	1		18	6.359	.967	2			
	31	6.589	.693	1		19	6.384	.980	2			
Apr.	5	6.576	.671	4		20	6.398	.978	2			
	7	6.583	.692	5		26	6.382	.985	1			
	9	6.593	.679	2		May	1	6.419	.985	1	6.408	.981
	15	6.584	.668	3		3	6.391	.980	3	6.394	.979	
	16	6.581	.675	2		5	6.408	.960	1	6.397	.963	
	18	6.560	.672	1	1959	May	28	6.373	.979	1		
	19	6.584	.674	1	1961	Apr.	12	6.391	+ 0.974	4	6.388	+ 0.973
	20	6.579	.678	2			mean	6.392	+ 0.977			
	26	6.583	.682	1								
May	1	6.601	.680	1								
	3	6.586	.678	3								
	5	6.608	.663	1								
1961												
Apr.	12	6.600	+ 0.678	4	6.597	+ 0.677						
					mean	6.584	+ 0.678					

Table V. Observations of Comparison Stars (Cont'd)

NA 57 = HD 121608							UA 58 = HD 75470							
DATE	V System of primary stds.	B-V	V System of 10-year stds.	B-V	DATE	V System of primary stds.	B-V	V System of 10-year stds.	B-V					
1956					1956									
Mar. 19	7.687	+0.556	4	7.690	+0.548	Jan. 17	6.703	+0.860	4	6.711	+0.857			
Apr. 5	7.690	.546	2	7.683	.548	Feb. 3	6.695	.885	4	6.693	.887			
18	7.694	.531	1	7.690	.540	10	6.730	.860	4	6.721	.859			
1957					14	6.726	.853	4	6.716	.849				
Mar. 23	7.685	.547	2	7.684	.549	Mar. 5	6.673	.870	2	6.711	.875			
24	7.692	.529	2	7.697	.523	19	6.713	.867	3	6.716	.859			
28	7.684	.519	2		Apr. 18	6.723	.852	2	6.719	.861				
Apr. 5	7.694	.534	2		1957									
6	7.671	.530	4		Mar. 23	6.711	.864	5	6.710	.866				
Apr. 18	7.677	.533	2		24	6.726	.865	4	6.731	.859				
24	7.811	.538	1		26	6.721	.879	1						
May 1	7.671	.524	1	7.660	.520	Apr. 5	6.710	.857	4					
3	7.671	.560	2	7.674	.559	9	6.719	.862	2					
5	7.664	.552	1	7.653	.555	15	6.723	.864	3	6.720	.868			
30	7.691	.527	2	7.691	.527	16	6.719	.862	1					
Jun. 1	7.681	.549	2		19	6.713	.867	1						
1961					20	6.714	.860	2						
Apr. 12	7.689	+0.547	2	7.686	+0.546	May 3	6.716	.864	3	6.719	.863			
				mean 7.684	+0.541	Dec. 27	6.697	.864	4					
NB 57 = HD 121496							1958							
DATE	V System of primary stds.	B-V	V System of 10-year stds.	B-V	Feb. 11	6.706	.868	2	6.715	.862				
					27	6.697	.867	4						
1956					Mar. 20	6.724	.874	2	6.718	.872				
Jan. 19	6.841	+0.470	2		Apr. 10	6.725	.864	4	6.725	.859				
1957					11	6.729	.865	4	6.729	.859				
Mar. 23	6.846	.479	2	6.845	+0.481	15	6.732	.859	3	6.728	.855			
24	6.854	.470	2	6.859	.464	May 4	6.733	.853	3	6.735	.852			
28	6.813	.470	2		9	6.734	.849	3	6.731	.857				
Apr. 5	6.861	.459	2		10	6.732	.879	3	6.726	.874				
6	6.836	.462	4		Jun. 8	6.720	.866	1						
18	6.837	.471	2		9	6.708	.878	3						
24	6.960	.482	1		16	6.711	.865	4	6.714	.865				
May 1	6.829	.463	1	6.818	.459	1961	Apr. 12	6.732	+0.864	4	6.729	+0.863		
3	6.844	.476	2	6.847	.475			mean 6.720	+0.863					
5	6.850	.452	1	6.839	.455	UB 58 = HD 75974								
30	6.844	.464	2	6.844	.464	DATE	V System of primary stds.	B-V	V System of 10-year stds.	B-V				
Jun. 1	6.843	.483	2		1955									
1961					Nov. 16	6.671	+0.706	5						
Apr. 12	6.853	+0.478	3	6.850	+0.477	1956								
				mean 6.846	+0.470	Jan. 17	6.661	.712	4	6.669	+0.709			
UA 58 = HD 75470							Feb. 3	6.676	.700	4	6.674	.702		
DATE	V System of primary stds.	B-V	V System of 10-year stds.	B-V	10	6.711	.698	4	6.702	.697				
					14	6.687	.685	4	6.677	.681				
1955					Mar. 5	6.637	.697	2	6.675	.702				
Nov. 16	6.705	+0.873	5											

**Table V. Observations of Comparison Stars (Cont'd)**

UB 58 = HD 75974								NA 58 = HD 123453							
DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V	DATE	V System of primary stds.	B-V	Wt.	V System of 10-year stds.	B-V				
1956						1958									
Mar. 19	6 <sup>m</sup> .678	+ 0 <sup>m</sup> .704	3	6 <sup>m</sup> .681	+ 0 <sup>m</sup> .696	Feb. 27	7 <sup>m</sup> .647	+ 0 <sup>m</sup> .594	1						
Apr. 18	6.683	.698	2	6.679	.707	Mar. 20	7.633	.585	1	7 <sup>m</sup> .627	+ 0 <sup>m</sup> .583				
1957						Apr. 10	7.626	.577	2	7.626	.572				
Mar. 23	6.683	.690	5	6.682	.692		11	7.628	.584	2	7.628	.578			
24	6.691	.704	4	6.696	.698		15	7.623	.586	2	7.619	.582			
	6.686	.716	1			May 4	7.629	.586	2	7.631	.585				
Apr. 5	6.666	.697	4				9	7.619	.583	2	7.616	.591			
9	6.680	.701	2				10	7.625	.589	2	7.619	.584			
	6.671	.703	2	6.668	.707	Jun. 9	7.620	.586	2						
16	6.683	.709	1				16	7.640	.574	2	7.643	.574			
19	6.673	.707	1												
20	6.677	.702	2												
May 3	6.664	.700	3	6.667	.699										
Dec. 27	6.666	.706	4												
1958															
Feb. 11	6.666	.710	2	6.675	.704										
27	6.654	.713	4												
Mar. 20	6.685	.709	2	6.679	.707										
Apr. 11	6.692	.713	4	6.692	.707										
15	6.686	.712	3	6.682	.708	May 30	5 <sup>m</sup> .462	+ 0 <sup>m</sup> .334	2	5 <sup>m</sup> .462	+ 0 <sup>m</sup> .334				
May 4	6.698	.699	4	6.700	.698	Jun. 1	5.469	.348	2						
9	6.689	.686	3	6.686	.694		13	5.463	.352	3					
10	6.714	.693	3	6.708	.688		14	5.477	.346	0					
Jun. 8	6.669	.700	1				22	5.474	.351	2					
9	6.664	.717	3				24	5.464	.348	1					
16	6.677	.707	4	6.680	.707		25	5.454	.354	1					
1961															
Apr. 12	6.693	+ 0.704	4	6.690	+ 0.703	1958									
				mean 6.684	+ 0.699	Feb. 11	5.446	.347	4	5.455	.341				
						27	5.459	.344	1						
						Mar. 20	5.474	.356	2	5.468	.354				
						Apr. 10	5.452	.353	2	5.452	.348				
						11	5.472	.353	2	5.472	.347				
						15	5.474	.343	2	5.470	.339				
						May 4	5.456	.355	3	5.458	.354				
						9	5.461	.336	2	5.458	.344				
						10	5.478	.330	2	5.472	.325				
1956						Jun. 9	5.458	.345	2						
Jan. 20	7 <sup>m</sup> .629	+ 0 <sup>m</sup> .588	2			16	5.468	.346	2	5.471	.346				
Mar. 14	7.655	.593	2												
	19	7.629	.590	4	7 <sup>m</sup> .632	+ 0 <sup>m</sup> .582									
Apr. 5	7.650	.581	2	7.643	.583										
18	7.657	.556	1	7.653	.565										
1957															
Mar. 24	7.635	.574	2	7.640	.568										
Apr. 5	7.650	.598	1												
6	7.656	.593	4												
May 30	7.632	.583	1	7.632	.583										
Jun. 1	7.629	.572	1												
	25	7.632	.588	0											
1958															
Feb. 11	7.633	.555	4	7.642	.549										

Table V. Observations of Comparison Stars (Cont'd)

UA 59 = HD 79096							UB 59 = HD 79499						
	V	B-V	V	B-V		V	B-V	V	B-V		V	B-V	
DATE	System of primary stds.	Wt.	System of 10-year stds.		DATE	System of primary stds.	Wt.	System of 10-year stds.		DATE	System of primary stds.	Wt.	
1958					1959					1959			
Mar. 20	6.502	+ 0.743	2	6.496	+ 0.741	Mar. 3	8.290	+ 0.434	4	8.313	+ 0.448		
Apr. 10	6.497	.747	4	6.497	.739	1961							
11	6.508	.732	4	6.508	.726	Mar. 23	8.303	.443	3	8.299	.442		
15	6.510	.735	3	6.506	.731	Apr. 12	8.307	+ 0.457	3	8.304	+ 0.456		
May 4	6.517	.737	3	6.519	.736					mean	8.315	+ 0.446	
9	6.509	.726	2	6.506	.734								
10	6.525	.729	3	6.519	.724								
1959										NA 59 = HD 124401			
Jan. 5	6.479	.725	4	6.511	.729	V	B-V	V	B-V				
12	6.498	.722	2	6.503	.724	System of primary stds.		System of 10-year stds.					
24	6.500	.718	3	6.504	.734	DATE		Wt.					
29	6.490	.729	5			1957							
31	6.495	.719	4	6.504	.724	Mar. 23	6.990	+ 1.018	2	6.989	+ 1.020		
Feb. 4	6.486	.727	5	6.501	.732	24	7.006	0.989	2	7.011	0.983		
5	6.504	.721	5	6.509	.722	1958							
26	6.490	.733	5	6.506	.725	Feb. 27	7.009	1.018	1				
Mar. 1	6.496	.730	2	6.498	.740	Mar. 20	7.008	1.022	2	7.002	1.020		
3	6.492	.720	4	6.515	.734	Apr. 10	6.967	1.037	2	6.967	1.032		
1961						11	6.971	1.029	2	6.971	1.023		
Mar. 23	6.514	.728	3	6.510	.727	15	6.996	1.022	2	6.992	1.018		
31	6.510	.734	2	6.504	.749	May 4	6.984	1.014	2	6.986	1.013		
Apr. 12	6.500	+ 0.742	3	6.497	+ 0.741	9	6.981	1.015	2	6.978	1.023		
				mean	6.506	+ 0.731	10	6.975	1.022	2	6.969	1.017	
	UB 59 = HD 79499												
	V	B-V	V	B-V									
DATE	System of primary stds.	Wt.	System of 10-year stds.										
1958													
Feb. 11	8.299	+ 0.459	2	8.308	+ 0.453	1959							
27	8.284	.446	4			Jan. 5	6.900	0.993	0	6.932	0.997		
Mar. 20	8.318	.443	2	8.312	.441	12	6.976	1.006	2	6.981	1.008		
Apr. 10	8.315	.460	4	8.315	.455	24	6.977	1.016	2	6.981	1.032		
11	8.308	.450	4	8.308	.444	31	6.972	1.010	1	6.981	1.015		
15	8.333	.440	3	8.329	.436	Feb. 4	6.975	1.015	1	6.990	1.020		
May 4	8.330	.439	3	8.332	.438	5	6.969	1.023	0	6.974	1.024		
9	8.318	.438	2	8.315	.446	26	6.974	0.994	0	6.990	0.986		
10	8.325	.465	3	8.319	.460	28	6.996	1.003	2	6.990	1.004		
1959						Mar. 1	6.969	1.005	2	6.971	1.015		
Jan. 5	8.277	.440	4	8.309	.444	3	6.952	1.016	2	6.975	1.030		
12	8.315	.425	2	8.320	.427	1961							
24	8.306	.427	3	8.310	.443	Apr. 12	6.994	+ 1.018	2	6.991	+ 1.017		
29	8.303	.440	5			mean	6.984	+ 1.017					
31	8.298	.444	4	8.307	.449								
Feb. 4	8.283	.451	5	8.298	.456								
5	8.312	.437	5	8.317	.438								
26	8.303	.456	5	8.319	.448								
28	8.356	.439	3	8.350	.440								
Mar. 1	8.322	.435	2	8.324	.445								
	NB 59 = HD 125337 = $\lambda$ Vir												
	V	B-V	V	B-V									
	System of primary stds.	Wt.	System of 10-year stds.										
1958													
Feb. 27	4.511	+ 0.125	1										
Mar. 20	4.523	.130	1	4.517	+ 0.128								
Apr. 10	4.498	.148	2	4.498	.143								
11	4.500	.145	2	4.500	.139								
15	4.527	.123	2	4.523	.119								
May 4	4.512	.132	2	4.514	.131								

**Table V. Observations of Comparison Stars (Cont'd)**

NB 59 = HD 125337 = $\lambda$ Vir								UB 60 = HD 82140									
DATE	V	B-V	System of primary stds.	Wt.	V	B-V	System of 10-year stds.	DATE	V	B-V	System of primary stds.	Wt.	V	B-V	System of 10-year stds.		
1958																	
May	9	4 <sup>m</sup> .510	+0 <sup>m</sup> .122	2	4 <sup>m</sup> .507	+0 <sup>m</sup> .130		May	15	8 <sup>m</sup> .395	+0 <sup>m</sup> .629	1					
	10	4.505	.142	2	4.499	.137			16	8.394	.638	1					
1959																	
Jan.	5	4.483	.136	0	4.475	.140		Jan.	20	8.380	.628	4					
	12	4.518	.126	2	4.523	.128			28	8.372	.641	3	8 <sup>m</sup> .375	+0 <sup>m</sup> .641			
	24	4.534	.118	1	4.538	.134			29	8.375	.641	4					
	31	4.497	.129	0	4.506	.134		Feb.	17	8.376	.638	2	8.374	.643			
Feb.	4	4.527	.119	1	4.542	.124			18	8.380	.624	3					
	5	4.521	.111	0	4.526	.112			21	8.390	.628	2	8.381	.628			
	26	4.509	.118	0	4.525	.110			23	8.385	.633	2	8.386	.631			
	28	4.533	.108	1	4.527	.109			25	8.377	.641	1					
Mar.	1	4.515	.120	2	4.517	.130		Mar.	15	8.370	.640	3	8.371	.647			
	3	4.497	.110	1	4.520	.124			17	8.386	.634	3	8.380	.640			
1961																	
Apr.	12	4.526	+0.138	2	4.523	+0.137			18	8.403	.628	2	8.388	.627			
					mean 4.515	+0.131			19	8.365	.640	2	8.373	.641			
UA 60 = HD 81563																	
DATE	V	B-V	System of primary stds.	Wt.	V	B-V	System of 10-year stds.	Apr.	3	8.370	.614	4	8.376	.618			
									4	8.373	.630	2					
									6	8.395	.626	3	8.387	.632			
1959																	
May	9	8 <sup>m</sup> .319	+0 <sup>m</sup> .466	3				1961	Apr.	12	8.373	+0.644	3	8.370	+0.643		
	15	8.284	.483	1									mean 8.377	+0.635			
	16	8.310	.468	2													
1960																	
Jan.	20	8.282	.479	4					NA 60 = HD 126251								
	28	8.278	.480	3	8 <sup>m</sup> .281	+0 <sup>m</sup> .480											
	29	8.281	.479	4				DATE	V	B-V	System of primary stds.	Wt.	V	B-V	System of 10-year stds.		
Feb.	17	8.272	.477	2	8.270	.482											
	18	8.276	.486	3				1959	Jan.	5	6 <sup>m</sup> .451	+0 <sup>m</sup> .410	0	6 <sup>m</sup> .483	+0 <sup>m</sup> .414		
	21	8.285	.493	2	8.276	.493			12	6.477	.415	2	6.482	.417			
	23	8.294	.472	2	8.295	.470			Feb.	4	6.482	.412	1	6.497	.417		
	25	8.277	.491	2						5	6.468	.418	0	6.473	.419		
Mar.	15	8.277	.479	3	8.278	.486			1960	Jan.	28	6.493	.407	2	6.496	.407	
	17	8.293	.473	3	8.287	.479				Feb.	17	6.475	.419	1	6.473	.424	
	18	8.306	.477	2	8.291	.476					21	6.489	.417	2	6.480	.417	
	19	8.270	.477	2	8.278	.478					23	6.496	.410	1	6.497	.408	
	21	8.272	.500	3	8.278	.475			Mar.	15	6.489	.413	2	6.490	.420		
	26	8.284	.474	2	8.277	.476				17	6.487	.414	2	6.481	.420		
	30	8.268	.482	3	8.273	.484				18	6.490	.422	3	6.475	.421		
Apr.	3	8.278	.456	4	8.284	.460				19	6.497	.414	3	6.505	.415		
	4	8.270	.485	2						21	6.492	.444	2	6.498	.419		
	6	8.292	.461	3	8.284	.467				30	6.491	.417	2	6.496	.419		
1961																	
Apr.	12	8.268	-0.497	3	8.265	+0.496		Apr.	3	6.511	.422	1	6.517	.426			
					mean 8.280	+0.478											

Table V. Observations of Comparison Stars (Cont'd)

NA 60 = HD 126251							UA 61 = HD 83683								
DATE	V	B-V	System of primary stds.	Wt.	V	B-V	System of 10-year stds.	DATE	V	B-V	System of primary stds.	Wt.	V	B-V	System of 10-year stds.
1961								1960							
Apr. 12	6 <sup>m</sup> .497	+0 <sup>m</sup> .412	2		6 <sup>m</sup> .494	+0 <sup>m</sup> .411		Mar. 30	6 <sup>m</sup> .966	+0 <sup>m</sup> .469	3	6 <sup>m</sup> .971	+0 <sup>m</sup> .471		
					mean 6.491	+0.417		Apr. 3	6.956	.470	4	6.962	.474		
								6	6.994	.451	3	6.986	.457		
	NB 60 = HD 126766							1961							
DATE	V	B-V	System of primary stds.	Wt.	V	B-V	System of 10-year stds.	Feb.	6	6.962	.492	3	6.962	.490	
								10	6.965	.483	3	6.958	.477		
1959								24	6.966	.487	3	6.974	.473		
Jan. 5	6 <sup>m</sup> .633	+0 <sup>m</sup> .423	0		6 <sup>m</sup> .665	+0 <sup>m</sup> .427		Mar. 23	6.989	.457	3	6.985	.456		
12	6.643	.422	2		6.648	.424		31	6.987	.460	3	6.981	.475		
Feb. 4	6.650	.415	1		6.665	.420		Apr. 4	6.972	.460	2	6.965	.455		
5	6.641	.417	0		6.646	.418		9	6.948	.474	2	6.954	.480		
1960								12	6.964	+0.480	3	6.961	+0.479		
Jan. 28	6.642	.425	2		6.645	.425						mean 6.968	+0.473		
Feb. 17	6.636	.428	1		6.634	.433									
21	6.643	.416	1		6.634	.416									
23	6.648	.413	1		6.649	.411									
Mar. 15	6.652	.429	2		6.653	.436									
17	6.654	.405	2		6.648	.411									
18	6.660	.411	3		6.645	.410									
19	6.647	.432	3		6.655	.433									
21	6.662	.454	2		6.668	.429									
30	6.659	.422	2		6.664	.424									
Apr. 3	6.660	.441	0		6.666	.445									
1961															
Apr. 12	6.655	+0.424	2		6.652	+0.423									
					mean 6.652	+0.423									
	UA 61 = HD 83683														
DATE	V	B-V	System of primary stds.	Wt.	V	B-V	System of 10-year stds.								
1960															
Jan. 20	6 <sup>m</sup> .983	+0 <sup>m</sup> .468	1												
28	6.965	.470	3		6 <sup>m</sup> .968	+0 <sup>m</sup> .470									
Feb. 17	6.958	.479	1		6.956	.484									
18	6.966	.469	3												
21	6.981	.465	2		6.972	.465									
23	6.971	.476	2		6.972	.474									
25	6.972	.463	1												
Mar. 15	6.970	.458	3		6.971	.465									
17	6.959	.480	3		6.953	.486									
18	6.983	.475	2		6.968	.474									
19	6.958	.474	2		6.966	.475									
21	6.960	.500	2		6.966	.475									
26	6.967	.479	1		6.960	.481									

**Table V. Observations of Comparison Stars (Cont'd)**

NA 61 = HD 128596							NB 61 = HD 128429								
DATE	V	B-V	System of primary stds.	Wt.	V	B-V	System of 10-year stds.	DATE	V	B-V	System of primary stds.	Wt.	V	B-V	
	System of primary stds.	Wt.			System of 10-year stds.				mean	Wt.			System of 10-year stds.		
1960								1961							
Jan. 28	7.464	+0.651	2	7.467	+0.651			Apr. 12	6.206	-0.456	2	6.203	+0.455		
Feb. 17	7.470	.645	1	7.468	.650					mean	6.201	+0.461			
21	7.474	.660	1	7.465	.660										
23	7.480	.641	1	7.481	.639										
Mar. 15	7.473	.646	2	7.474	.653										
17	7.481	.646	2	7.475	.652										
18	7.480	.647	3	7.465	.646										
19	7.479	.650	3	7.487	.651										
21	7.483	.680	2	7.489	.655										
30	7.483	.657	2	7.488	.659										
Apr. 3	7.492	.665	1	7.498	.669										
1961								1961							
Feb. 6	7.454	.667	0	7.454	.665			Feb. 6	8.099	+0.547	4	8.099	+0.545		
8	7.505	.652	2	7.500	.643			8	8.076	.545	3	8.071	.536		
22	7.459	.634	1	7.469	.630			15	8.092	.543	3	8.079	.543		
Mar. 1	7.508	.656	2	7.501	.653			24	8.086	.558	3	8.094	.544		
23	7.538	.644	2	7.534	.643			Mar. 23	8.106	.543	3	8.102	.542		
31	7.494	.643	2	7.488	.658			31	8.099	.533	3	8.093	.547		
Apr. 9	7.480	.634	2	7.486	.640			Apr. 4	8.092	.547	4	8.085	.542		
12	7.480	+0.652	2	7.477	+0.651			9	8.076	.537	2	8.082	.543		
				mean	7.485	+0.650		12	8.083	+0.555	3	8.080	+0.554		
											mean	8.088	+0.544		
	NB 61 = HD 128429								UB 62 = HD 86898						
DATE	V	B-V	System of primary stds.	Wt.	V	B-V	System of 10-year stds.	DATE	V	B-V	System of primary stds.	Wt.	V	B-V	
	System of primary stds.	Wt.			System of 10-year stds.				mean	Wt.			System of 10-year stds.		
1960								1961							
Feb. 17	6.188	+0.457	1	6.186	+0.462			Feb. 15	7.794	+1.042	3	7.781	+1.042		
21	6.201	.450	2	6.192	.450			24	7.791	1.056	3	7.799	1.042		
23	6.191	.455	1	6.192	.453			Mar. 23	7.818	1.036	3	7.814	1.035		
Mar. 15	6.185	.469	2	6.186	.476			31	7.805	1.032	2	7.799	1.046		
17	6.180	.478	2	6.174	.484			Apr. 4	7.800	1.040	4	7.793	1.035		
18	6.202	.450	3	6.187	.449			9	7.786	1.022	2	7.792	1.028		
19	6.201	.460	3	6.209	.461			12	7.799	+1.047	3	7.796	+1.046		
21	6.220	.488	2	6.226	.463					mean	7.796	+1.039			
30	6.208	.457	2	6.213	.459										
Apr. 3	6.216	.469	1	6.222	.473										
1961									NA 62 = HD 129271						
DATE	V	B-V	System of primary stds.	Wt.	V	B-V	System of 10-year stds.	DATE	V	B-V	System of primary stds.	Wt.	V	B-V	
	System of primary stds.	Wt.			System of 10-year stds.				mean	Wt.			System of 10-year stds.		
1960								1961							
Feb. 17	6.188	+0.457	1	6.186	+0.462			Feb. 8	8.058	+0.837	2	8.053	+0.828		
21	6.201	.450	2	6.192	.450			10	8.052	.817	1	8.045	.811		
23	6.191	.455	1	6.192	.453			15	8.063	.843	2	8.050	.843		
Mar. 15	6.185	.469	2	6.186	.476			22	8.021	.812	1	8.031	.808		
17	6.180	.478	2	6.174	.484			Mar. 1	8.070	.842	2	8.063	.839		
18	6.202	.450	3	6.187	.449			23	8.051	.815	2	8.047	.814		
19	6.201	.460	3	6.209	.461			Apr. 9	8.041	.805	1	8.047	.811		
21	6.220	.488	2	6.226	.463			12	8.058	+0.811	2	8.055	+0.810		
30	6.208	.457	2	6.213	.459					mean	8.051	+0.823			
Apr. 3	6.216	.469	1	6.222	.473										
1961															
Feb. 8	6.202	.487	2	6.197	.478										
22	6.180	.464	1	6.190	.460										
Mar. 1	6.234	.474	2	6.227	.471										
23	6.204	.438	2	6.200	.437										
31	6.218	.450	2	6.212	.465										
Apr. 9	6.195	.449	2	6.201	.455										

the nights when the extinction was determined,  $\sigma(Q_{y1})$  in this equation was replaced by the mean error with which  $Q_{y1}$  was determined.

From equations (25) and (26) it can be concluded, that the highest accuracy of photometric observations is obtained when all the stars are observed at the same, fairly high, altitude, except for the few standard stars observed at different altitudes for determining the extinction coefficient.

*B. Results for Ten-Year Standards and Comparison stars.* The observations of the Ten-Year Standards are listed in Table IV. Only those nights are included for which at least 4 Ten-Year Standards are observed. However, in forming averages and for further discussion only the nights with observations of at least 12 Ten-Year Standards were used. Observations of comparison stars are listed in Table V.

The values V, B-V and their weight, given in the first three columns for each star in Tables IV and V, were computed from equations (3), (4) and (27), respectively, using the transformation coefficients derived from the observations of primary standard stars. The weighted mean of these magni-

tudes and colors was formed for each Ten-Year Standard using only the nights when at least 12 Ten-Year Standards were observed; these averages are given with their mean errors in Table VI. The deviations of individual observations of the Ten-Year Standards from these averages were computed. The weighted means of these deviations were formed for each night (given in the last two columns of Table IV) and added to all the observations of the Ten-Year Standards and comparison stars observed that night. The magnitudes and colors corrected in this way are given in Tables IV and V in the fifth and sixth columns for each star. These are the values which can be considered as determined by comparison with the mean magnitudes and colors of the Ten-Year Standards. These magnitudes and colors will henceforth be considered to be reduced to the system of the Ten-Year Standards.

The weighted mean magnitudes and colors of the Ten-Year Standards are given in Table IV for each season; they are averages of values reduced to the system of Ten-Year Standards. These mean values were used for computing the mean blue magni-

TABLE VI. Magnitudes and Colors of Standard and Comparison Stars  
TEN-YEAR STANDARDS (in system of primary standards)

HD	Star	V		B-V		n
			m.e.	m.e.		
58946	$\rho$ Gem	4.181	$\pm 0^m 002$	$\pm 0^m 315$	$\pm 0^m 001$	61
76943	10 UMa	3.973	.002	.434	.001	64
82885	11 LMi	5.409	.002	.769	.001	64
89449	40 Leo	4.794	.001	.449	.001	66
90839	36 UMaA	4.837	.002	.518	.001	67
102870	$\beta$ Vir	3.609	.002	.550	.001	62
109358	$\beta$ CVn	4.257	.001	.587	.001	66
113139	78 UMa	4.929	.002	.366	.001	67
114710	$\beta$ Com	4.252	.001	.572	.001	66
115043		6.827	.002	.604	.001	66
115617	61 Vir	4.742	.003	.709	.002	48
117176	70 Vir	4.976	.002	.713	.001	66
120136	$\tau$ Boo	4.496	.002	.479	.002	67
121370	$\eta$ Boo	2.683	.002	.583	.001	65
128167	$\sigma$ Boo	4.463	.002	.363	.001	61
121156	$\xi$ Boo	4.548	$\pm 0^m 002$	$+ 0^m 765$	$\pm 0^m 001$	57

COMPARISON STARS (in system of Ten-Year Standards)

HD	Star	V	m.e.	B-V	m.e.	n
55052	UA 53=48 Gem	5.859	$\pm 0^m 003$	$\pm 0^m 370$	$\pm 0^m 002$	11
50692	UB 53					0
61997	UA 54	7.129	.002	0.411	.002	11
60914	UB 54	6.964	.003	1.048	.003	11
58551	UA' 54	7.160	.003	0.925	.003	10
58899	UB' 54	6.545	.004	0.460	.003	11
63772	UA 55	8.953	.002	0.350	.004	11

TABLE VI. Magnitudes and Colors of Standard and Comparison Stars (Cont'd)

HD	Star	V	m.e.	B-V	m.e.	n
62720	UB 55	7.409	±0.002	0.381	±0.004	11
67228	UA 56=μ Cnc.	5.304	.003	0.635	.002	17
68256	UB 56=ξ Cnc	4.674	.003	0.530	.003	15
72779	UA 57=35 Cnc	6.584	.002	0.678	.002	15
73665	UB 57=39 Cnc	6.392	.003	0.977	.002	13
75470	UA 58	6.720	.002	0.863	.002	21
75974	UB 58	6.684	.003	0.699	.002	20
79096	UA 59	6.506	.002	0.731	.002	20
79499	UB 59	8.315	.003	0.446	.002	20
81563	UA 60	8.280	.002	0.478	.003	14
82140	UB 60	8.377	.002	0.635	.002	14
83683	UA 61	6.968	.002	0.473	.002	21
83509	UB 61	7.037	.002	0.471	.002	21
87176	UA 62	8.088	.004	0.544	.002	9
86898	UB 62	7.796	.004	1.039	.002	7
116681	NA 53					0
118704	NA 54	8.494	.004	0.541	.002	9
118705	NB 54	9.171	.004	0.461	.003	9
120186	NA 55	7.714	.004	0.546	.002	8
119869	NB 55	8.897	.006	0.459	.005	8
119638	NA 56	6.912	.003	0.543	.003	12
121111	NB 56	7.691	.004	0.549	.002	12
121608	NA 57	7.684	.003	0.541	.004	11
121496	NB 57	6.846	.004	0.470	.003	7
123453	NA 58	7.632	.003	0.576	.003	15
123255	NB 58=95 Vir	5.464	.002	0.343	.002	11
124401	NA 59	6.984	.003	1.017	.003	20
125337	NB 59=λ Vir	4.515	.003	0.131	.002	18
126251	NA 60	6.491	.003	0.417	.001	16
126766	NB 60	6.652	.002	0.423	.002	16
128596	NA 61=NB 62	7.485	.004	0.650	.002	19
128429	NB 61	6.201	.004	0.461	.003	17
129271	NA 62	8.051	±0.003	+0.823	±0.005	8

## OTHER STARS (in system of primary standards)

HD	Star	V	B-V	n	HD	Star	V	B-V	n
12929	α Ari	2.017	+1.137	11	87777		8.400	+0.578	9
18331	HR 875	5.171	+0.079	11	110379-80	γ VirAB	2.749	+0.353	35
55156	49 Gem	7.068	+0.036	3	114449		7.614	+0.403	4
62345	K Gem	3.574	+0.931	38	116658	α Vir*	0.994:	-0.227	40
61421	α CMi	0.384	+0.418	39	121325		6.188	+0.526	7
		7.630	+0.948	5	121865		7.047	+0.978	2
69994		5.830	+1.134	3	121981		6.957	+0.997	2
73344		6.899	+0.542	10	124897	α Boo	-0.039	+1.227	34
73666	40 Cnc	6.610	+0.001	41	126660	δ Boo	4.047	+0.499	28
75528	54 Cnc.	6.376	+0.636	10	128752		6.729	+1.005	11
76508		6.172	+1.002	9	128986		6.982	+1.631	5
79009		6.886	+0.062	9	129157		8.410	+0.867	8
81361		6.288	+0.966	7	130109	109 Vir	3.702	-0.009	29
81581		7.754	+0.308	7	134083	45 Boo	4.929	+0.429	25
83343		6.641	+0.438	16					
84722		6.450	+0.428	15					
							* Eclipsing variable		

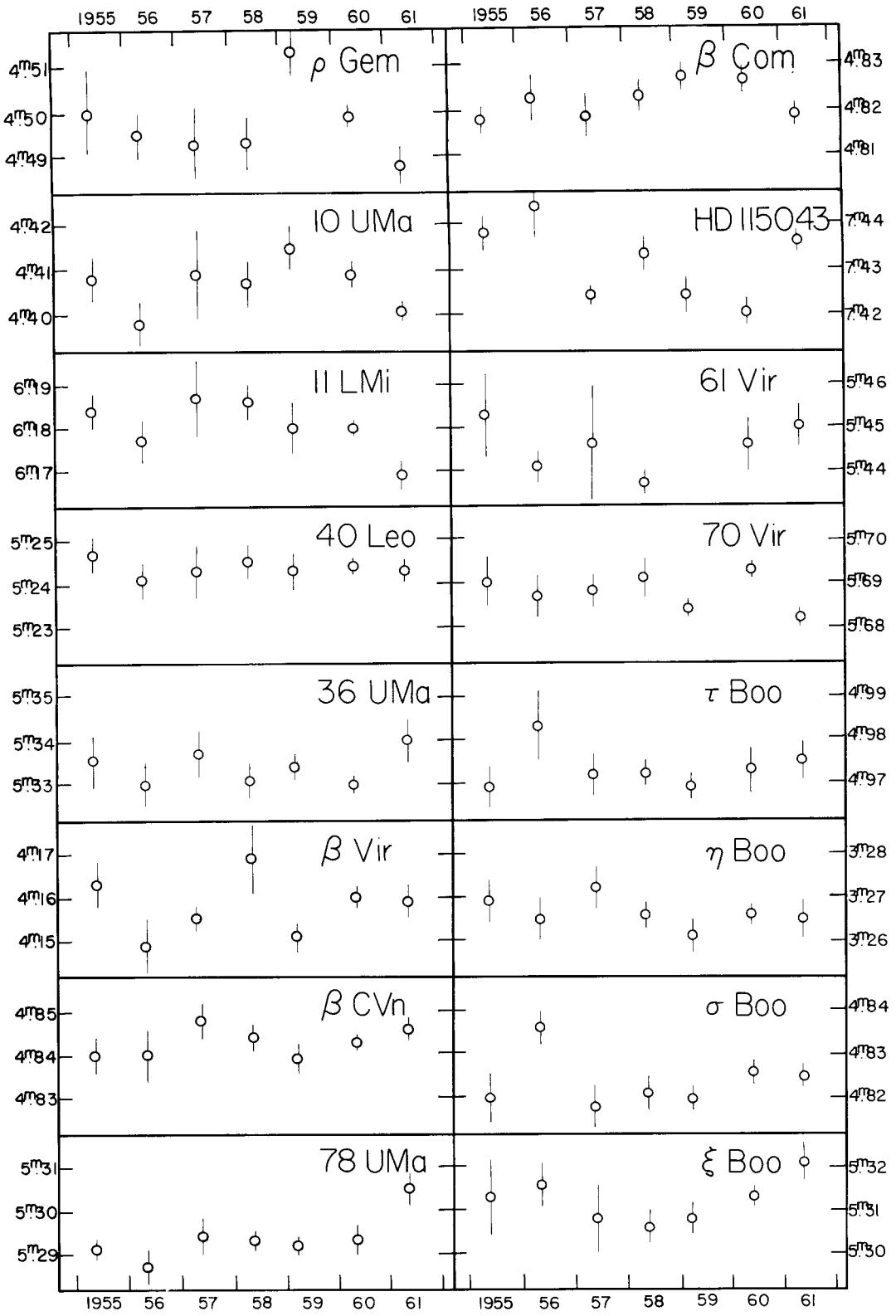


Figure 5. The yearly mean blue magnitudes of the Ten-Year Standard stars. Every point is based on 5 to 13 nights. The vertical lines indicate mean errors.

tudes which are plotted in Figure 5. Examining this figure, and also the nightly values given in Table IV, we can see that there seems to be no variability in the Ten-Year Standards which can not be explained by the errors of observations. For none of these stars does the r.m.s deviation of mean seasonal values from the final mean value exceed  $\pm 0.009$ .

The weighted mean magnitudes and colors of the comparison stars, reduced to the system of Ten-Year Standards, are given with their mean errors in Table VI. The last column in this table gives the number,  $n$ , of nights included in forming the average. Table VI also contains the weighted mean magnitudes and colors, not reduced to the system of Ten-Year Standards, for some other stars which were observed occasionally in this program. Most of these stars were candidates for comparison stars or Ten-Year Standards which were later rejected.

*C. Improved Magnitudes and Colors for Primary Standards of the UBV System.* The deviations,  $\delta_V$  and  $\delta_{B-V}$ , of observed magnitudes and colors of primary standard stars from the catalogue values were used for improving these last values. Only those 32 nights were used on which at least 6 different primary standard stars were observed and nightly extinction coefficients were determined. The mean  $\delta_V$  (computed by assuming  $Q_{V2} = -0.002$ ) and the mean  $\delta_{B-V}$  were computed for every star; let us denote these mean values by  $\bar{\delta}_V$  and  $\bar{\delta}_{B-V}$ . The values  $\bar{\delta}_V$  and  $\bar{\delta}_{B-V}$  were averaged for the stars  $\beta$  Cnc,  $\eta$  Hya,  $\alpha$  Ser,  $\beta$  Lib,  $\epsilon$  CrB and  $\tau$  Her; these average values are denoted by  $\delta^o_V$  and  $\delta^o_{B-V}$ . The quantities  $\Delta_V = \bar{\delta}_V - \delta^o_V$  and  $\Delta_{B-V} = \bar{\delta}_{B-V} - \delta^o_{B-V}$  are now treated as corrections which should be added to the magnitudes and colors given by Johnson and Harris (7). These corrections and also the corrected magnitudes and colors are given in Table VII, together with their mean errors computed from the scatter of values  $\delta_V$  and  $\delta_{B-V}$ . The mean magnitude and mean color of stars denoted as primary standards by Johnson and Harris remain, of course, unchanged. The last column of Table VII gives the number of observations used for

deriving the improved values. The corrections derived from observations in different years and with different multiplier tubes are in good agreement.

The improved values given in Table VII were obtained after finishing all the reductions of the observations of the present program. Therefore all the values of the transformation coefficients, the mean errors of single observations and the magnitudes and colors of Ten-Year Standards and comparison stars given in this paper are based on the data for primary standard stars given by Johnson and Harris (7).

#### IV. Magnitudes and Colors of Uranus and Neptune

*A. Color Indices of Uranus and Neptune.* The yellow magnitudes and color indices of Uranus and Neptune are given in Table VIII. They were derived exactly in the same way as the values for the comparison stars in Table V and both of these tables have the same arrangement.

The energy distribution in the spectra of Uranus and Neptune differs strongly from the energy distribution in spectra of color type stars. H. L. Johnson in Progress Report from 1955 writes:

"There has always been some doubt whether the blue-yellow color index B-V that has been determined from the standard U B V system filters, satisfactorily represents the planetary gradient in the blue filter spectral region. We would, in fact, expect the gradient determined from the ordinary B-V to be too blue because of the methane absorption in the yellow region. This is an important matter since this gradient enters into both the extinction and systemic reductions. We have, therefore, made color measures using two filters formed by splitting (from the spectral standpoint) the blue filter into two filters. The exact procedure that we used is the following: We selected a yellow-transmitting sharp cut-off filter glass whose cut-off wavelength lies approximately in the center of the blue filter. During these special observations of the planets and standard stars, deflections were taken with the standard blue filter and with the standard blue filter plus the above yellow filter. By subtraction, we obtain deflections corres-

Table VII. Improved Magnitudes and Colors for Primary Standards of the UBV System

Star	V	$\Delta_V$	B-V	$\Delta_{B-V}$	n
$\beta$ Cnc	3 <sup>m</sup> 534	$\pm 0.002$	+ 0 <sup>m</sup> 014	+ 1 <sup>m</sup> 477 $\pm 0.001$	- 0 <sup>m</sup> 003
$\eta$ Hya	4.299	.002	- .001	- 0.195 .001	.000
90 Leo AB	5.947	.003	—	- 0.159 .001	—
HR 4550	6.448	.003	—	+ 0.753 .002	—
$\beta$ Lib	2.608	.003	- .002	- 0.109 .002	- .001
$\alpha$ Ser	2.640	.002	- .010	+ 1.170 .002	+ .002
$\epsilon$ CrB	4.144	.002	- .006	+ 1.231 .001	+ .001
$\tau$ Her	3.895	.003	+ .005	- 0.151 .002	+ .001

**TABLE VIII. Two-color Observations of Uranus and Neptune**  
**(Uncorrected for distance, phase and oblateness effects)**

Uranus						Uranus						
DATE	V System of primary stds.	B-V	V System of 10-year stds.	B-V	DATE	V System of primary stds.	B-V	V System of 10-year stds.	B-V	DATE	V System of 10-year stds.	
1953					1957							
Oct. 30	5.518	+0.568	2		Apr. 15	5.525	+0.521	2	5.522	+0.525		
Nov. 6	5.520	.551	2		16	5.531	.524	2				
1954					19	5.524	.525	1				
Feb. 17	5.537	.558	2		20	5.529	.523	2				
23	5.563	.544	3		May 3	5.568	.523	3	5.571	.522		
24	5.572	.554	3		5	5.584	.497	1	5.573	.500		
28	5.577	.533	2		1959							
Mar. 30	5.539	.532	2		Jan. 5	5.394	.523	4	5.426	.527		
Apr. 1	5.537	.532	3		12	5.411	.513	2	5.416	.515		
6	5.545	.528	2		24	5.413	.507	3	5.417	.523		
22	5.584	.535	2		29	5.393	.527	5				
26	5.586	.543	3		31	5.391	.518	4	5.400	.523		
1955					Feb. 4	5.388	.521	5	5.403	.526		
Apr. 11	5.502	.520	4		5	5.393	.531	5	5.398	.532		
12	5.520	.534	5		26	5.396	.535	5	5.412	.527		
20	5.547	.535	4	5.556 +0.534		Mar. 1	5.407	.526	2	5.409	.536	
21	5.566	.537	4	5.567 .533		3	5.410	.512	4	5.433	.526	
25	5.546	.532	3		1960							
May 4	5.594	.535	2	5.597 .538		Jan. 20	5.407	.526	4			
Oct. 27	5.537	.523	3		28	5.408	.522	3	5.411	.522		
28	5.541	.523	3		29	5.400	.523	4				
30	5.528	.523	2		Feb. 17	5.380	.531	1	5.378	.536		
31	5.519	.527	3		18	5.390	.532	3				
Nov. 4	5.536	.531	3		21	5.414	.520	2	5.405	.520		
16	5.501	.534	5		23	5.411	.530	2	5.412	.528		
20	5.501	.528	4		25	5.400	.542	1				
19	5.507	.532	4		Mar. 15	5.412	.537	3	5.413	.544		
27	5.480	.529	5		17	5.418	.533	3	5.412	.539		
1956					18	5.442	.524	2	5.427	.523		
Jan. 17	5.413	.541	4	5.421 .538		19	5.422	.527	1	5.430	.528	
Feb. 3	5.405	.549	4	5.403 .551		21	5.422	.549	3	5.428	.524	
10	5.440	.523	4	5.431 .522		26	5.434	.533	1	5.427	.535	
14	5.454	.518	5	5.444 .514		30	5.435	.529	3	5.440	.531	
Mar. 5	5.433	.527	2	5.491 .532		Apr. 3	5.436	.524	4	5.442	.528	
14	5.474	.523	4		6	5.476	.515	3	5.468	.521		
19	5.487	.532	3	5.490 .524		1961						
Apr. 5	5.511	.529	4	5.504 .531		Feb. 8	5.384	.525	3	5.379	.516	
18	5.550	.527	2	5.546 .536		10	5.387	.540	3	5.380	.534	
1957						24	5.389	.538	3	5.397	.524	
Mar. 23	5.470	.525	5	5.469 .527		Mar. 23	5.441	.512	3	5.437	.511	
24	5.486	.533	4	5.491 .527		31	5.415	.538	3	5.409	.553	
26	5.482	.526	1		Apr. 4	5.430	.517	2	5.423	.512		
31	5.499	.529	1		8	5.426	.510	2				
Apr. 5	5.496	.516	4		12	5.456 +0.531	3	5.453 +0.530				
9	5.519	.517	2									

TABLE VIII. Two-color Observations of Uranus and Neptune (Cont'd)

(Uncorrected for distance, phase and oblateness effects)

DATE	Neptune				DATE	Neptune				
	V System of primary stds.	B-V Wt.	V System of 10-year stds.	B-V Wt.		V System of primary stds.	B-V Wt.	V System of 10-year stds.	B-V Wt.	
<b>1954</b>										
May 31	7.901	+ 0.404	2			1957	30	7.869	+ 0.396	
Jun. 2	7.912	.396	2							
12	7.916	.396	1			1959	Jan. 12	7.942	.416	
14	7.911	.410	2				24	7.925	.414	
16	7.881	.412	2			Feb. 4	7.894	.429	1	
17	7.913	.398	1				5	7.902	.407	
							28	7.890	.409	
<b>1955</b>										
Apr. 21	7.864	.406	3	7.865	+ 0.402	Mar. 1	7.872	.402	2	
25	7.877	.391	2				3	7.854	.402	
May 4	7.899	.392	2	7.902	.395	1960	Jan. 28	7.910	.422	
Jun. 20	7.933	.379	2			Feb. 21	7.877	.424	1	
<b>1956</b>										
Jan. 17	7.944	.398	2	7.952	.395		23	7.860	.436	
19	7.935	.392	2			Mar. 15	7.845	.438	2	
20	7.922	.407	2				17	7.845	.429	
Feb. 3	7.921	.411	2	7.919	.413		18	7.852	.431	
10	7.911	.389	2	7.902	.388		19	7.854	.437	
Mar. 19	7.896	.401	4	7.899	.393		21	7.854	.468	
Apr. 5	7.853	.427	2	7.846	.429		30	7.846	.432	
	18	7.884	.381	0	7.880	.390	<b>1960</b>			
<b>1957</b>										
Mar. 23	7.869	.415	2	7.868	.417	Apr. 3	7.854	.442	1	
24	7.862	.410	2	7.867	.404	1961	Mar. 1	7.897	.450	
28	7.928	.393	2				8	7.875	.428	
Apr. 5	7.870	.402	2				23	7.880	.419	
6	7.852	.393	4				31	7.856	.422	
May 1	7.844	.380	1	7.833	.376		Apr. 8	7.831	.421	
	3	7.859	+ 0.412	2	7.862	+ 0.411		12	7.849	+ 0.424
										2
										7.846
										+ 0.423

ponding to the amount of light coming through the two halves of the blue filter. We can then compute the planetary gradient in the blue filter region compared with stars observed in the same spectral region. This procedure gives considerably redder values of B-V than the standard U B V filters. The new values lead to more accordant results especially for large hour angles.\*

The new values of color-indices obtained by Johnson are

$$(B-V)' = + 0.721 \text{ for Uranus, and} \quad (28)$$

$$(B-V)' = + 0.628 \text{ for Neptune;}^* \quad (29)$$

they are called henceforth the gradient color indices and are used in the transformation of the blue magnitudes in all subsequent reductions. Similar gradient color indices for transforming yellow magnitudes have not yet been determined with satisfactory pre-

cision and therefore these magnitudes cannot be properly transformed to the BV system. The yellow magnitudes given in Table VIII were corrected for color dependence of extinction and reduced to the BV system using the directly observed color-indices given in the same table. The discussion of data contained in Table VIII will therefore be postponed to the further papers of this series.

*B. Observations of Planets in Blue Color.* The brightness of Uranus and Neptune in blue color is

\*Special observations of Uranus and Neptune at the low altitudes of 20° to 25° were made with the blue filter on four nights in 1961 for determining what values of color-indices would be necessary for correcting the observations for color-dependence of extinction. The color indices  $(B-V)' = + 0.82 \pm 0.11$  (m.e.) for Uranus and  $(B-V)' = + 0.50 \pm 0.05$  (m.e.) for Neptune were obtained. Taking into account the low accuracy of these values they are in satisfactory agreement with Johnson's gradient color-indices.

compared with that of comparison stars at least once a week over a period of several months around each opposition. The blue filter is the same as that used for two-color observations. The comparison of each planet with two comparison stars takes an average of about 80 minutes; the duration of such observations should not exceed 1½ hours.

All observations are made with a diaphragm 48 seconds of arc in diameter (2mm at the cassegrainian focus of the 21-inch telescope). Therefore the satellites Miranda, Ariel, Umbriel and Triton are always measured together with their respective planets. All the photometric data for planets given in this paper refer to planets measured together with the aforementioned satellites. Uranus' satellites Titania and Oberon, which are occasionally within the diaphragm together, may increase the planet's brightness by about 0.0008; this fact is neglected in the present discussion.

Denoting by A the measure for comparison star A, by B that for comparison star B, by P that for planet and by S the deflection for radioactive standard source, we can represent a single observation as:

S A B A B A ... B A P A P A ... P A B A B A ... B A S  
5 x B                  10 x P                  5 x B

A total of 10 measures of the planet, 10 measures of star B, and 21 measures of star A are made. Each measure consists of two deflections separated by the deflection for sky background. In the middle of each measurement of sky background the hour angle is read on the telescope's setting circle. The standard time is read during the 11th measure of star A.

All observations of planets in blue color made from January 1953 to June 1961 are listed in Tables IX and X. The first column gives the Universal Time at the middle of the observation with an accuracy of a hundredth of a day. The second column gives the hour angle of the planet at the middle of the observation (E denotes East, W West). The 3rd, 4th and 5th columns give the amplifier's gain used when observing the planet, the comparison star A and comparison star B, respectively. The first number in each of these columns gives the setting of the 2.5 magnitude gain steps, the second number that of the 0.5 magnitude steps. They are given to permit a future rediscussion of the effects of possible non-linearity of the amplifier and errors in its calibration. The 6th column of Tables IX and X gives the number of measures of the planet, which is usually 10. The number is smaller than 10 when the observations were stopped by clouds or when the planet could be observed only at low altitude; in this latter case the increase in the number of measures would mean that the planet is observed at still lower altitudes which would only reduce the accuracy.

Since the earth, sun, and planets change their relative positions, the measured brightness is corrected to some arbitrary fixed distance. The opposition position for 1950 has been used for this purpose. The correction is determined according to the formula

$$\Delta_d = 5(\log R_o - \log R + \log \Delta_o - \log \Delta), \quad (30)$$

where R is the distance of the planet from the sun,  $\Delta$  is the distance of the planet from the earth and  $R_o$  and  $\Delta_o$  are the arbitrary fixed distances to which the corrections are made; their values are given by Hardie and Gicias (3). The distance corrections computed from equation (30) are given in the 9th column of Tables IX and X.

The 7th and 8th columns of these tables give the differential extinction correction,  $-Q_{b1} \Delta M$ , and the correction for the color dependence of extinction,  $-Q_{b2} \bar{M} \Delta(B-V)$ , respectively. The values  $Q_b^2 = -0.033$  and the mean seasonal values of  $Q_{b1}$  are used, except for the year 1961 when the use is made of the nightly values of this coefficient. The difference  $\Delta(B-V)$  is defined as

$$\Delta(B-V) = (B-V)' - (B-V) \text{ star A}, \quad (31)$$

where for the gradient color index  $(B-V)'$  the values given by equations (28) and (29) are used. The difference in air mass is computed from the approximate formula

$$\Delta M = M_{\text{planet}} - M_{\text{star A}} \approx,$$

$$\approx (A \sin \bar{H} + B \cos \bar{H} + C) \bar{M}^2, \quad (32)$$

where  $\bar{H}$  is the mean hour angle of the planet and star A, positive westward,  $\bar{M}$  is the mean air mass of the planet and star A, while

$$A = \sin(\alpha_{\text{star A}} - \alpha_{\text{planet}}) \cos \phi \cos \bar{\delta}, \quad (33)$$

$$B = -\sin(\delta_{\text{star A}} - \delta_{\text{planet}}) \cos \phi \sin \bar{\delta}, \quad (34)$$

$$C = \sin(\delta_{\text{star A}} - \delta_{\text{planet}}) \sin \phi \cos \bar{\delta}; \quad (35)$$

here  $\alpha$  and  $\delta$  are right ascension and declination,  $\bar{\delta}$  is the mean declination of the planet and star A and  $\phi$  is geographic latitude of the observatory.

The corrections given in the 7th and 8th columns of Tables IX and X refer to the middle of the observation. Similar corrections are, however, computed for every measurement of the planet. The individual measures corrected for extinction are used for computing the nightly mean value of the magnitude difference between the planet and comparison











TABLE IX. Magnitude Differences in Blue Color Between Uranus Comparison Stars (Unit =  $10^{-4}$ ) (cont'd)

DATE U.T. 1	Hour Angle 2	Amplifier Gains					Number of measures of planet 6	$\Delta B(U-U_A)$ Corrected for differential extinction 7	$\Delta B(U-U_A)$ Corrected for color dependence of extinc- tion 8	Distance correction 9	$\Delta B(UB-U_A)$ corrected for extinction and distance 10	$\Delta B(UB-U_A)$ corrected for extinction 11	Color correction to the BV system		Phase Correction 14	$\Delta B(U-U_A)$ Final Value 16	$\Delta B(U-U_B)$ Final Value 17	Remarks 18
		Uranus 3	Star U.A. 4	Star U.B. 5	U.UA 12	U.UB 13						U-U.A 12	U-U.B 13					
1956																		
Dec.	12.38	1.5E	45	53	53	10	-2	+2	+116	-1151	+2	+113	+2	+1	-4	-2	-56	-1208 -1326
	13.39	1.2E	45	53	53	5	-2	+2	+118	-1156	+2	+109	+3	+1	-4	-1	-56	-1212 -1326
	14.39	1.4E	45	53	53	10	-2	+2	+119	-1150	+1	+109	+2	+1	-4	-1	-56	-1206 -1320
	20.37	1.2E	45	53	53	10	-2	+2	+128	-1153	+2	+112	+2	+1	-4	-1	-55	-1208 -1325
	21.40	0.3E	45	53	53	10	-2	+2	+129	-1154	+1	+111	+2	+1	-4	-1	-55	-1209 -1325
	29.43	0.9W	45	53	53	10	0	+1	+138	-1154	+1	+108	+2	+1	-4	-1	-55	-1209 -1322
1957																		
Jan.	16.36	1.0W	45	53	53	4	0	+1	+151	-1159	+2	+110	+3	+1	-4	0	-54	-1212 -1327
	17.31	0.9E	45	52	52	10	0	+1	+152	-1161	+1	+107	+2	+1	-4	0	-54	-1214 -1326
	18.34	0.1W	45	53	53	10	0	+1	+152	-1151	+2	+110	+2	+1	-4	0	-54	<u>-1204 -1319</u>
																mean	-1210 -1324	
Feb.	1.40	2.4W	45	53	53	10	-3	+2	+147	-1155	+2	+112	+3	+1	-4	0	-53	-1207 -1324
	14.30	1.0W	44	52	52	10	-1	+1	+147	-1155	+2	+106	+3	+1	-4	0	-52	-1206 -1317
	28.33	2.6W	45	53	53	3	-7	+2	+133	-1154	+2	+109	+4	+1	-4	-1	-51	-1205 -1319
Mar.	27.14	0.2E	45	53	53	7	0	+1	+93	-1160	+2	+102	+3	+1	-4	-2	-50	-1211
Apr.	5.24	2.9W	45	53	53	10	-11	+2	+75	-1155	+2	+103	+3	+1	-4	-3	-50	-1207
	15.14	1.4W	45	53	53	10	-5	+2	+57	-1156	+2	+103	+3	+1	-4	-3	-50	-1208 -1316
	28.18	2.8W	45	53	53	10	-12	+2	+31	-1154	+4	+105	+4	+1	-4	-3	-50	-1206 -1316
May	2.18	3.2W	45	53	53	10	-13	+2	+24	-1143	+4	+102	+4	+1	-4	-3	-50	<u>-1199</u>
																mean	-1207 -1319	
Sep.	21.50	4.2E	46	54	54	9	-8	-9	-	-	-	-198	+4	-2	0	-2	-62	
	22.49	4.4E	46	54	54	10	-9	-9	-32	-1468	+4	-198	+2	-2	0	-2	-62	-1534
	28.48	4.2E	46	54	54	10	-13	-9	-23	-1455	+4	-198	+2	-2	0	-2	-63	-1522 -1322
	29.47	4.0E	46	54	54	10	-12	-8	-21	-1455	+4	-196	+2	-2	0	-2	-63	-1522 -1324
Oct.	9.48	3.6E	46	54	53	10	-12	-8	-4	-1465	+4	-198	+1	-2	0	-2	-64	-1533 -1333
	10.47	3.7E	46	54	54	10	-13	-8	-2	-1456	+4	-197	+2	-2	0	-2	-64	-1524 -1325
	23.47	3.0E	45	53	53	4	-9	-6	+23	-1460	+3	-201	+3	-2	0	-3	-64	-1529 -1326
	24.47	3.0E	45	53	53	8	-9	-6	+25	-1464	+3	-202	+2	-2	0	-3	-64	-1533 -1329
Nov.	7.49	1.4E	45	53	53	10	-4	-5	+53	-1468	+2	-194	+1	-2	0	-3	-65	-1538 -1342
	8.45	2.4E	45	53	53	10	-7	-6	+55	-1466	+3	-197	+1	-2	0	-3	-65	-1536 -1337
	10.43	2.8E	45	53	53	9	-9	-6	+59	-1468	+3	-196	+2	-2	0	-3	-65	-1538 -1340
	17.46	1.6E	45	53	53	10	-5	-5	+74	-1467	+2	-194	+1	-2	0	-3	-65	-1537 -1341
	21.39	3.1E	45	53	53	10	-11	-7	+80	-1470	+4	-198	+2	-2	0	-3	-65	-1540 -1340
	24.50	0.2E	45	53	53	10	-1	-5	+88	-1461	+1	-200	+1	-2	0	-3	-65	-1531 -1329
	25.39	2.6E	45	53	53	10	-8	-6	+89	-1469	+3	-196	+1	-2	0	-2	-65	-1539 -1340
	27.43	1.7E	45	53	53	10	-4	-5	+93	-1464	+2	-196	+1	-2	0	-2	-65	-1533 -1335
Dec.	3.39	2.3E	45	53	53	10	-7	-6	+105	-1463	+3	-197	+1	-2	0	-2	-65	-1532 -1333
	4.39	2.1E	45	53	53	10	-6	-6	+106	-1465	+2	-194	+1	-2	0	-2	-65	-1534 -1338
	10.38	1.6E	45	53	53	10	-4	-5	+117	-1468	+2	-198	+1	-2	0	-2	-64	-1536 -1336
	12.39	1.5E	45	53	53	10	-4	-5	+120	-1466	+2	-199	+1	-2	0	-2	-64	<u>-1534 -1333</u>
																mean	-1533 -1334	
1958																		
Jan.	8.39	0.2W	43	51	51	10	-5	0	+155	-1474	+1	-197	+1	-2	0	0	-63	-1539 -1340
	9.35	0.7E	43	51	51	10	-5	-1	+156	-1471	+1	-197	+1	-2	0	0	-63	-1536 -1337
	10.31	1.6E	43	51	51	10	-6	-1	+157	-1471	+2	-199	+1	-2	0	0	-63	-1536 -1335
	12.44	1.9W	43	51	51	10	-2	-1	+158	-1470	+1	-200	+1	-2	0	0	-62	-1534 -1332
	14.45	2.0W	43	51	51	10	-2	-1	+160	-1472	+1	-198	+1	-2	0	0	-62	-1536 -1336
	16.41	1.2W	43	51	51	10	+1	-5	+161	-1471	+1	-198	+1	-2	0	0	-62	-1535 -1335
	20.45	2.4W	43	51	51	10	+2	-6	+163	-1474	+3	-198	+1	-2	0	0	-62	-1538 -1338
	23.44	2.4W	43	51	51	11	+2	-6	+164	-1476	+2	-200	+2	-2	0	0	-62	-1540 -1338
	28.41	2.0W	43	51	51	10	+2	-6	+165	-1470	+3	-197	+1	-2	0	0	-61	-1533 -1334
Feb.	10.37	2.0W	44	52	52	9	0	-6	+163	-1468	+4	-200	+3	-2	0	0	-60	-1530 -1328
	11.22	1.6E	44	52	52	10	0	-5	+163	-1468	+1	-196	+2	-2	0	0	-60	-1530 -1332
	12.21	1.7E	44	52	52	10	+1	-5	+163	-1467	+1	-196	+1	-2	0	0	-60	-1529 -1331
	13.20	1.5E	44	52	52	10	+1	-5	+162	-1468	+1	-197	+2	-2	0	0	-60	-1530 -1331
																	clouds	

TABLE IX. Magnitude Differences in Blue Color between Uranus Comparison Stars (Unit =  $0.^m.001$ ) (cont'd)

DATE U.T.	Hour Angle	Amplifier Gains					Number of measures of planet	$\Delta B'(UB-UA)$ corrected for extinction and distance	$\Delta B'(UB-UA)$ corrected for extinction	$\Delta B'(UB-UA)$ Final Value	Remarks	
		Uranus		Star UA	Star UB							
		3	4	5	6	7						
1958												
Feb. 16.19	2.0E	44	52	52	10	- 2	- 1	+160	-1465	+ 2	- 195	+ 2 - 2
18.24	0.5E	44	52	52	10	- 3	0	+159	-1467	+ 1	- 197	+ 1 - 2
27.18	1.5E	44	52	52	10	+ 3	- 5	+151	-1469	+ 1	- 199	+ 1 - 2
Mar. 2.20	0.7E	44	52	52	10	+ 2	- 5	+148	-1469	+ 1	- 199	+ 2 - 2
5.28	0.8W	44	52	52	10	0	- 5	+145	-1464	+ 2	- 195	+ 2 - 2
10.15	1.4E	45	53	53	10	- 1	- 1	+139	-1468	+ 2	- 194	+ 2 - 2
18.20	0.0	44	52	52	10	+ 1	- 5	+127	-1465	+ 1	- 196	+ 1 - 2
19.15	1.3E	44	52	52	10	+ 4	- 5	+126	-1466	+ 1	- 195	+ 1 - 2
20.12	1.3E	44	52	52	10	+ 4	- 5	+124	-1462	+ 2	- 195	+ 2 - 2
27.14	0.6E	45	53	53	10	+ 2	- 5	+113	-1466	+ 1	- 198	+ 1 - 2
Apr. 6.16	0.6W	45	53	53	10	0	- 5	+ 94	-1468	+ 1	- 195	+ 1 - 2
9.14	0.5W	45	53	53	10	0	- 5	+ 89	-1465	+ 2	- 195	+ 1 - 2
10.15	0.6W	44	52	52	10	0	- 5	+ 88	-1464	+ 1	- 197	+ 1 - 2
11.18	1.3W	44	52	52	10	- 1	- 5	+ 86	-1465	+ 2	- 198	+ 1 - 2
14.15	0.7W	45	53	52	10	0	- 5	+ 80	-1466	+ 3	- 194	+ 1 - 2
15.19	2.1W	45	53	53	10	- 3	- 6	+ 78	-1465	+ 2	- 202	+ 2 - 2
23.15	1.4W	45	53	53	10	- 7	- 1	+ 62	-1465	+ 2	- 195	+ 1 - 2
24.14	1.4W	45	53	53	10	- 7	- 1	+ 60	-1465	+ 2	- 197	+ 1 - 2
28.16	2.1W	45	53	53	10	- 4	- 6	+ 53	-1463	+ 3	- 197	+ 1 - 2
May 4.19	3.1W	45	53	52	10	- 7	- 7	+ 41	-1466	+ 3	- 201	+ 2 - 2
9.16	2.9W	45	52	52	10	- 6	- 6	+ 31	-1463	+ 3	- 198	+ 3 - 2
10.18	3.2W	45	53	53	10	- 7	- 7	+ 29	-1465	+ 3	- 197	+ 1 - 2
17.20	3.6W	45	53	53	10	- 8	- 8	+ 16	-1469	+ 3	- 198	+ 2 - 2
23.17	3.8W	45	53	53	10	- 10	- 8	+ 6	-1459	+ 4	- 201	+ 2 - 2
25.19	4.0W	45	52	52	10	- 11	- 8	+ 1	-1472	+ 4	- 197	+ 2 - 2
Jun. 1.19	4.4W	45	52	52	8	- 11	- 9	- 10	-1451	+ 7	- 192	+ 4 - 2
mean												
1959												
Oct. 3.48	4.3E	44	46		10	+ 8	- 1	- 12	-1121	+ 3		
4.50	4.0E	44	46		9	+ 6	- 1	- 10	-1130	+ 3		
7.49	4.0E	43	45		10	+ 4	- 1	- 5	-1122	+ 2		
8.47	4.1E	43	45		10	+ 3	- 1	- 4	-1122	+ 2		
14.47	3.3E	43	45		10	+ 2	- 1	+ 7	-1122	+ 1		
15.43	4.3E	44	46		10	+ 2	- 1	+ 9	-1126	+ 1		
21.48	3.0E	44	46	54	9	- 2	- 1	+ 20	-1120	+ 1	+1522	+ 2 2
22.45	3.0E	44	46	54	10	- 2	- 1	+ 22	-1127	+ 2	+1521	+ 1 2
Nov. 2.45	3.1E	44	46	54	10	- 2	- 1	+ 44	-1130	+ 2	+1525	+ 1 2
3.47	2.4E	44	46	54	10	- 1	- 1	+ 46	-1124	+ 1	+1523	+ 1 2
7.48	2.1E	44	46	54	10	- 1	- 1	+ 54	-1127	+ 1	+1519	+ 1 2
8.49	1.8E	44	46	54	10	0	- 1	+ 56	-1128	+ 1	+1519	+ 1 2
13.50	1.1E	44	46	54	10	+ 1	- 1	+ 66	-1129	+ 1	+1519	+ 1 2
14.46	2.2E	44	52	55	10	- 1	- 1	+ 68	-1118	+ 1	+1527	+ 1 3
20.43	2.3E	44	52	55	10	- 1	- 1	+ 81	-1114	+ 12	+1531	+ 3 3
23.48	1.0E	44	46	54	10	+ 1	- 1	+ 87	-1127	+ 1	+1524	+ 2 2
29.52	0.3W	44	46	54	10	+ 3	- 1	+ 99	-1120	+ 2	+1525	+ 1 2
30.40	2.4E	44	46	54	10	- 1	- 1	+100	-1119	+ 2	+1525	+ 1 3
Dec. 10.44	0.9E	44	52	55	10	+ 1	- 1	+119	-1115	+ 3	+1528	+ 1 2
11.40	1.7E	44	52	55	10	0	- 1	+121	-1116	+ 1	+1530	+ 1 2
15.41	1.5E	43	45	53	10	0	- 1	+128	-1113	+ 1	+1527	+ 1 2
17.34	2.0E	43	45	53	10	+ 1	- 1	+131	-1120	+ 1	+1526	+ 1 2
19.46	0.2W	43	45	53	10	+ 3	- 1	+140	-1115	+ 1	+1521	+ 1 2
23.37	1.4E	42	45	53	10	+ 1	- 1	+141	-1123	+ 1	+1526	+ 1 2
29.33	2.3E	42	44	54	10	+ 1	- 1	+149	-1122	+ 1		
31.37	1.2E	42	44	52	10	+ 2	- 1	+152	-1123	+ 1	+1522	+ 3
mean												
Jan. 5.36	1.1E	42	45	53	10	+ 2	- 1	+158	-1118	+ 2	+1525	+ 2
11.35	0.9E	43	51	54	10	+ 2	- 1	+164	-1122	+ 1	+1525	+ 2
18.40	0.7W	43	51	54	10	+ 3	- 1	+169	-1125	+ 1	+1525	+ 2
31.36	0.8W	43	51	54	10	+ 3	- 1	+176	-1124	+ 2	-1518	+ 2
mean												
clouds												
bad seeing												
windy												

TABLE IX. Magnitude Differences in Blue Color between Uranus Comparison Stars (Unit =  $0.^m001$ ) (cont'd)

DATE U.T. 1	Hour Angle 2	Amplifier Gains					Number of measures of planet 6	Correction for dif- ferential extinction 7	Distance correction 8	$\Delta B'(U-UA)$ corrected for extinction and distance 9	$\Delta B'(UB-UA)$ corrected for extinction 10	Color correction to the RV system		Phase Correction 14	Oblateness correction 15	$\Delta B(U-UA)$ Final Value 16	$\Delta B(U-UB)$ Final Value 17	Remarks 18	
		Uranus 3	Star UA 4	Star UB 5								U-UA 11	U-UA 12	U-UB 13					
1959																			
Feb.	4.27	1.1E	43	51	54	10	+ 4	- 1	+176	-1124	+ 1	+1522	+ 2	0	0	-68	-1192 -2714		
	5.25	1.5E	43	51	54	10	+ 4	- 1	+176	-1124	+ 1	+1522	+ 2	0	0	-68	-1192 -2714		
	11.36	1.6W	44	51	54	10	+ 3	- 1	+175	-1126	+ 2	+1511	+ 2	0	0	-68	-1194 -2705	hazy?	
	14.38	2.0W	44	41	54	10	+ 3	- 1	+174	-1125	+ 2	+1518	+ 3	0	0	-67	-1192 -2710		
	19.37	2.2W	44	51	54	10	+ 2	- 1	+172	-1125	+ 2	+1513	+ 3	0	0	-67	-1192 -2705		
	24.44	4.2W	44	52		9	+ 4	- 1	+169	-1125	+ 4			0	0	-67	-1192	clouds	
	25.36	2.3W	44	52		10	+ 2	- 1	+168	-1119	+ 1			0	0	-67	-1186	clouds	
Mar.	3.34	2.4W	45	52	55	10	+ 1	- 1	+163	-1124	+ 1	+1512	+ 3	0	0	-1	-67	-1192 -2704	
	7.41	3.4W	45	52	55	10	- 1	- 1	+159	-1124	+ 1	+1512	+ 3	0	0	-1	-67	-1192 -2704	
	13.41	4.6W	45	53		10	- 4	- 1	+152	-1123	+ 2			0	0	-1	-66	-1190	
	15.35	3.3W	45	52	55	10	- 2	- 1	+149	-1127	+ 1	+1509	+ 3	0	0	-1	-66	-1194 -2703	
	19.33	3.2W	45	53		10	- 2	- 1	+144	-1124	+ 2			0	0	-1	-66	-1191	
	26.31	3.0W	45	52	55	10	- 2	- 1	+133	-1115	+ 1	+1523	+ 3	0	0	-2	-66	-1183 -2706	
Apr.	9.25	2.7W	45	52	54	10	- 2	- 1	+109	-1126	+ 2	+1517	+ 4	0	+ 1	- 2	-66	-1194 -2711	
	10.25	2.7W	45	52	54	10	- 2	- 1	+107	-1120	+ 1	+1518	+ 2	0	+ 1	- 2	-66	<u>-1188 -2706</u>	
														mean	-1190	-2707			
1960																			
Jan.	28.34	0.2E	43	54	54	10	+ 1	+ 9	+183	-2640	+ 2	+ 252	+ 1	- 5	- 2	0	-77	-2722 -2971	
	29.42	0.1E	43	54	54	3	+ 1	+ 9	+183	-2635	+ 2	+ 254	+ 3	- 5	- 2	0	-77	-2717 -2968	clouds, wind
Feb.	17.33	0.9W	43	54	54	10	0	+ 9	+185	-2637	+ 2	+ 254	+ 2	- 5	- 2	0	-75	-2717 -2968	
	18.28	0.4W	43	54	54	7	0	+ 9	+184	-2637	+ 2	+ 257	+ 2	- 5	- 2	0	-75	-2717 -2971	
	21.32	0.8W	43	54	54	10	0	+ 9	+183	-2639	+ 2	+ 253	+ 2	- 5	- 2	0	-74	-2718 -2968	p seeing
	22.31	0.8W	43	54	54	10	0	+ 9	+183	-2636	+ 2	+ 258	+ 2	- 5	- 2	0	-74	-2715 -2970	clds at end
	23.34	1.4W	43	54	54	6	- 1	+ 9	+183	-2637	+ 2	+ 256	+ 2	- 5	- 2	0	-74	-2716 -2969	
	25.35	2.0W	43	54	54	6	- 2	+ 9	+181	-2636	+ 2	+ 255	+ 3	- 5	- 2	0	-74	-2715 -2967	
Mar.	3.34	2.0W	43	54	54	6	- 3	+10	+176	-2637	+ 2	+ 251	+ 3	- 5	- 2	0	-74	-2715 -2964	
	15.26	1.1W	43	54	54	10	0	+ 9	+164	-2642	+ 2	+ 255	+ 2	- 5	- 2	- 1	-73	-2721 -2973	
	17.27	1.5W	43	54	54	7	- 2	+ 9	+162	-2638	+ 2	+ 251	+ 2	- 5	- 2	- 1	-73	-2717 -2965	
	19.28	1.8W	43	54	54	6	- 3	+ 9	+159	-2638	+ 2	+ 254	+ 3	- 5	- 2	- 1	-73	-2717 -2968	
	21.28	1.9W	44	55	54	6	- 4	+ 9	+156	-2636	+ 2	+ 255	+ 2	- 5	- 2	- 1	-73	-2715 -2967	
Apr.	1.26	2.0W	44	54	54	6	- 5	+10	+141	-2638	+ 2	+ 252	+ 2	- 2	- 1	- 2	-72	-2714 -2965	clouds
	3.16	0.2E	44	54	55	6	+ 2	+ 9	+136	-2639	+ 3	+ 253	+ 2	- 2	- 1	- 2	-72	-2715 -2967	
	4.20	0.6W	44	54	55	10	0	+ 9	+135	-2635	+ 2	+ 252	+ 1	- 2	- 1	- 2	-72	-2711 -2962	
	6.27	1.4W	44	54	55	6	- 2	+ 9	+131	-2638	+ 2	+ 258	+ 3	- 2	- 1	- 2	-72	-2714 -2971	
	21.22	2.4W	44	55	55	6	- 8	+10	+104	-2643	+ 4	+ 250	+ 2	- 2	- 1	- 3	-72	-2720 -2969	
May	7.18	2.6W	44	54	54	10	- 8	+10	+ 72	-2640	+ 4	+ 251	+ 2	- 2	- 1	- 3	-72	-2717 -2967	
	12.18	2.8W	44	54	54	7	- 11	+ 11	+ 63	-2641	+ 4	+ 250	+ 3	- 2	- 1	- 3	-72	-2718 -2967	
	18.18	3.3W	44	54	54	7	- 16	+ 12	+ 51	-2638	+ 6	+ 251	+ 2	- 2	- 1	- 3	-72	-2715 -2965	
	21.19	3.6W	44	54	54	6	- 19	+ 13	+ 46	-2643	+ 6	+ 253	+ 3	- 2	- 1	- 3	-73	-2721 -2973	
	23.18	3.8W	44	54	54	6	- 19	+ 13	+ 42	-2640	+ 6	+ 248	+ 3	- 2	- 1	- 3	-73	-2718 -2965	
	25.18	3.7W	44	54	54	6	- 19	+ 13	+ 38	-2638	+ 6	+ 256	+ 2	- 2	- 1	- 3	-73	-2716 -2971	
	28.18	4.0W	44	54	54	6	- 24	+ 15	+ 32	-2641	+ 8	+ 246	+ 5	- 2	- 1	- 3	-73	<u>-2719 -2964</u>	
														mean	-2717	-2968			
Oct.	20.48	3.3E	51	53	53	8	-15	+13	+ 20	-1316	+ 6	+ 62	+ 8	- 5	- 5	- 2	-85	-1408 -1470	clouds
Nov.	17.51	0.9E	45	53	53	10	- 2	+ 9	+ 75	-1313	+ 2	+ 68	+ 2	- 5	- 5	- 3	-85	-1406 -1474	
	18.50	1.3E	45	53	53	10	- 2	+ 9	+ 77	-1314	+ 2	+ 66	+ 2	- 5	- 5	- 3	-85	-1407 -1473	
	22.51	0.8E	45	53	53	10	- 2	+ 9	+ 85	-1314	+ 2	+ 67	+ 2	- 5	- 5	- 3	-85	-1407 -1474	
	26.47	1.5E	45	53	53	10	- 6	+ 9	+ 93	-1316	+ 3	+ 66	+ 3	- 5	- 5	- 3	-86	-1410 -1476	
Dec.	15.51	0.9W	45	53	53	10	+ 4	+ 9	+131	-1313	+ 3	+ 68	+ 1	- 5	- 5	- 2	-85	-1405 -1473	
	16.51	0.7W	45	53	53	10	+ 4	+ 9	+133	-1316	+ 3	+ 66	+ 1	- 5	- 5	- 2	-85	-1408 -1474	
	31.43	0.3E	45	53	53	10	+ 1	+ 9	+158	-1316	+ 2	+ 66	+ 3	- 5	- 5	- 2	-85	-1408 -1474	
1961																			
Jan.	3.48	1.4W	46	54	54	10	+ 2	+ 9	+163	-1315	+ 3	+ 66	+ 1	- 5	- 5	- 1	-85	-1406 -1472	p seeing
	4.44	0.6W	45	53	53	10	+ 1	+ 9	+164	-1316	+ 2	+ 69	+ 1	- 5	- 5	- 1	-85	-1407 -1476	clouds
	10.38	0.8E	46	54	54	10	+ 1	+ 9	+172	-1319	+ 4	+ 68	+ 2	- 5	- 5	- 1	-85	-1410 -1478	
	11.43	0.6W	46	54	54	10	+ 3	+ 9	+173	-1320	+ 6	+ 68	+ 1	- 5	- 5	- 1	-85	-1411 -1479	
	29.36	0.0	45	53	53	10	+ 2	+ 9	+190	-1320	+ 3	+ 68	+ 1	- 5	- 5	0	-84	-1409 -1477	
Feb.	6.33	0.2W	45	53	53	10	+ 2	+ 9	+193	-1323	+ 3	+ 68	+ 1	- 5	- 5	0	-83	-1411 -1479	
	8.36	0.5W	45	53	53	7	+ 2	+ 9	+194	-1322	+ 2	+ 66	+ 1	- 5	- 5	0	-83	-1410 -1476	

TABLE IX. Magnitude Differences in Blue Color between Uranus Comparison Stars (Unit = 0.001) (cont'd)

DATE U.T.	Hour Angle	Amplifier Gains					Number of measures of planet	Correction for dif- ferential extinction U-UA	Correction for color dependence of ex- tinction	Distance correction	$\Delta B'(UB-UA)$ corrected for extinction and distance	$\Delta B'(UB-UA)$ corrected for extinction	U-UA	Color correction to the BV system	Phase Correction	Oblateness correction	$\Delta B'(UA)$ Final Value	$\Delta B'(UB)$ Final Value	Remarks	
		3	4	5	Star UA	Star UB														
		Uranus																		
1961																				
Feb.	15.35	0.8W	45	53	53	8	+ 4	+ 9	+195	-1321	+ 3	+ 70	+ 1	- 5	- 5	0	-83	-1409	-1479	
	21.36	1.4W	45	53	53	7	+ 4	+ 9	+193	-1319	+ 3	+ 69	+ 1	- 5	- 5	0	-82	-1406	-1475	
	22.30	0.3W	46	54	54	10	+ 4	+ 9	+193	-1321	+ 3	+ 68	+ 1	- 5	- 5	0	-82	-1408	-1476	
Mar.	8.28	0.6W	45	53	53	8	+ 3	+ 9	+185	-1318	+ 2	+ 68	+ 1	- 5	- 5	0	-82	-1406	-1474	
	23.20	0.3E	46	54	54	10	+ 0	+ 9	+169	-1320	+ 2	+ 70	+ 1	- 5	- 5	- 1	-81	-1407	-1477	
	31.17	0.4E	46	54	54	10	+ 3	+ 9	+158	-1318	+ 2	+ 70	+ 1	- 5	- 5	- 2	-81	-1406	-1476	
Apr.	4.17	0.3E	46	53	53	10	+ 3	+ 9	+151	-1319	+ 3	+ 68	+ 1	- 5	- 5	- 2	-81	-1407	-1475	
	9.19	0.5W	46	54	54	10	+ 3	+ 9	+143	-1323	+ 3	+ 67	+ 1	- 5	- 5	- 2	-80	-1410	-1477	
May	12.18	2.5W	46	54	54	8	+ 4	+11	+ 81	-1312	+ 3	+ 69	+ 1	- 5	- 5	- 3	-80	-1400	-1469	
	19.16	3.2W	46	54	54	6	+ 6	+13	+ 68	-1314	+ 4	+ 67	+ 2	- 5	- 5	- 3	-80	-1399	-1467	clouds
	21.18	3.3W	46	54	54	10	+ 6	+13	+ 64	-1311	+ 4	+ 68	+ 2	- 5	- 5	- 3	-80	-1399	-1467	
																mean	-1407	-1475		

Lall. = Lallemand tube, p seeing = Poor seeing, 2 = Amplifier No. 2, 1 = Amplifier No. 1, 4 = Amplifier No. 4,  
 3C = Amplifier No. 3C, \* = Wind, \*\* = Clouds, \*\*\* = 3mm diaphragm.

TABLE X. Magnitude Differences in Blue Color Between Neptune and Comparison Stars (Unit =  $10^{-4}$ )

DATE UT.	Hour Angle	Amplifier Gains					Number of measures of planet	Correction for dif- ferential extinction N-NA	Correction for color dependence of ex- tinction N-NA	Distance correction 9	$\Delta B'(N-NA)$ corrected for extinction and distance 10	$\Delta B'(NB-NA)$ corrected for extinction 11	N-NA	Color correction to the BV system		$\Delta B(N-NA)$ Final Value 15	$\Delta B(N-NB)$ Final Value 16	Remarks 18		
		Neptune 3	Star NA 4	Star NB 5											N-NA	N-NB				
1953																				
Mar.	6.43	0.5W					9	+ 3	+ 7	- 51	- 523	+ 2	+ 7	- 1	- 517	42 <sup>10</sup> tel.				
	7.42	0.2W					10	+ 3	+ 7	- 50	- 518	+ 1 2	+ 7	- 1	- 512	42 <sup>10</sup> tel.				
	12.42	0.5W					10	+ 3	+ 7	- 45	- 518	+ 1 2	+ 7	- 1	- 512	42 <sup>10</sup> tel.				
Jun.	10.20	1.3W	45	45			6	+ 2	+ 8	- 68	- 513	+ 2	+ 4	- 2	- 511	21 <sup>10</sup> tel.				
	13.19	1.0W	45	46			8	+ 3	+ 8	- 71	- 524	+ 2	+ 4	- 2	- 522	21 <sup>10</sup> tel.*				
	15.18	1.0W	45	45			10	+ 3	+ 8	- 73	- 518	+ 2	+ 4	- 2	- 516	21 <sup>10</sup> tel.				
	16.18	1.1W	45	45			10	+ 3	+ 8	- 74	- 519	+ 2	+ 4	- 2	- 517	21 <sup>10</sup> tel.				
	17.19	1.5W	45	45			10	+ 2	+ 8	- 75	- 520	+ 2	+ 4	- 2	- 518	21 <sup>10</sup> tel.				
	18.20	1.6W	45	45			10	+ 2	+ 8	- 76	- 517	+ 2	+ 4	- 2	- 515	21 <sup>10</sup> tel.				
	19.18	1.4W	45	45			10	+ 2	+ 8	- 77	- 517	+ 2	+ 4	- 2	- 515	21 <sup>10</sup> tel.				
	22.19	1.6W	44	45			10	+ 2	+ 8	- 81	- 518	+ 2	+ 4	- 2	- 516	21 <sup>10</sup> tel.				
	24.19	1.8W	44	45			10	+ 2	+ 8	- 83	- 519	+ 2	+ 4	- 2	- 517	21 <sup>10</sup> tel.				
	25.19	1.8W	44	45			10	+ 2	+ 8	- 84	- 516	+ 2	+ 4	- 2	- 514	21 <sup>10</sup> tel.				
	27.19	2.0W	44	45			10	+ 2	+ 8	- 86	- 518	+ 2	+ 4	- 2	- 516	21 <sup>10</sup> tel.				
	28.17	1.7W	44	45			10	+ 2	+ 8	- 88	- 517	+ 2	+ 4	- 2	- 515	21 <sup>10</sup> tel.				
Jul.	1.18	2.0W	45	45			10	+ 2	+ 8	- 91	- 512	+ 2	+ 4	- 2	- 510	21 <sup>10</sup> tel.				
	2.18	2.1W	45	45			10	+ 2	+ 8	- 92	- 518	+ 2	+ 4	- 2	- 516	21 <sup>10</sup> tel.				
	3.17	2.0W	45	45			10	+ 2	+ 8	- 93	- 522	+ 2	+ 4	- 2	- 520	21 <sup>10</sup> tel.				
														mean	- 516					
1954																				
Feb.	4.49	0.5E					10	- 2	+ 5	- 84	- 785	+ 2	+ 600	+ 3	+ 1	+ 3	- 2	- 786	- 1384 R	
	5.51	0.1E					10	- 1	+ 5	- 83	- 784	+ 2	+ 597	+ 2	+ 1	+ 3	- 2	- 785	- 1380 R	
	6.49	0.3E					10	- 2	+ 5	- 82	- 784	+ 2	+ 601	+ 3	+ 1	+ 3	- 2	- 785	- 1384 R	
	7.52	0.5W					10	- 1	+ 5	- 80	- 794	+ 3	+ 584	+ 2	+ 1	+ 3	- 2	- 795	- 1377 U	
	10.52	0.8W					5	0	+ 5	- 77	- 798	+ 4	+ 588	+ 5	+ 1	+ 3	- 2	- 799	- 1385 U	
	11.45	0.8E					10	- 3	+ 5	- 76	- 788	+ 2	+ 600	+ 3	+ 1	+ 3	- 2	- 787	- 1385	
	17.50	0.8W					10	0	+ 5	- 69	- 785	+ 2	+ 600	+ 5	+ 1	+ 3	- 2	- 786	- 1384 R	
	18.50	0.7W					3	0	+ 5	- 68	- 787	+ 6	+ 597	+ 6	+ 1	+ 3	- 2	- 788	- 1383 R	
	20.52	1.4W					6	0	+ 5	- 66	- 782	+ 2	+ 594	+ 3	+ 1	+ 3	- 1	- 786	- 1374 R	
	21.49	0.8W					10	0	+ 5	- 65	- 787	+ 2	+ 594	+ 6	+ 1	+ 3	- 1	- 791	- 1389	
	24.48	0.7W					10	0	+ 5	- 62	- 776	+ 2	+ 602	+ 4	+ 1	+ 3	- 1	- 776	- 1376 U	
	28.48	0.9W					10	0	+ 5	- 58	- 794	+ 2	+ 592	+ 3	+ 1	+ 3	- 1	- 794	- 1384 U	
Mar.	1.46	0.6W					10	0	+ 5	- 57	- 786	+ 2	+ 609	+ 3	+ 1	+ 3	- 1	- 786	- 1393 U	
	2.46	0.6W					10	0	+ 5	- 56	- 787	+ 3	+ 602	+ 4	+ 1	+ 3	- 1	- 787	- 1387	
	5.44	0.4W					10	0	+ 5	- 54	- 793	+ 2	+ 594	+ 3	+ 1	+ 3	- 1	- 793	- 1385 R	
	6.45	0.5W					12	0	+ 5	- 53	- 802	+ 3	+ 602	+ 1	+ 1	+ 3	0	- 801	- 1377 U	
														mean	- 788	- 1383				
Apr.	28.41	1.2W					10	0	+ 5	- 39	- 787	+ 2	+ 610	+ 2	+ 5	+ 10	0	- 797	- 1387 Lall.	
	1.37	0.4W					10	+ 1	+ 5	- 38	- 786	+ 2	+ 609	+ 2	+ 5	+ 10	0	- 781	- 1385 Lall.	
	4.37	0.7W					10	+ 1	+ 5	- 37	- 785	+ 2	+ 607	+ 2	+ 5	+ 10	0	- 780	- 1382 Lall.	
	6.35	0.2E					5	+ 1	+ 5	- 36	- 780	+ 2	+ 608	+ 5	+ 5	+ 10	0	- 775	- 1378 Lall.	
	9.38	1.2W					10	0	+ 5	- 36	- 781	+ 3	+ 616	+ 2	+ 5	+ 10	0	- 776	- 1387 Lall.	
	12.28	1.0E					6	+ 2	+ 5	- 36	- 781	+ 3	+ 611	+ 3	+ 5	+ 10	0	- 776	- 1382 Lall.	
	13.33	0.2W					5	+ 1	+ 5	- 36	- 770	+ 2	+ 614	+ 3	+ 5	+ 10	0	- 775	- 1384 Lall.	
	14.34	0.7W					5	0	+ 5	- 36	- 781	+ 2	+ 614	+ 3	+ 5	+ 10	0	- 775	- 1385 Lall.	
	16.28	0.8E					10	+ 2	+ 5	- 35	- 783	+ 3	+ 612	+ 2	+ 5	+ 10	0	- 778	- 1385 Lall.	
	17.31	0.0					10	+ 1	+ 5	- 35	- 782	+ 2	+ 603	+ 4	+ 5	+ 10	0	- 777	- 1375 Lall.	
	22.32	0.6W					10	+ 1	+ 5	- 36	- 777	+ 1	+ 609	+ 2	+ 5	+ 10	0	- 773	- 1377 Lall.	
	23.38	2.2W					5	- 1	+ 6	- 36	- 779	+ 2	+ 617	+ 3	+ 5	+ 10	0	- 773	- 1386 Lall.	
	26.30	0.6W					5	+ 1	+ 5	- 36	- 787	+ 2	+ 612	+ 2	+ 5	+ 10	0	- 781	- 1389 Lall.	
May	2.31	1.1W					5	+ 1	+ 5	- 39	- 788	+ 2	+ 601	+ 2	+ 5	+ 10	0	- 783	- 1389 Lall.	
	3.27	0.2W					5	+ 1	+ 5	- 39	- 792	+ 3	+ 600	+ 2	+ 5	+ 10	0	- 786	- 1382 Lall.	
	4.23	0.8E					5	+ 4	+ 5	- 39	- 779	+ 2	+ 615	+ 2	+ 5	+ 10	0	- 774	- 1384 Lall.	
	5.25	0.1E					5	+ 2	+ 5	- 40	- 791	+ 1	+ 607	+ 1	+ 5	+ 10	0	- 785	- 1388 Lall.	
	6.23	0.6E					5	+ 3	+ 5	- 40	- 794	+ 1	+ 600	+ 3	+ 5	+ 10	0	- 789	- 1384 Lall.	
	8.28	0.8W					5	+ 1	+ 5	- 41	- 787	+ 2	+ 610	+ 2	+ 5	+ 10	0	- 782	- 1387 Lall.	
	10.28	0.7W					8	+ 1	+ 5	- 42	- 794	+ 2	+ 602	+ 3	+ 5	+ 10	0	- 788	- 1386 Lall.	
	11.27	0.8W					5	+ 1	+ 5	- 43	- 789	+ 2	+ 610	+ 4	+ 5	+ 10	0	- 784	- 1389 Lall.	

TABLE X. Magnitude Differences in Blue Color Between Neptune and Comparison Stars (Unit = 0.001) (cont'd)

DATE U.T.	Hour Angle	Amplifier Gains					Number of measures of planet	Correction for dif- ferential extinction	$\Delta B'(N-NA)$ corrected for extinction and distance	$\Delta B'(NB-NA)$ corrected for extinction	Phase Correction	$\Delta B(N-NA)$ Final Value	$\Delta B(N-NB)$ Final Value	Remarks				
		Neptune	Star NA	Star NB	N-NA	N-B												
May 12.24	0.2W	5	+ 2	+ 5	- 43	- 779	+ 2	+ 615	+ 3	+ 5	+ 10	0	- 774	- 1384	Lall.			
25.22	0.7W	4	+ 1	+ 5	- 52	- 789	+ 2	+ 603	+ 3	+ 5	+ 10	- 1	- 784	- 1383	Lall.			
28.24	1.3W	5	0	+ 5	- 54	- 787	+ 2	+ 609	+ 3	+ 5	+ 10	- 1	- 783	- 1387	Lall.			
31.23	1.2W	5	0	+ 5	- 57	- 788	+ 2	+ 598	+ 3	+ 5	+ 10	- 1	- 784	- 1377	Lall.			
Jun. 1.24	1.4W	5	0	+ 5	- 58	- 786	+ 2	+ 601	+ 5	+ 5	+ 10	- 1	- 782	- 1378	Lall.			
2.25	1.8W	5	- 1	+ 5	- 58	- 790	+ 3	+ 608	+ 3	+ 5	+ 10	- 1	- 786	- 1389	Lall.			
6.24	0.8W	5	+ 2	+ 5	- 62	- 784	+ 2	+ 610	+ 3	+ 5	+ 10	- 1	- 779	- 1385	Lall.			
7.26	1.1W	5	+ 1	+ 5	- 63	- 790	+ 2	+ 603	+ 5	+ 5	+ 10	- 1	- 785	- 1384	Lall.			
8.19	0.8W	5	- 2	+ 5	- 64	- 785	+ 3	+ 596	+ 3	+ 5	+ 10	- 1	- 780	- 1372	Lall.			
9.19	0.7W	2	- 2	+ 5	- 65	- 779	+ 2	+ 607	+ 12	+ 5	+ 10	- 2	- 775	- 1378	Lall.			
12.22	1.5W	7	- 1	+ 5	- 68	- 789	+ 4	+ 604	+ 3	+ 5	+ 10	- 2	- 786	- 1380	Lall.			
14.23	2.3W	5	- 4	+ 6	- 70	- 779	+ 6	+ 610	+ 4	+ 5	+ 10	- 2	- 775	- 1383	Lall.			
16.19	1.4W	5	0	+ 5	- 72	- 784	+ 2	+ 607	+ 5	+ 5	+ 10	- 2	- 780	- 1383	Lall.			
17.22	2.0W	5	- 2	+ 5	- 73	- 790	+ 4	+ 604	+ 5	+ 1	+ 3	- 2	- 792	- 1395	U			
18.17	1.5W	5	- 1	+ 5	- 74	- 781	+ 4	+ 605	+ 2	+ 1	+ 3	- 2	- 782	- 1385	U			
19.22	2.2W	10	- 4	+ 6	- 75	- 776	+ 3	+ 614	+ 2	+ 1	+ 3	- 2	- 778	- 1389	U			
20.21	2.1W	10	- 3	+ 6	- 77	- 778	+ 3	+ 612	+ 3	+ 1	+ 3	- 2	- 779	- 1389	U			
21.22	2.2W	10	- 4	+ 6	- 78	- 783	+ 3	+ 605	+ 2	+ 1	+ 3	- 2	- 784	- 1387	U			
22.20	1.9W	10	- 3	+ 5	- 79	- 772	+ 3	+ 618	+ 1	+ 1	+ 3	- 2	- 773	- 1389	R			
28.19	2.2W	10	- 4	+ 6	- 86	- 789	+ 4	+ 766	+ 4	+ 1	+ 3	- 2	- 790		U			
29.18	1.9W	5	- 2	+ 5	- 87	- 788	+ 3	+ 677	+ 2	+ 1	+ 3	- 2	- 789		U			
30.19	2.4W	11	- 6	+ 6	- 88	- 783	+ 4	+ 619	+ 4	+ 1	+ 3	- 2	- 784	- 1403	U			
Jul. 1.19	2.4W	10	- 6	+ 6	- 89	- 795	+ 3	+ 598	+ 2	+ 1	+ 3	- 2	- 796	- 1392	U			
												mean	- 782	- 1386				
1955																		
Jan. 12.53	1.2E	45	45	52	10	- 10	+ 3	- 116	- 17	+ 2	+ 1099	+ 2	+ 6	+ 13	- 2	- 13	- 1105	R
Mar. 3.39	0.1E	45	45	52	10	- 7	+ 3	- 58	- 16	+ 1	+ 1102	+ 1	+ 6	+ 13	- 1	- 11	- 1106	R
13.36	1.2E	45	45	52	10	- 8	+ 3	- 50	- 14	+ 2	+ 1100	+ 2	+ 6	+ 13	- 1	- 9	- 1102	R
17.37	0.8E	45	45	52	10	- 7	+ 3	- 47	- 13	+ 2			+ 6	+ 13	- 1	- 8		R
27.33	1.0E	45	45	52	10	- 7	+ 3	- 41	- 16	+ 1	+ 1096	+ 1	+ 6	+ 13	0	- 10	- 1099	R
Apr. 4.31	0.9E	45	45	52	10	- 5	+ 3	- 38	- 15	+ 2	+ 1097	+ 2	+ 6	+ 13	0	- 9	- 1099	R
11.30	0.6E	45	45	52	10	- 6	+ 3	- 37	- 12	+ 3	+ 1098	+ 2	+ 6	+ 13	0	- 6	- 1097	R
13.27	1.2E	45	45	52	10	- 5	+ 3	- 36	- 17	+ 2	+ 1102	+ 2	+ 6	+ 13	0	- 11	- 1106	R
20.36	1.3W	45	45	52	10	- 10	+ 3	- 36	- 14	+ 1	+ 1097	+ 1	+ 6	+ 13	0	- 8	- 1098	R
21.36	1.4W	45	45	52	10	- 10	+ 4	- 36	- 18	+ 1	+ 1099	+ 2	+ 6	+ 13	0	- 12	- 1104	R
25.33	1.1W	45	45	52	10	- 8	+ 3	- 36	- 14	+ 1	+ 1101	+ 1	+ 6	+ 13	0	- 8	- 1102	R
26.34	1.2W	45	45	52	10	- 10	+ 3	- 37	- 16	+ 2	+ 1098	+ 3	+ 6	+ 13	0	- 8	- 1098	U S-Vf
28.32	0.9W	45	45	52	10	- 7	+ 3	- 37	- 15	+ 1	+ 1097	+ 1	+ 7	+ 14	0	- 10	- 1106	U UBV
May 3.31	1.0W	45	45	52	10	- 8	+ 3	- 38	- 10	+ 2	+ 1097	+ 2	+ 2	+ 5	0	- 10	- 1105	U UBV
4.33	1.6W	45	45	52	10	- 11	+ 4	- 39	- 12	+ 2	+ 1099	+ 1	+ 2	+ 5	0	- 13	- 1108	U UBV
9.30	1.2W	45	45	52	10	- 8	+ 3	- 41	- 15	+ 2	+ 1098	+ 2	+ 2	+ 5	0	- 13	- 1105	U UBV
12.28	1.0W	45	45	52	10	- 8	+ 3	- 42	- 15	+ 1	+ 1095	+ 2	+ 2	+ 5	0	- 10	- 1105	U UBV
16.30	1.7W	45	45	52	10	- 12	+ 4	- 44	- 11	+ 2	+ 1098	+ 2	+ 2	+ 5	0	- 8	- 1101	U S-Vf
Jun. 7.22	1.2W	45	45	52	10	- 8	+ 3	- 61	- 14	+ 3	+ 1100	+ 2	+ 7	+ 14	- 1	- 11	- 1103	U S-Vf
16.20	1.3W	45	45	52	10	- 10	+ 3	- 70	- 17	+ 1	+ 1098	+ 2	+ 7	+ 14	- 2	- 13	- 1105	U S-Vf
17.19	1.3W	45	45	52	10	- 10	+ 3	- 71	- 18	+ 1	+ 1099	+ 2	+ 7	+ 14	- 2	- 10	- 1101	U S-Vf
20.22	2.2W	45	45	52	10	- 17	+ 4	- 74	- 15	+ 2	+ 1098	+ 2	+ 7	+ 14	- 2	- 15	- 1107	U S-Vf
21.22	2.2W	45	45	52	10	- 17	+ 4	- 76	- 20	+ 2	+ 1099	+ 2	+ 7	+ 14	- 2	- 11	- 1102	U S-Vf
22.19	1.5W	45	45	52	10	- 11	+ 4	- 77	- 16	+ 2	+ 1098	+ 1	+ 7	+ 14	- 2	- 15	- 1105	U S-Vf
27.19	1.9W	45	45	52	10	- 14	+ 4	- 82	- 20	+ 2	+ 1097	+ 2	+ 7	+ 14	- 2	- 15	- 1105	U S-Vf
Jul. 1.20	2.2W	45	45	52	10	- 18	+ 4	- 87	- 20	+ 2	+ 1097	+ 2	+ 7	+ 14	- 2	- 14	- 1106	U S-Vf
2.20	2.3W	45	45	52	10	- 19	+ 4	- 88	- 19	+ 3	+ 1099	+ 1	+ 7	+ 14	- 2	- 17	- 1110	U S-Vf
6.20	2.6W	45	45	52	10	- 23	+ 4	- 93	- 22	+ 3	+ 1100	+ 1	+ 7	+ 14	- 2	- 16		U S-Vf
7.19	2.4W	45	45	10	- 20	+ 4	- 94	- 21	+ 3			+ 7	+ 14	- 2		- 17		U S-Vf
8.18	2.2W	45	45	10	- 18	+ 4	- 95	- 22	+ 2			+ 7	+ 14	- 2				
												mean	- 11	- 1105				







TABLE X. Magnitude Differences in Blue Color Between Neptune and Comparison Stars (Unit = 0.001) (cont'd)

DATE U.T. 1	Hour Angle 2	Amplifier Gains			Number of measures of planet 6	Correction for dif- ferential extinction 7	Correction for color dependence of ex- tinction 8	Distance correction 9	$\Delta B(N-NA)$ corrected for extinction and distance 10	$\Delta B(NB-NA)$ corrected for extinction 11	Color correction to the BV system 12	Phase Correction 14	$\Delta B(N-NA)$ Final Value 16	$\Delta B(NB-NA)$ Final Value 17	Remarks 18		
		Neptune 3	Star NA 4	Star NB 5													
1960																	
Jul.	1.22	2.2W	53	45	45	6	-12	+13	-76	+1326	+ 2	+ 168	+ 3	-2	-1	-2	+1324 +1155
	12.19	2.1W	53	45	45	7	-12	+13	-89	+1322	+ 1	+ 171	+ 4	-2	-1	-2	+1318 +1148
	16.19	2.4W	53	45	45	6	-14	+14	-94	+1325	+ 1	+ 168	+ 4	-2	-1	-2	+1318 +1154
																mean +1327 +1162	
1961																	
Feb.	10.46	1.6E				9	-6	-1	-100	+ 101	+ 2	-1475	+ 2	0	-2	-2	+ 99 +1572
	15.47	1.0E				8	-4	-1	-91	+ 106	+ 1	-1473	+ 2	0	-2	-2	+ 104 +1575
	22.48	0.6E				6	-3	-1	-83	+ 109	+ 1	-1467	+ 2	0	-2	-2	+ 107 +1572
Mar.	1.46	0.3E				10	-2	-1	-75	+ 105	+ 1	-1473	+ 1	0	-2	-2	+ 103 +1574
	8.46	0.1W				8	-3	-1	-67	+ 107	+ 1	-1473	+ 2	0	-2	-1	+ 106 +1577
	22.41	0.3E				10	-4	-1	-54	+ 104	+ 1	-1471	+ 2	0	-2	-1	+ 103 +1572
	23.45	0.8W				7	-2	-1	-54	+ 103	+ 1	-1471	+ 2	0	-2	-1	+ 102 +1571
	31.44	0.9W				6	-3	-1	-48	+ 106	+ 2	-1476	+ 1	0	-2	-1	+ 105 +1579
Apr.	9.40	0.6W				10	-3	-1	-43	+ 104	+ 1	-1475	+ 1	0	-2	0	+ 104 +1577
May	9.28	0.1W				10	-1	-1	-38	+ 99	+ 1	-1474	+ 1	0	-2	0	+ 99 +1571
	16.31	1.7W				6	-10	-1	-40	+ 100	+ 2	-1476	+ 1	0	-2	0	+ 100 +1574
	17.28	0.5W				10	-4	-1	-40	+ 98	+ 1	-1476	+ 1	0	-2	0	+ 98 +1572
	21.28	0.6W				7	-2	-1	-42	+ 102	+ 1	-1473	+ 1	0	-2	0	+ 102 +1573
Jun.	2.32	2.4W				5	-14	-1	-48	+ 103	+ 4	-1479	+ 1	0	-2	+ 1	+ 100 +1577
	8.23	0.7W				10	-4	-1	-52	+ 100	+ 1	-1476	+ 2	0	-2	+ 1	+ 99 +1572
	10.21	0.5W				10	-4	-1	-54	+ 99	+ 2	-1471	+ 1	0	-2	+ 1	+ 98 +1568
	28.21	1.6W				8	-14	-1	-57	+ 114	+ 3	-1476	+ 2	0	-2	+ 2	+ 108 +1582
																mean + 102 +1574	

U = Unrefrigerated tube, R = Refrigerated tube, Lall. = Lallemand tube, \* = Clouds at end, + = dusty sky,  
f = Filter.

star A. Such mean values, corrected for extinction effects and with distance corrections added, are given in the 10th column of Tables IX and X. They are accompanied by their mean errors computed by taking into consideration the scatter of individual measures and uncertainty of extinction corrections. We can see in Figure 4 that the r.m.s. deviation of the nightly value of the extinction coefficient  $Q_{b1}$  from its mean seasonal value equals approximately 1/5 of this mean seasonal value. Assuming the same relative uncertainty for coefficient  $Q_{b2}$  the mean errors given in the 10th column are computed from the formula

$$\varepsilon = \{ \varepsilon_i^2 + (Q_{b1}\Delta M/5)^2 + [Q_{b2} \bar{M} \Delta(B-V)/5]^2 \}^{1/2} \quad (36)$$

where  $\varepsilon_i$  is the mean error of the mean magnitude difference between the planet and the comparison star A computed from the scatter of individual measures corrected for extinction. This intrinsic mean error  $\varepsilon_i$  is computed either from the sum of squares of the deviations or from the approximate formula [cf Pearson (12)]

$$\varepsilon_i \approx (\text{Max } X - \text{Min } X)/n, \quad (37)$$

where Max X and Min X are the largest and the smallest values of magnitude difference and n is the number of values used. The formula (37) gives for  $3 \leq n \leq 10$  the estimates of mean error of mean values which do not differ systematically from those computed from the sum of squares of the deviations, and are only slightly less accurate. For the observations near the meridian under good atmospheric conditions values of  $\varepsilon_i$  between  $\pm 0^m 0006$  and  $\pm 0^m 0012$  were obtained.

The 11th column of Tables IX and X gives the magnitude difference between the comparison stars B and A, corrected for extinction. The extinction corrections were computed for every measure of star B, using the equations (31) - (35), in which the values for the planet are replaced by corresponding values for comparison star B. The mean errors are computed from equation (36).

The magnitude difference between the planet and comparison star A reduced to the BV photometric system can be obtained by adding the corrective term  $A_s \Delta(B-V)$  to the values in the 10th column of Tables IX and X. The values  $\Delta(B-V)$  found from equation (31) and the mean seasonal values of the transformation coefficient  $A_s$  given in Table II are used for computing the values of this corrective term given in the 12th column of the above mentioned tables (for the years 1953 - 1954 the values of  $A_s$  given in the last column of Table I are used; the observations used for determining these values were not available

to the present author). A similar corrective term for the magnitude difference between the planet and comparison star B is given in the 13th column of Tables IX and X. As can be seen in Table II most of the determinations of the coefficient  $A_s$  were made between January and April. Therefore the assumed values of this coefficient for other months are usually highly uncertain. For estimating errors of corrective terms given in the 12th and 13th columns we should also take into consideration the uncertainty of the gradient color indices,  $(B-V)'$ , of the planets.

The uncertainty of terms  $A_s \Delta(B-V)$  seems to be the principal source of inaccuracy in our observations.

*C. Phase and Oblateness Effects.* The coefficients describing the dependence of brightness of Uranus and Neptune on their phase were computed by Sinton (13), who used the observations made in the present program. These coefficients,  $a$ , and the phase angles,  $i$ , are used for computing the phase corrections,  $\Delta_i = -a_i^2$ , which are given in the 14th column of Tables IX and X. The maximum phase angle is about  $3^\circ$  for Uranus and  $2^\circ$  for Neptune.

The polar axis of Uranus is situated approximately in the plane of the ecliptic and in November 1945 it was directed approximately towards the earth. For the purpose of eliminating the photometric effects of the changing direction of Uranus' polar axis relative to the earth, two assumptions are made. First, we assume that the equatorial plane of Uranus coincides with the mean orbital plane of its four brightest satellites. Second, because of the lack of quantitative information about the brightness distribution on the disk of Uranus, we assume that it is uniform and hence that the total brightness of Uranus is proportional to its projected area, as seen from the earth. The last assumption is, of course, a very rough approximation to the real situation.

On these assumptions Uranus is brightest when its polar axis is directed towards the earth. The corrections are added to all the observed magnitudes of Uranus to give them the values to be expected (on our simplifying assumptions) if the polar axis of Uranus is directed towards the earth. These oblateness corrections are computed from the following formula, resulting from the expression for the projected area and from the definition of stellar magnitude,

$$\Delta_{\text{obl}} = 1.25 \log [1 - (1-b^2) \cos^2(\vartheta_1 - L)], \quad (38)$$

where the square of the ratio of the polar-to-equatorial axis is  $b^2 = 0.833$  according to Lowell (14) and the longitude of the node of the orbits of Uranus' satellites referred to the earth equator is  $\vartheta_1 = 165^\circ 81$  for the epoch 1900 and  $166^\circ 53$  for the epoch 1950 according to Newcomb (15); for the period covered

by the observations here reported we assume  $\theta_1 = 166^{\circ}65$ . By  $L$  is denoted the geocentric longitude of Uranus which can be found from the formula obtained by simple trigonometry,

$$\tan(L - L_{\text{hel}}) = \frac{\sin(L_o - L_{\text{hel}})}{R + \cos(L_o - L_{\text{hel}})}, \quad (39)$$

where  $L_{\text{hel}}$  is the heliocentric longitude of Uranus (given in Nautical Almanac),  $L_o$  is longitude of the sun and  $R$  is the distance of Uranus from the sun, expressed in astronomical units.

The oblateness corrections computed from equation (38) are given in the 15th column of Table IX. The 16th and 17th columns give the finally

adopted magnitude differences between the planet and the comparison stars, computed from the formulas

$$\Delta B(\text{planet-star A}) = \Delta B'(\text{planet-star A}) + A_8[(B-V)' - (B-V)_{\text{star A}}] + \Delta_i + \Delta_{\text{obl}}, \quad (40)$$

$$\Delta B(\text{planet-star B}) = \Delta B'(\text{planet-star A}) - \Delta B'(\text{star B-star A}) + A_8[(B-V)_{\text{star A}} - (B-V)_{\text{star B}}] + \Delta_i + \Delta_{\text{obl}} \quad (41)$$

TABLE XI. Blue Magnitudes of Uranus

Observations From	To	Mean of Final Values	Mean of Final Values	Blue Magnitude of Star A (in System of Ten- Year Standards)	Blue Magnitude of Star B (in System of Ten- Year Standards)	Blue Magnitude of Uranus derived from Star A	Blue Magnitude of Uranus derived from Star B	Correction to be added to B if N directly observed color for Uranus is used	Correction to be added to B if N photometric oblateness of Uranus is twice smaller than geometric
		$\Delta B(U-U_A)$	$\Delta B(U-U_B)$	$B_{U_A}$ m.e.	$B_{U_B}$ m.e.	$B_U$ m.e.	$B_U$ m.e.	$\delta$ color	$\delta$ $\text{obl}$
1953, Jan. 21	1953, Feb. 19	-0.161	-0.272	6.229 $\pm 0.004$	.....	6.068 $\pm 0.004$	.....	-0.013	+0.012
1953, Feb. 23	1953, Apr. 25	-0.159	6.229 $\pm 0.004$	m.	6.070	6.067 $\pm 0.004$	6.067	-0.017	+0.011
1953, Oct. 30	1954, Mar. 5	-1.472	-1.945	7.540 $\pm 0.003$	8.012 $\pm 0.004$	6.068 $\pm 0.004$	6.067 $\pm 0.004$	-0.012	+0.015
1954, Mar. 15	1954, Apr. 26	-0.934	-1.997	7.005 $\pm 0.004$	8.085 $\pm 0.005$	6.071 $\pm 0.005$	6.082 $\pm 0.005$	-0.019	+0.013
1954, Oct. 11	1955, Jan. 12	-3.234	-1.722	9.303 $\pm 0.004$	7.790 $\pm 0.004$	6.069 $\pm 0.004$	6.068 $\pm 0.004$	-0.021	+0.020
1955, Feb. 10	1955, Apr. 26	-3.226	-1.714	9.303 $\pm 0.004$	7.790 $\pm 0.004$	6.077 $\pm 0.004$	6.076 $\pm 0.004$	-0.021	+0.017
1955, Sep. 23	1955, Nov. 8	+0.120	+0.850	5.939 $\pm 0.004$	5.204 $\pm 0.004$	6.059 $\pm 0.004$	6.054 $\pm 0.004$	-0.010	+0.024
1955, Nov. 16	1956, May 17	+0.132	+0.855	5.939 $\pm 0.004$	5.204 $\pm 0.004$	6.071 $\pm 0.004$	6.059 $\pm 0.004$	-0.010	+0.022
1956, Sep. 22	1957, Jan. 18	-1.210	-1.324	7.262 $\pm 0.003$	7.369 $\pm 0.003$	6.052 $\pm 0.003$	6.045 $\pm 0.003$	-0.010	+0.028
1957, Feb. 1	1957, May 2	-1.207	-1.319	7.262 $\pm 0.003$	7.369 $\pm 0.003$	6.055 $\pm 0.003$	6.050 $\pm 0.003$	-0.010	+0.025
1957, Sep. 21	1957, Dec. 12	-1.533	-1.334	7.583 $\pm 0.003$	7.383 $\pm 0.003$	6.050 $\pm 0.003$	6.049 $\pm 0.003$	-0.010	+0.033
1957, Dec. 26	1958, Jun. 1	-1.531	-1.332	7.583 $\pm 0.003$	7.383 $\pm 0.003$	6.052 $\pm 0.003$	6.051 $\pm 0.003$	-0.010	+0.030
1958, Oct. 3	1959, Apr. 10	-1.190	-2.707	7.237 $\pm 0.003$	8.761 $\pm 0.004$	6.047 $\pm 0.004$	6.054 $\pm 0.004$	-0.008	+0.034
1960, Jan. 28	1960, May 28	-2.717	-2.968	8.758 $\pm 0.004$	9.012 $\pm 0.003$	6.041 $\pm 0.003$	6.044 $\pm 0.003$	-0.006	+0.038
1960, Oct. 20	1961, May 21	-1.407	-1.475	7.441 $\pm 0.003$	7.508 $\pm 0.003$	6.034 $\pm 0.003$	6.033 $\pm 0.003$	-0.004	+0.042

TABLE XII. Blue Magnitudes of Neptune

Observations From	To	Mean of Final Values	Mean of Final Values	Blue Magnitude of Star A (in System of Ten- Year Standards)	Blue Magnitude of Star B (in System of Ten- Year Standards)	Blue Magnitude of Neptune derived from Star A	Blue Magnitude of Neptune derived from Star B	Correction to be added to B if N directly observed color for Neptune is used
		$\Delta B(N-N_A)$	$\Delta B(N-N_B)$	$B_{N_A}$ m.e.	$B_{N_B}$ m.e.	$B_N$ m.e.	$B_N$ m.e.	$\delta$ color
1953, Mar. 6	1953, Jul. 3	-0.516	(8.784)*	.....	.....	(8.268)*	.....	-0.017
1954, Feb. 4	1954, Mar. 6	-0.788	-1.383	9.035 $\pm 0.004$	9.632 $\pm 0.005$	8.247 $\pm 0.005$	8.249 $\pm 0.005$	-0.015
1954, Mar. 28	1954, Jul. 1	-0.782	-1.386	9.035 $\pm 0.004$	9.632 $\pm 0.005$	8.253 $\pm 0.005$	8.246 $\pm 0.005$	-0.022
1955, Jan. 12	1955, Jul. 8	-0.011	-1.105	8.260 $\pm 0.004$	9.356 $\pm 0.008$	8.249 $\pm 0.008$	8.251 $\pm 0.008$	-0.026
1956, Jan. 17	1956, Jun. 27	+0.790	-0.001	7.455 $\pm 0.004$	8.240 $\pm 0.004$	8.245 $\pm 0.004$	8.239 $\pm 0.004$	-0.013
1957, Jan. 17	1957, Jun. 24	+0.017	+0.920	8.225 $\pm 0.005$	7.316 $\pm 0.005$	8.242 $\pm 0.005$	8.236 $\pm 0.005$	-0.017
1958, Jan. 8	1958, Jul. 5	+0.020	+2.427	8.208 $\pm 0.004$	5.807 $\pm 0.003$	8.228 $\pm 0.003$	8.234 $\pm 0.003$	-0.013
1959, Jan. 11	1959, Jun. 10	+0.234	+3.580	8.001 $\pm 0.004$	4.646 $\pm 0.004$	8.235 $\pm 0.004$	8.226 $\pm 0.004$	-0.011
1960, Jan. 29	1960, Jul. 16	+1.327	+1.162	6.908 $\pm 0.003$	7.075 $\pm 0.003$	8.235 $\pm 0.003$	8.237 $\pm 0.003$	-0.009
1961, Feb. 10	1961, Jun. 28	+0.102	+1.574	8.135 $\pm 0.004$	6.662 $\pm 0.005$	8.237 $\pm 0.005$	8.236 $\pm 0.005$	-0.007

\* Not to be reduced to the system of Ten-Year Standards

Since the errors of the coefficients  $A_8$ , of the planetary gradient color indices  $(B-V)'$  and of the oblateness corrections are not known, the mean errors of the final magnitude differences, computed from equations (40) and (41), can not be estimated. Their weighted mean values are computed for every interval of several months during which  $A_8$  apparently did not change rapidly. These mean values are given in the corresponding columns of Tables IX and X and in the 2nd and 3rd columns of Tables XI and XII. The weights used for computing the mean values of  $\Delta B(\text{planet} - \text{star A})$  are taken as inverse squares of mean errors given in the 10th column of Tables IX and X. The mean errors given in the 10th and 11th columns of these tables are used for computing the weights of  $\Delta B(\text{planet} - \text{star B})$ .

*D. Variations in Brightness of Uranus and Neptune.* The mean magnitude differences between the planets and their comparison stars, the blue magnitudes of comparison stars and the blue magnitudes of the planets obtained by comparison with star A and, separately, by comparison with star B, are given in the 3rd to 8th columns of Tables XI and XII. The blue magnitudes of the comparison stars, given in these tables, are reduced to the system of Ten-Year Standards and obtained by adding the yellow magnitudes and blue-yellow colors of the same stars given in Table VI. The blue magnitudes of the planets given in the 7th and 8th columns of Tables XI and XII are obtained by adding the 3rd and 5th or 4th and 6th columns of these tables, respectively. The values for the year 1954 marked by colons are obtained with Lallemand's multiplier tube and may not be directly comparable with other values which were all obtained with the 1P21 multiplier tubes. There is no obvious evidence of variability of any of the comparison stars except for UA56 =  $\mu$ Cnc. The values obtained by comparison with this star are marked by colons in Table XI. The blue magnitudes reduced to the system of Ten-Year Standards are not available for the comparison stars UB 53 = HD 50692 and NA 53 = HD 116681. The value for this last star obtained by comparisons with primary standard stars during the year 1953 is given in Table XII in parentheses.

For every opposition the average of blue magnitudes given in the 7th and 8th columns of Tables XI and XII is computed (omitting the values with colons and in parentheses) and given in Table XIII. The blue magnitudes of the planets given in this table are plotted in Figure 6. This figure shows an increase in brightness of Uranus by about 0<sup>m</sup>035 and an increase in brightness of Neptune by about 0<sup>m</sup>014.

Examining the mean values of the transformation coefficient  $A_8$  given in Tables I and II we no-

TABLE XIII.

The mean blue magnitudes, B, for the two planets, obtained assuming H. L. Johnson's gradient color-indices of planets and photometric oblateness of Uranus equal to geometric.

Opposition	Uranus	Neptune
1953	6.069	
1954	6.068	8.248
1955	6.072	8.250
1956	6.056	8.242
1957	6.050	8.239
1958	6.050	8.231
1959	6.050	8.230
1960	6.042	8.236
1961	6.034	8.236

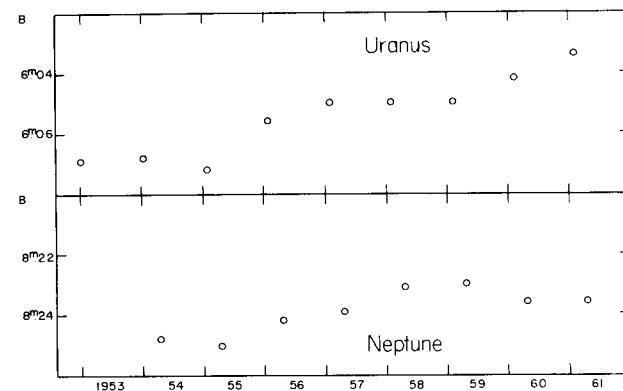


Figure 6. The mean blue magnitudes for the two planets obtained using H.L. Johnson's values of the gradient of the energy distribution in the spectra of these planets within the blue filter spectral region. For Uranus it is assumed that the distribution of brightness over the apparent disc of the planet is uniform.

tice that the values of this coefficient are decreasing during the period covered by the present observations. Therefore, if we assume another value for the gradient color-index of the planets we would obtain another rate of increase in their brightness. In particular, assuming for the reduction to the BV system and for correction for the color-dependence of extinction the directly observed color-indices of the planets (as given in Table VIII), instead of gradient color-indices  $(B-V)'$  given by the equations (28) and (29), we must add to the blue magnitudes of

the planets the corrections given in the 9th column of Tables XI and XII. They are computed from the formula

$$\delta_{\text{color}} = (A_s - Q_{b_2} \bar{M}) [(B-V)' - (B-V)], \quad (42)$$

where  $B-V$  is the directly observed color index of the planet. The mean blue magnitudes of Neptune corrected in this way are given in Table XIV and in

TABLE XIV.

The mean blue magnitudes,  $B$ , for the two planets, obtained assuming the directly observed color-indices of planets and photometric oblateness of Uranus half as great as geometric.

Opposition	Uranus	Neptune
1953	6.066	—
1954	6.071	8.233
1955	6.070	8.224
1956	6.069	8.229
1957	6.067	8.222
1958	6.072	8.218
1959	6.076	8.219
1960	6.074	8.227
1961	6.072	8.229

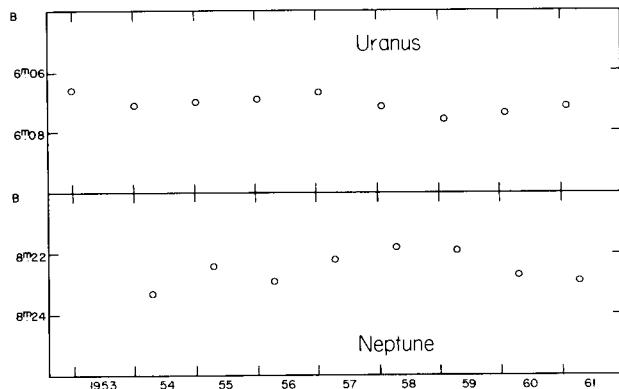


Figure 7. The mean blue magnitudes for the two planets obtained assuming that the gradient of the energy distribution in the spectrum of each of these planets within the blue filter spectral region is the same as for a star having the same  $B-V$  color-index as the planet. For Uranus it is assumed that the photometric effects due to its oblateness are half as great as those for the uniform distribution of brightness over the apparent disc of the planet.

Figure 7. No significant change in brightness of Neptune can be seen in these data.

The brightness of Uranus also will not show the significant change in the years 1953 - 1961 if, in addition to assuming the directly observed color-index, we shall assume that the photometric effects of oblateness of Uranus are half those computed on the assumption of uniform brightness distribution on the disc of the planet. Taking into consideration the presence of equatorial belts and limb darkening on Uranus this last assumption seems reasonable.

The corrections which should be added to the blue magnitudes of Uranus if "photometric" oblateness is half the geometric, are given in the last column of Table XI. The final blue magnitudes of Uranus with the corrections from the last two columns of Table XI are given in Table XIV and in Figure 7.

Our present knowledge of the distribution of energy in the spectra of Uranus and Neptune and of the distribution of brightness on the disc of Uranus is not satisfactory for deciding whether there are any changes in brightness of Uranus and Neptune caused by variability of solar energy output. As soon as these can be determined by actual observations the degree of solar variability can be more accurately found from existing data.

#### V. Suggestions for Future Observers in this Program

Special observations should be undertaken for redetermining the gradient colors  $(B-V)'$  of the planets which are used for transforming the observations into the BV system and for computing the corrections for the color-dependence of extinction. It should be taken into account that these gradient colors depend critically on the width of the band-pass for the combination of the filter, multiplier tube and the telescope.

The limb darkening of Uranus and, hence, the photometric effects of its oblateness should, if possible, be determined from photoelectric scans across the disc of the planet.

Every series of measurements of the planet in blue color should be accompanied by observations of at least 4 standard stars, preferably in 2 colors. Two low altitude ( $22^\circ - 28^\circ$ ) standard stars and two high altitude stars should be observed. One of the low altitude stars should be situated in the East and the other in the West. One of the standard stars observed at high altitude should be blue, the other red. These observations will be used for determining the coefficients  $Q_{b_1}$  and  $A_s$ . Planetary observations at hour angles larger than  $1\frac{1}{2}$  hours for negative declinations and larger than 3 hours for positive declinations should not be made.

The comparison stars should be chosen so that their colors do not differ much from the gradient

color-index (B-V)' of the planet; one of them should be redder than the planet, another bluer. They should have very nearly the same declination as the planet.

The telescope, filters, multiplier tube or amplifier should not be changed without serious reason. If the multiplier tube must be changed, the new tube should be selected so that the coefficients  $A_6$  and  $A_8$  are very near to zero. The image of the telescope's mirror on the cathode should be carefully centered for maximum output.

The calibration of the amplifier's gain and tests for linearity of the amplifier and recorder should be made often, preferably every 2 or 3 months. These tests should be done at widely differing temperatures for checking any temperature effect. The battery used as a power supply for the multiplier tube and that used in the recorder should often be checked.

Every observation should be made at such gain of the amplifier that the deflection of the recorder is not smaller than 0.55 and not larger than 0.95 of full scale. This means that in most cases the observations with the yellow filter should be made at an amplifier's gain different from that used for the observations of the same star with blue filter.

To avoid fatigue of the multiplier tube no standard stars should be used for which  $B < 3^m 7$ . In particular,  $\eta$  Boo,  $\beta$  Lib and  $\alpha$  Ari should not be used as standards. The list of stars used for determining the extinction and transformation coefficients should contain not only primary standards of the UBV system but at least 8 other well observed stars selected from among the secondary standards. The stars  $\kappa$  Gem, HD 73665,  $\vartheta$  Hya,  $\rho$  Leo, 109 Vir,  $\iota$  Her,  $\beta$  Oph and  $\mu$  Her A may be suitable for that purpose.

The accuracy of two-color observations can be highly increased if comparison stars and standard stars are all measured at the same fixed altitude. Only a few (4 to 6) standard stars should be observed at substantially different altitudes to make possible determining the extinction coefficients. On those nights when Neptune comparison stars are observed, the fixed altitude must be low because of the low declination of Neptune and extinction stars can be observed near the zenith. On these nights the Uranus comparison stars should not be observed; they should be observed on other nights when a higher value of fixed altitude is chosen. Every year each of the comparison stars used since 1953 and each of the comparison stars chosen for the next year should be observed on at least 6 nights. Eight stars or more should be included in the least-squares solution for simultaneous determination of extinction and transformation coefficients for every night.

The radioactive standard source should be observed through color filters. Usually the two-color

observations of a star can be accompanied by the measurement of the standard source with only one (e.g. blue) filter. However, before and after measuring the bright star (for which, say,  $B < 4^m 0$ ) the standard source should be measured with both yellow and blue filters for checking if the fatigue of the multiplier tube is the same in both colors.

The observations should be made only on photometric nights with no trace of clouds. It is much better to have fewer but superior observing nights. The observations made on poor nights may destroy the accuracy of the whole program. When the observations are made in the presence of the Moon, the sky background should be measured alternately north and south from the star (or planet) to eliminate the influence of moonlight reflected within the telescope tube.

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