

INSPACE

MADE

Optimast-SCI Science: New Discovery Frontiers with Sensitive, **Millarcsecond Resolution**

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Optical interferometry from a spacebased platform is freed from the limitations of the Earth's atmosphere.

One fundamental benefit is sensitivity.



Gains of 100-1000× or more are possible, compared to ground-based optical interferometry.

Imaging the Terrestrial Planet Forming Regions of Nearby YSOs

• ALMA observations are mapping out regions outside the water snow line; e.g. Isella + 2018, HD163296 above, with 40mas resolution (inset on lower right is our own solar system) Optimast will have > 5× greater resolution • Spanning 0.4-1.0µm, Optimast will image both scattered and thermal light

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Fundamental Parameters of Low Mass Stellar Binaries

Follows ground-based work of bright stars.



Single Object Targets Resolvable by Optimast (N>1,000)

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Detection

O AO

LC

100

O Rad

1000

•The orbit of 12 Boo (Boden et al. 2005) measured by the Palomar Testbed Interferometer, which determined component masses to 0.3%. •Despite having a shorter base-line, with a shorter operational wavelength, Optimast will have similar angular resolution but for far fainter targets (e.g. low-mass stars).

Solar Panel



Extreme optical resolution (~2mas) from a twoelement optical space interferometer: • Shapes, sizes of $\sim 10^3$ asteroids • Orbit and mass determination for all known V<16 binaries

• Spatially resolved spectra-photometry in >10 wavelength bands

• Space-based operation enables high sensitivity for reflected light objects

• Application of ground-based interferometry techniques (well developed but limited to emitted light objects, i.e. stars)

Main Belt Asteroids

• Sizes, shapes for any object > 10km (H<12.3) • Resolved surface mapping for > 30 km Rotation > 6 hours (<5° 'smear' in 300sec) • Detection of binaries, Keplerian solutions for binary orbits • Hundreds of possible targets

Near-Earth Asteroids

• Direct size determination for >10m objects (H<26) • Mapping of binary orbits

Jupiter Trojans • H<9.2 (~36 known targets)

Narrow emission line region dusty outflow UV & X-ray dusty torus snow starburst …… 会会会会 SMBH cold gas/dust disk

Probing the Inner Core Architecture of Active Galactic Nuclei

• At 20 Mpc, 1mas resolution probes the inner edge of the AGN disk • The feedback process between the dusty wind and host



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ONLINE RESOURCES

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