

LightBeam: Milliarcsecond Imaging in the NUV and Visible Leveraging Optical Interferometry and **In-Space Robotic Manufacturing and Assembly**

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has been supported by the NASA SBIR program, the Lowell Observatory, and Redwire / Made In Space.

Exploring the inner pixel of ALMA YSOs

Narrow Map the cores of règion UV & X-ray SMBH

__//-1 AU 0.1 pc 1 pc At 20Mpc 1.0 mas 10 mas 100 mas

Young Stellar Objects

 LightBeam will be able to explore the inner pixel of ALMA targets · Morphology of disk structures in the terrestrial planet forming region will complement ice giant maps Active Galactic Nuclei . LightBeam can probe the inner 0.1-1.0 parcsec of AGN out to 20 Mpc

· Unique mapping of the inner edge of the dusty torus, exploring core binarity

Low-Mass Binaries Direct orbit determinations for the lowest mass stars will provide mass measures

Main Belt Asteroids & Jupiter Trojans Sizes, shapes for any main belt object > 10km (H<12.3) Resolved surface mapping for > 30 km

 ~36 known Jupiter Trojans Detection of binaries, Keplerian solutions for

binary orbits · Additional targets: gas giant moons, ice dwarfs

Beam combiner and printed booms being tested together with target tracking tests

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rotation), and (c) its pole-to-equator intensity profile (due to gravity darkening).

LightBeam will use these techniques to map disk structures across the universe.

Orbit of 12 Boo from Boden et al. 2005.

Using the Palomar Testbed Interferometer, Boden et al. were able to map the orbit of 12 Boo and fit a Keplerian orbit, resulting in ~0.33% mass

LightBeam will have the sensitivity, resolution and contrast to map the orbits of



avionics bay, under the feedstock container on top. Below, a view of LightBeam with partially deployed

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Examples Results of Size/Shape, and a Binary Orbit from Stellar Astronomy: LightBeam will apply these demonstrated techniques to far fainter targets